



The Nickel Isotopes – Status of Theory -  
Then and Now

Alex Brown – Michigan State University

# KARLSRUHER NUKLIDKARTE 7. Auflage 2006, revised printing 2007

CHART OF THE NUCLIDES, 7<sup>th</sup> Edition 2006 / CARTE DES NUCLEIDES, 7<sup>ème</sup> Edition 2006

CARTA DE NUCLEIDOS, 7<sup>a</sup> Edición 2006 / Таблица радионуклидов, 7-е издание

核素图, 第7版

J. Magill<sup>1</sup>, G. Pfennig<sup>2</sup>, J. Galy<sup>1</sup>

<sup>1</sup> European Commission – Joint Research Centre – Institute for Transuranium Elements  
P.O. Box 2340, 76125 Karlsruhe, Germany

<sup>2</sup> formerly Forschungszentrum Karlsruhe in der Helmholtz-Gemeinschaft  
P.O. Box 3640, 76021 Karlsruhe, Germany

© European Communities 2006–2007

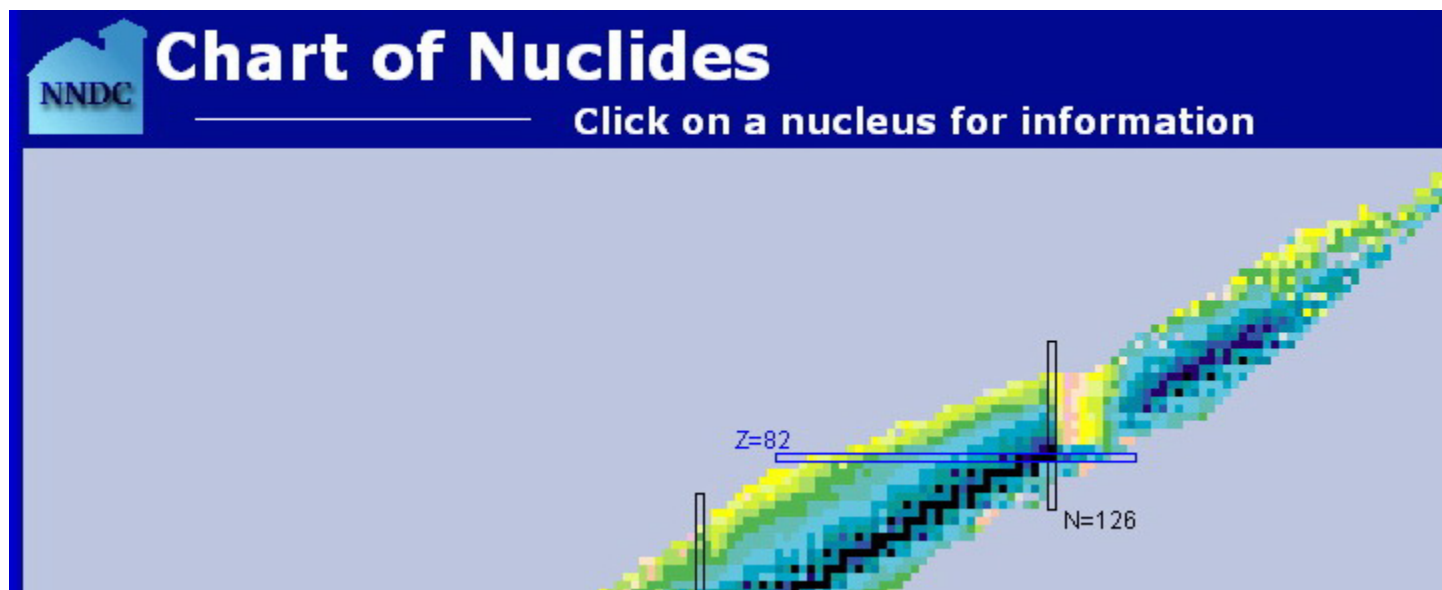
### Legal Notice

Neither the European Commission nor any person acting on behalf of the Commission is responsible for the use which might be made of the following information.

ISBN 92-79-02175-3 © European Communities, 2006 / Catalogue number: LC-NA-22276-EN-C

Li 6.941 <small><math>\sigma_{\text{abs}} 71</math></small>		Li 4 5.0 MeV $91 \cdot 10^{-24}$ s <small>p</small>	Li 5 1.23 MeV $370 \cdot 10^{-24}$ s <small>p</small>	Li 6 7.59 <small><math>\sigma 0.039</math> <math>\sigma_{\text{th}, \alpha} 940</math></small>	Li 7 92.41 <small><math>\sigma 0.045</math></small>	Li 8 840.3 ms <small><math>\beta^- 12.5</math> <math>\beta 2\alpha -1.6</math></small>	Li 9 178.3 ms <small><math>\beta^- 13.6...</math> <math>\beta n 0.7...</math> <math>\beta \alpha</math></small>	Li 10 230 keV $2.0 \cdot 10^{-21}$ s <small>n</small>	Li 11 8.5 ms <small><math>\beta^- 18.5; 20.4</math> <math>\beta 3369; 320...</math> <math>\beta n; \beta 2n; \beta 3n;</math> <math>\beta \alpha; \beta \beta; \beta \delta</math></small>
He 4.002602 <small><math>\sigma_{\text{abs}} &lt;0.05</math></small>		He 3 0.000134 <small><math>\sigma 0.00005</math> <math>\sigma_{\text{th}, p} 5330</math></small>	He 4 99.999866	He 5 648 keV $700 \cdot 10^{-24}$ s <small>n</small>	He 6 806.7 ms <small><math>\beta^- 3.5</math> <math>\beta \delta</math></small>	He 7 159 keV $2.9 \cdot 10^{-21}$ s <small>n</small>	He 8 119 ms <small><math>\beta^- 9.7...</math> <math>\gamma 981; 478^*</math> <math>\beta n; \beta \beta</math></small>	He 9 65 keV $7 \cdot 10^{-21}$ s <small>n</small>	He 10 0.17 MeV $2.7 \cdot 10^{-21}$ s <small>2n</small>

0.75E-4



Ni 48 ~2 ms ?	Ni 49 13 ms	Ni 50 12 ms	Ni 51 >200 ns	Ni 52 38 ms	Ni 53 45 ms	Ni 54 104 ms	Ni 55 209 ms	Ni 56 6.075 d
2p 1.35 ?	$\beta^+$ $\beta_p$ 3.7	$\beta^+$ $\beta_p$	$\beta^+$ ?	$\beta^+$ $\beta_p$ 1.34; 1.06	$\beta^+$ $\beta_p$ 1.90	$\beta^+$ $\gamma$ 937 g	$\beta^+$ 7.7... $\gamma$ (2919; 2976; 3303)	$\epsilon$ ; no $\beta^+$ $\gamma$ 158; 812; 750; 480; 270...

Ni 56 6.075 d	Ni 57 36.0 h	Ni 58 68.0769	Ni 59 $7.5 \cdot 10^4$ a	Ni 60 26.2231	Ni 61 1.1399	Ni 62 3.6345	Ni 63 100 a	Ni 64 0.9256	Ni 65 2.52 h	Ni 66 54.6 h	Ni 67 21 s	Ni 68 29 s
$\epsilon$ ; no $\beta^+$ $\gamma$ 158; 812; 750; 480; 270...	$\epsilon$ $\beta^+$ 0.8... $\gamma$ 1378; 1920; 127...	$\sigma$ 4.6 $\sigma_{n,\alpha} < 0.00003$	$\epsilon$ ; $\beta^+$ ... no $\gamma$ ; $\sigma$ 77.7 $\sigma_{n,\alpha}$ 14; $\sigma_{n,p}$ 2 $\sigma_{abs}$ 92	$\sigma$ 2.9	$\sigma$ 2.5 $\sigma_{n,\alpha}$ 0.00003	$\sigma$ 15	$\beta^-$ 0.07 no $\gamma$ $\sigma$ 20	$\sigma$ 1.6	$\beta^-$ 2.1... $\gamma$ 1482; 1115; 366... $\sigma$ 22	$\beta^-$ 0.2 no $\gamma$	$\beta^-$ 3.8... $\gamma$ (1937; 1115; 822...)	$\beta^-$ $\gamma$ 758; 84 g

Ni 68 29 s	Ni 69 11.4 s	Ni 70 6.0 s	Ni 71 2.56 s	Ni 72 1.57 s	Ni 73 0.84 s	Ni 74 0.9 s	Ni 75 344 ms	Ni 76 238 ms	Ni 77 128 ms	Ni 78 110 ms
$\beta^-$ $\gamma$ 758; 84 g	$\beta^-$ $\gamma$ 1871; 680; 1213; 1483...	$\beta^-$ 3.3... $\gamma$ 1036; 78... $m_2$	$\beta^-$ $\gamma$ 534; 2016	$\beta^-$ $\gamma$ 376; 94	$\beta^-$ $\gamma$ 166; 1010	$\beta^-$ $\gamma$ 166*; 694 $\beta_n$	$\beta^-$	$\beta^-$	$\beta^-$	$\beta^-$



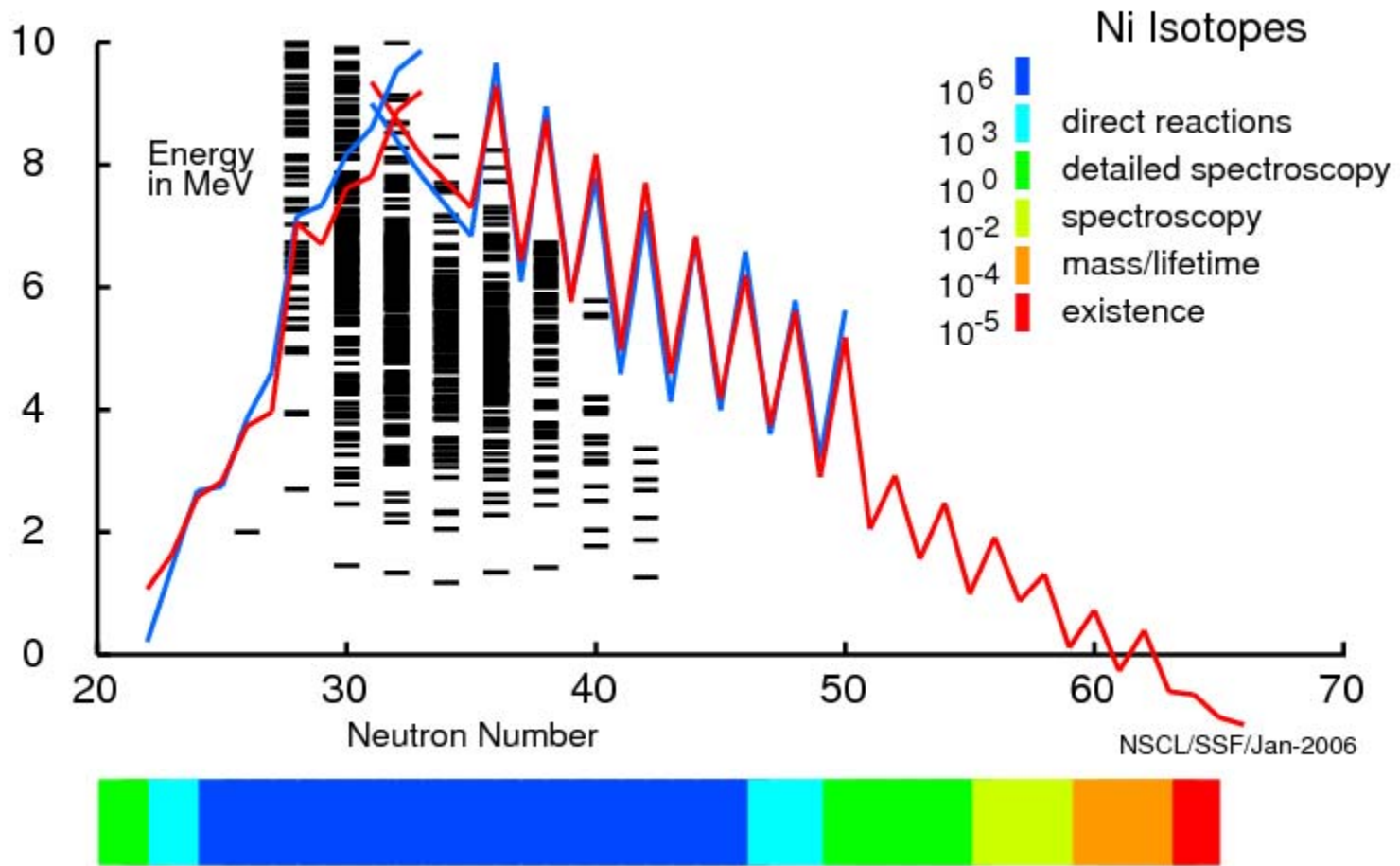
# Isotope Science Facility at Michigan State University

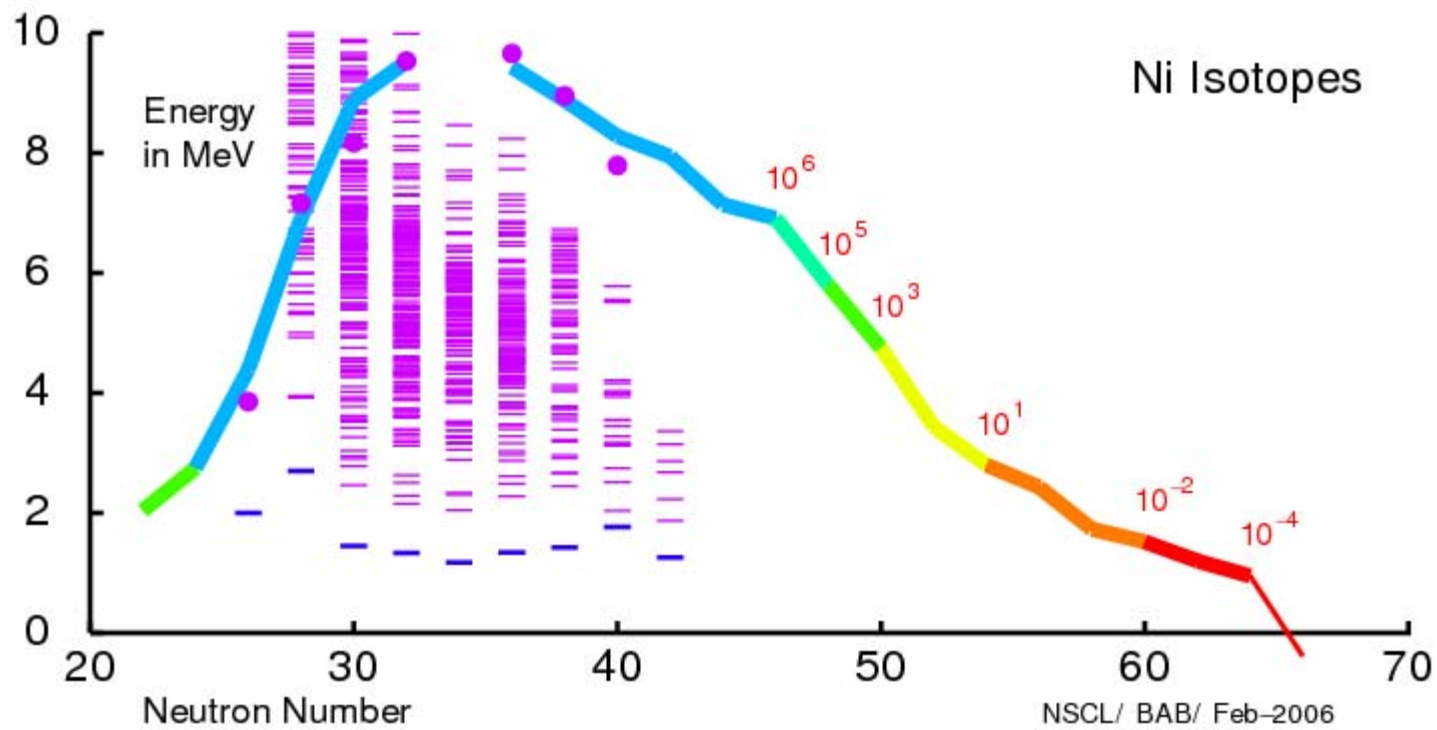
Upgrade of the NSCL rare isotope  
research capabilities

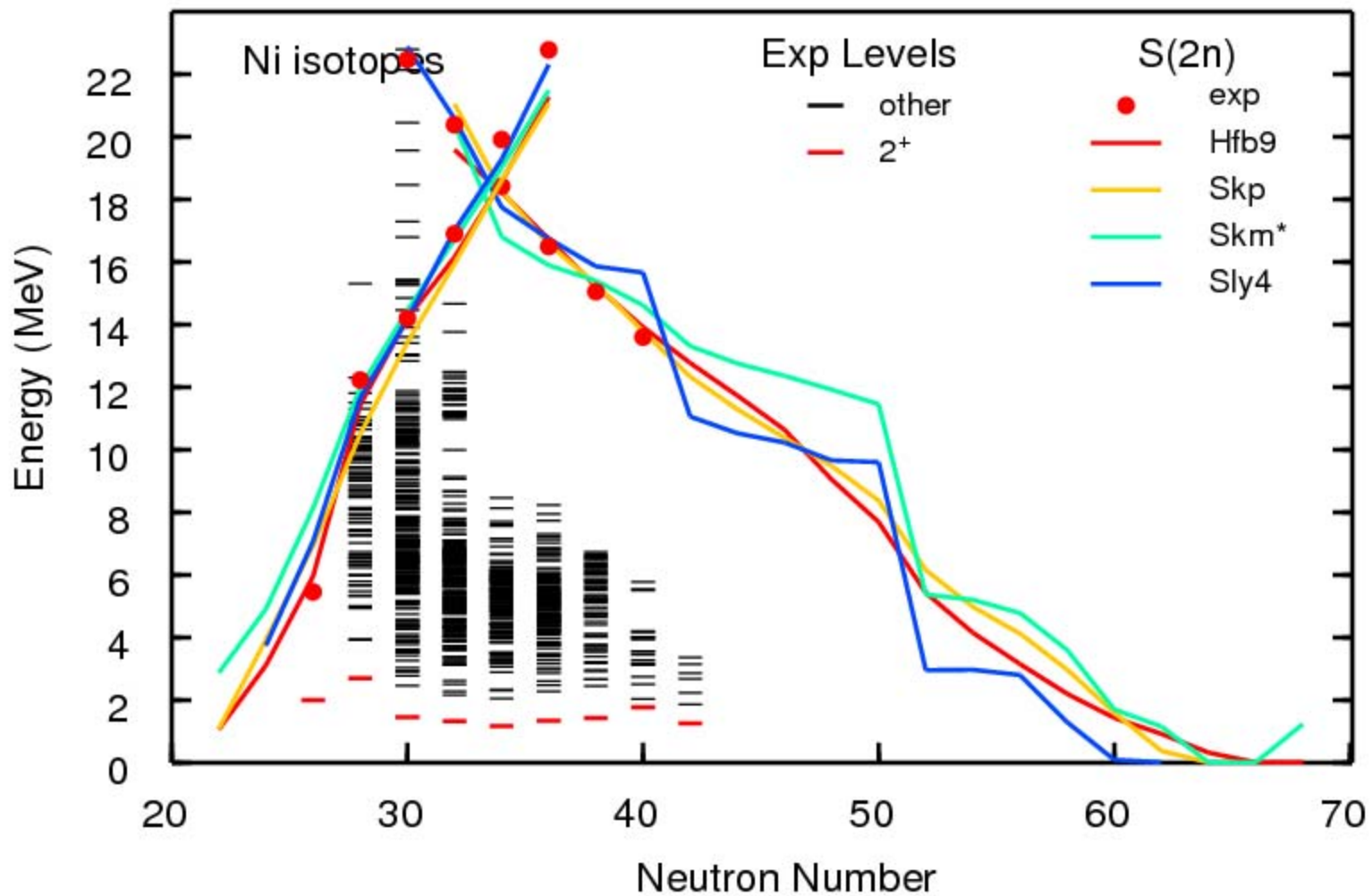


BNL, Nov 6, 2008

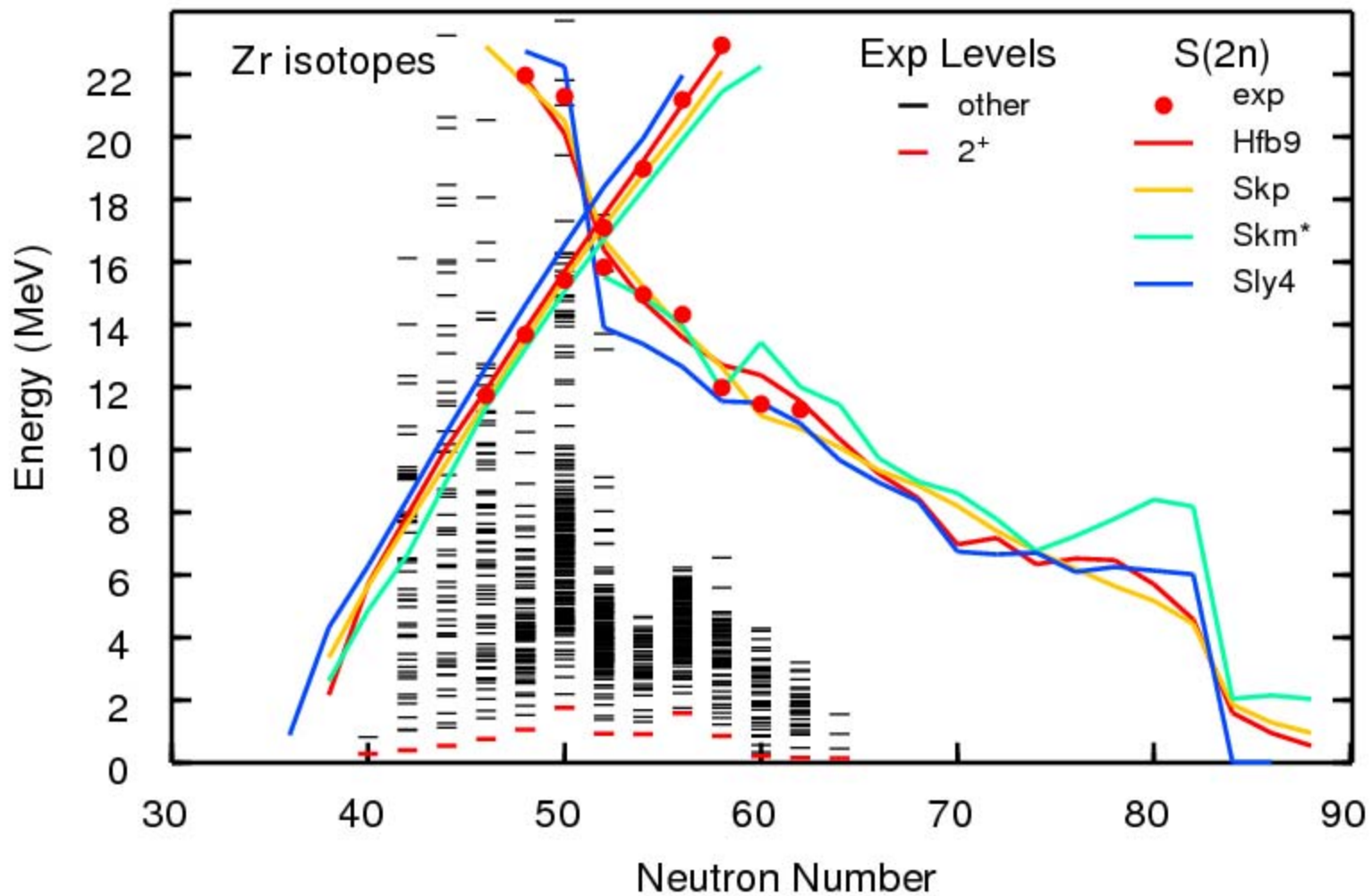




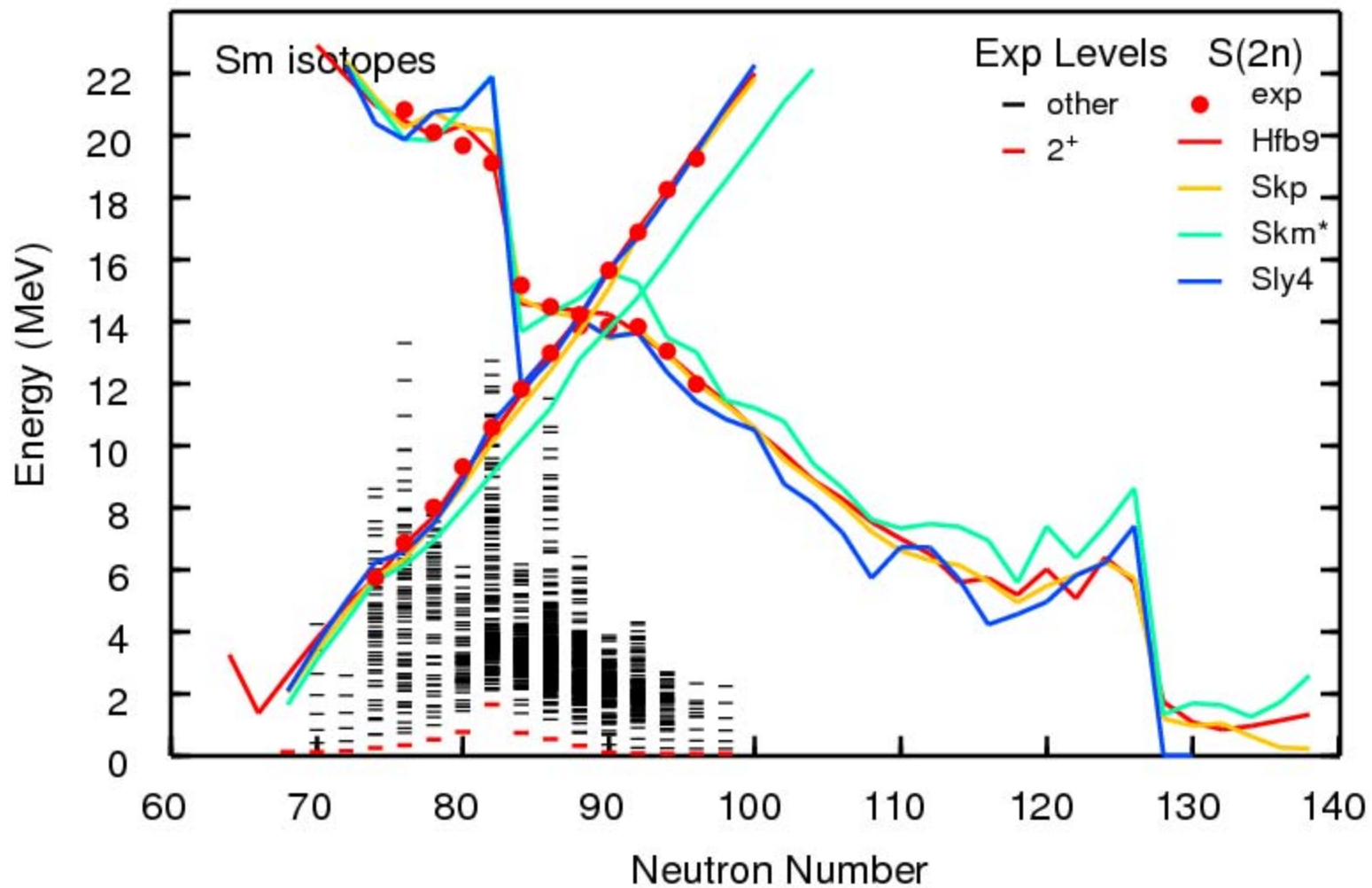


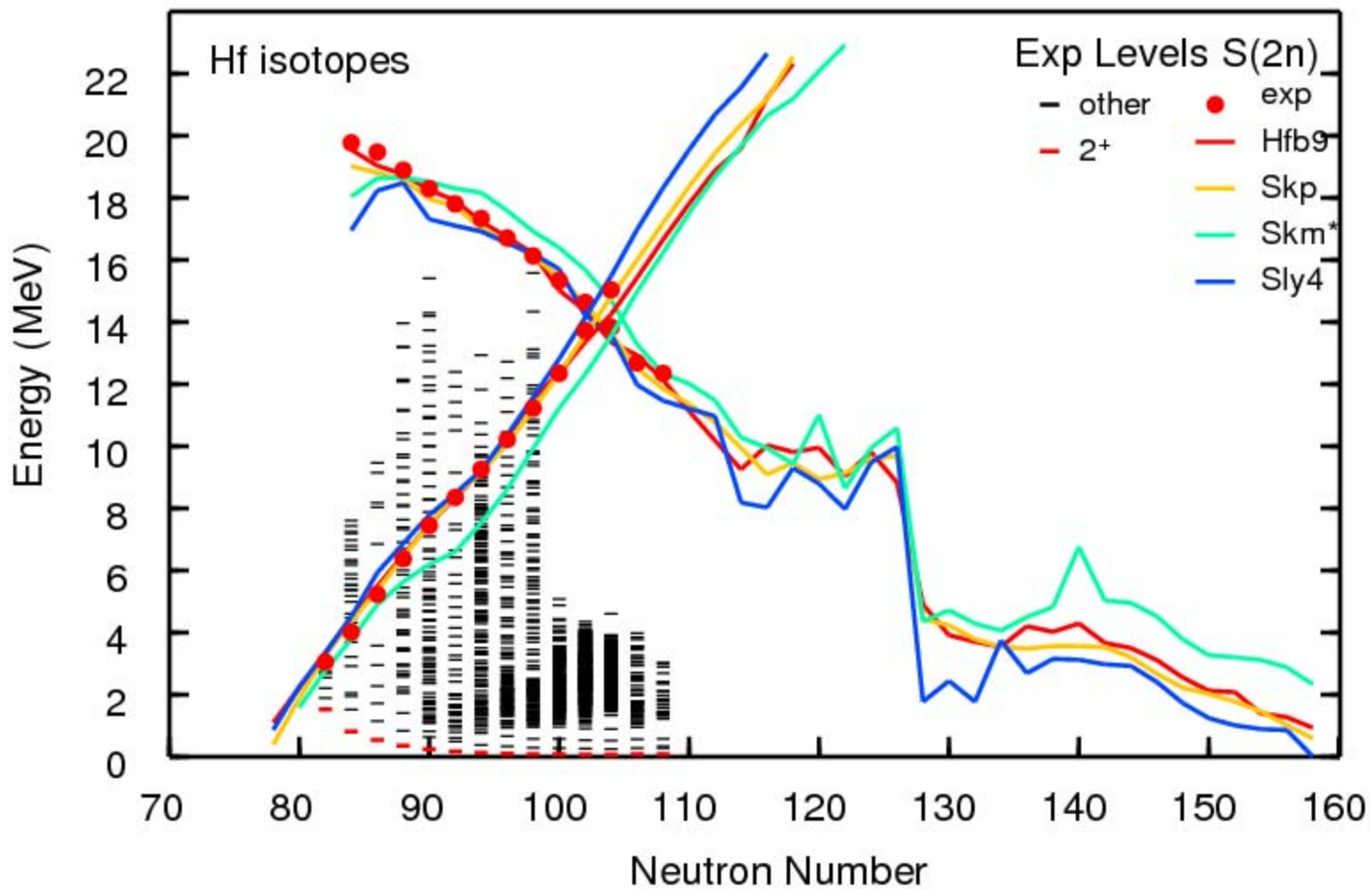


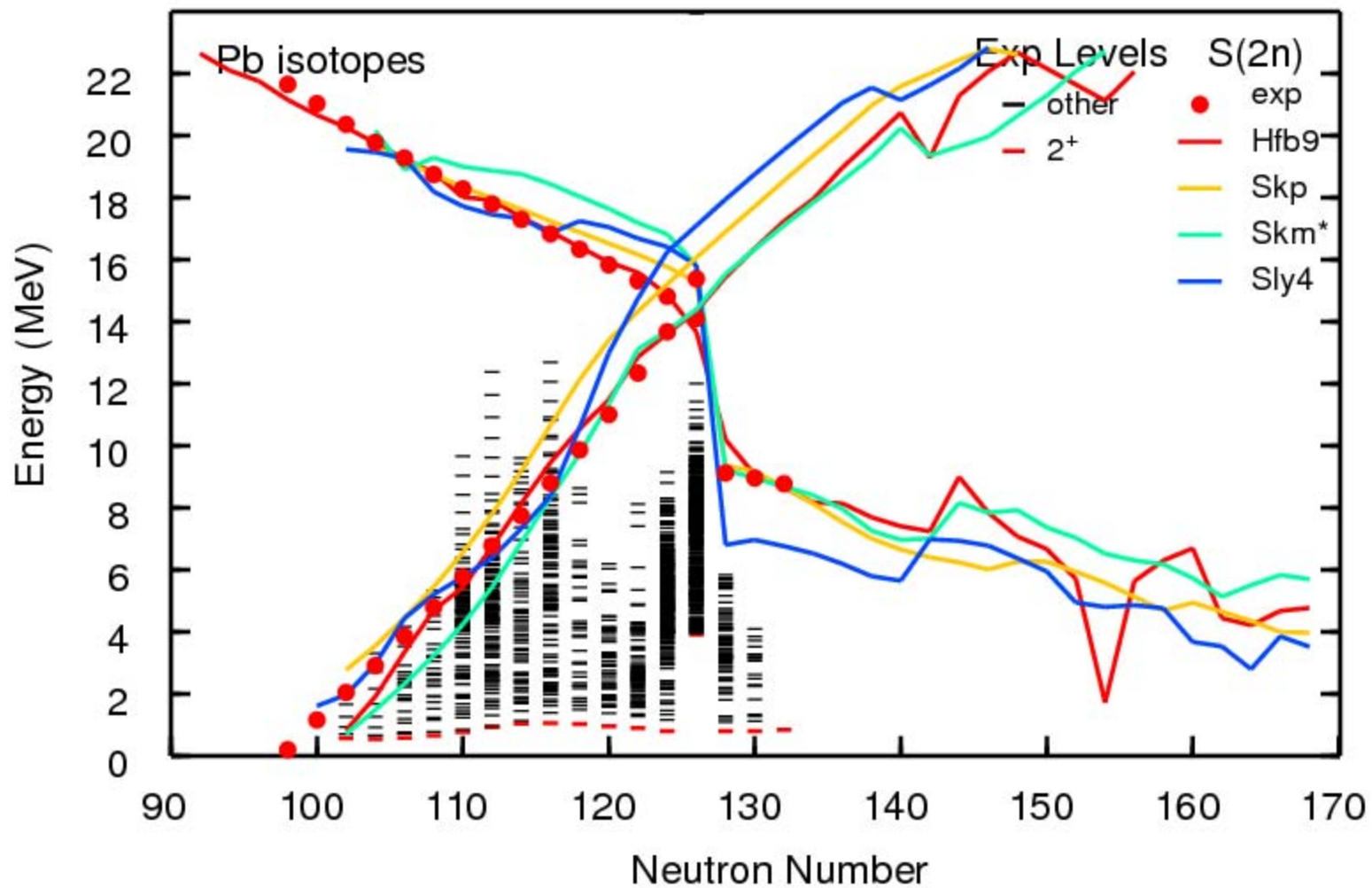






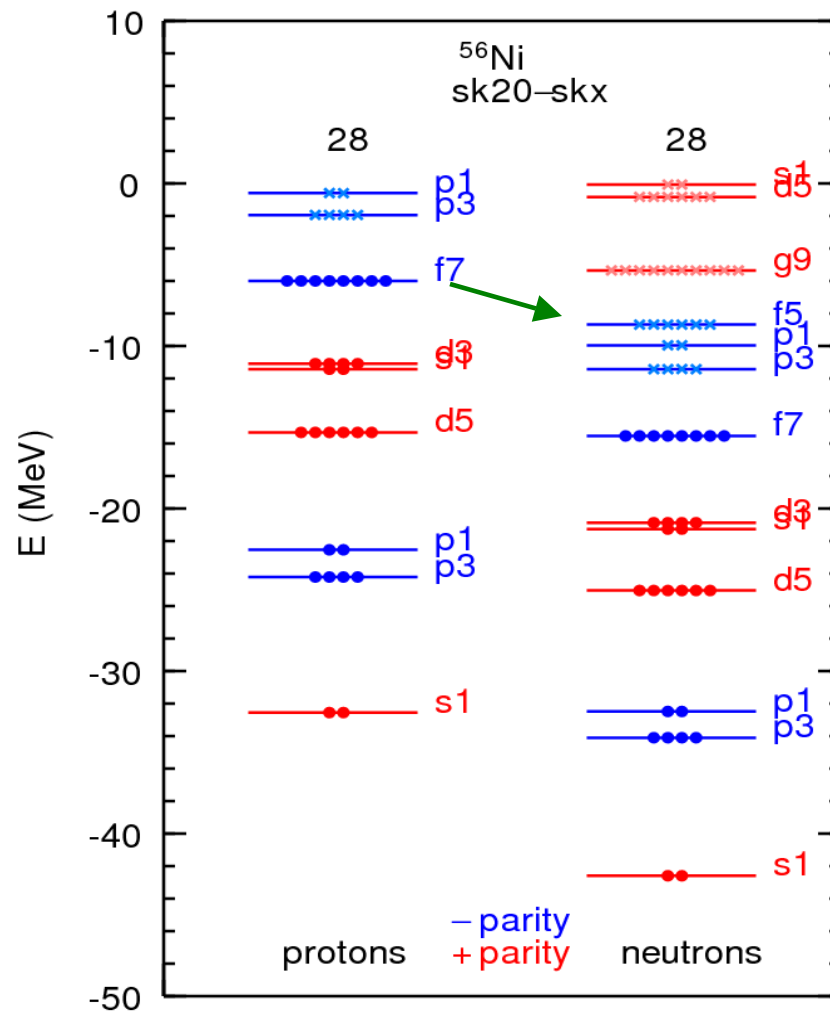


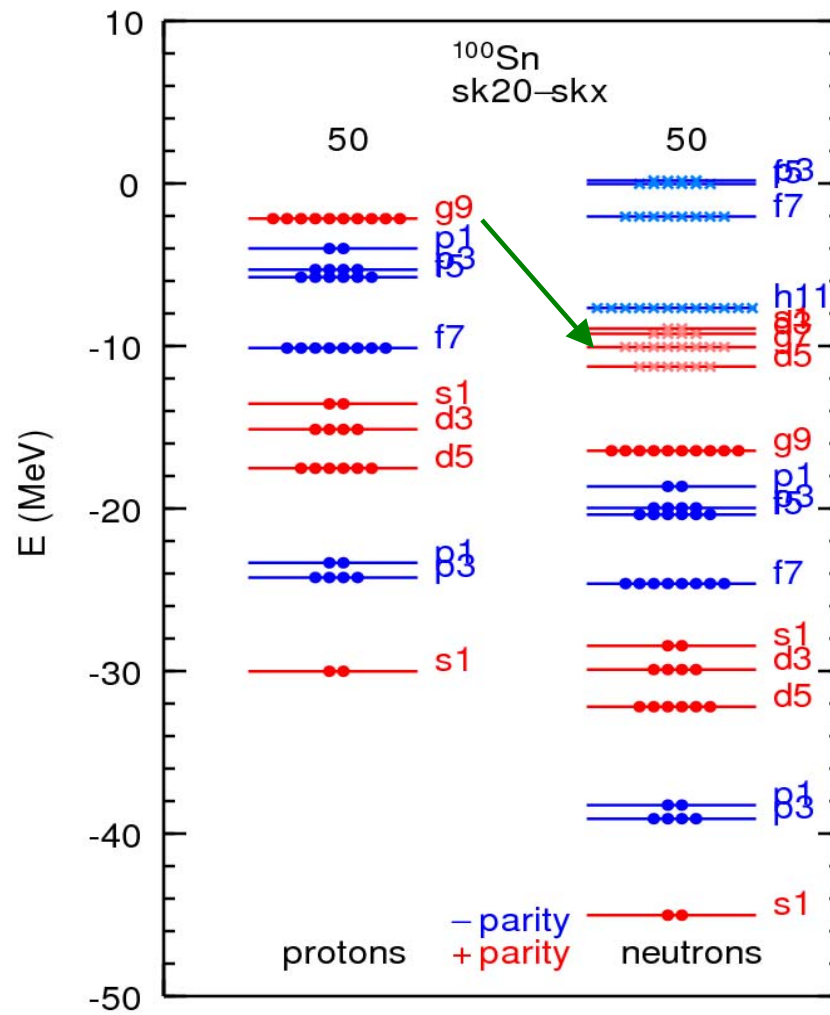


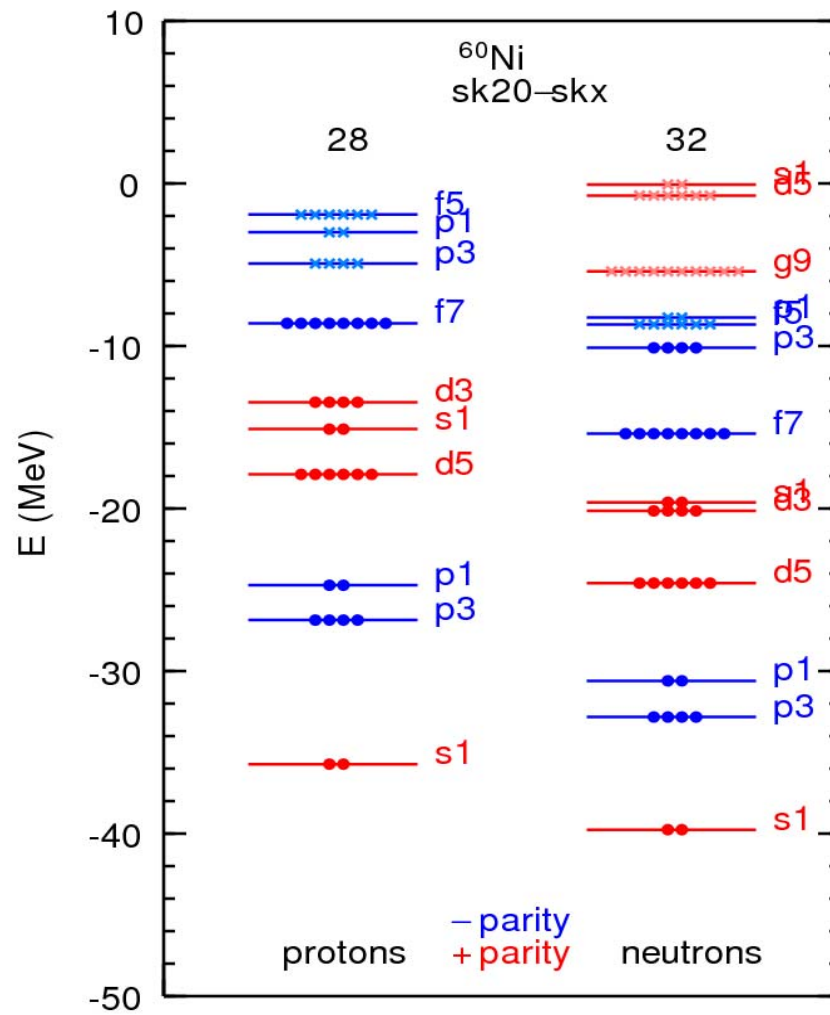






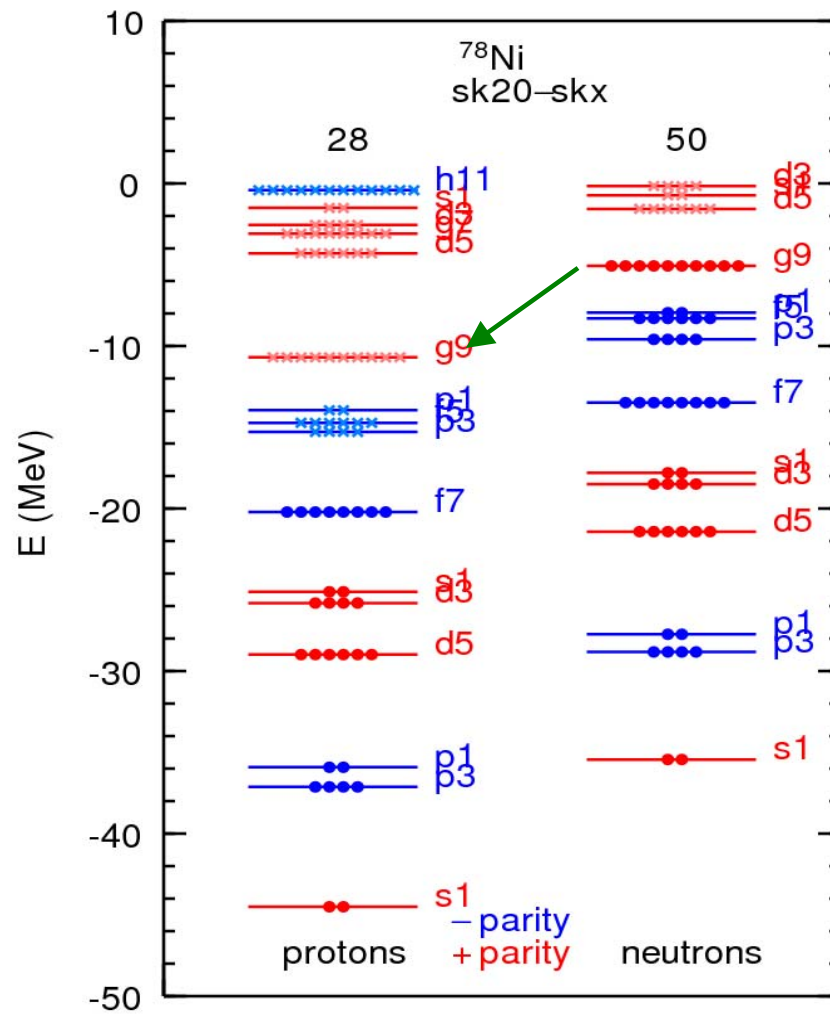












*Nuclear Physics* **A322** (1979) 61–91; © North-Holland Publishing Co., Amsterdam

Not to be reproduced by photoprint or microfilm without written permission from the publisher

**CHARGE-DEPENDENT TWO-BODY INTERACTIONS  
DEDUCED FROM DISPLACEMENT ENERGIES  
IN THE  $1f_{7/2}$  SHELL †**

B. A. BROWN ††

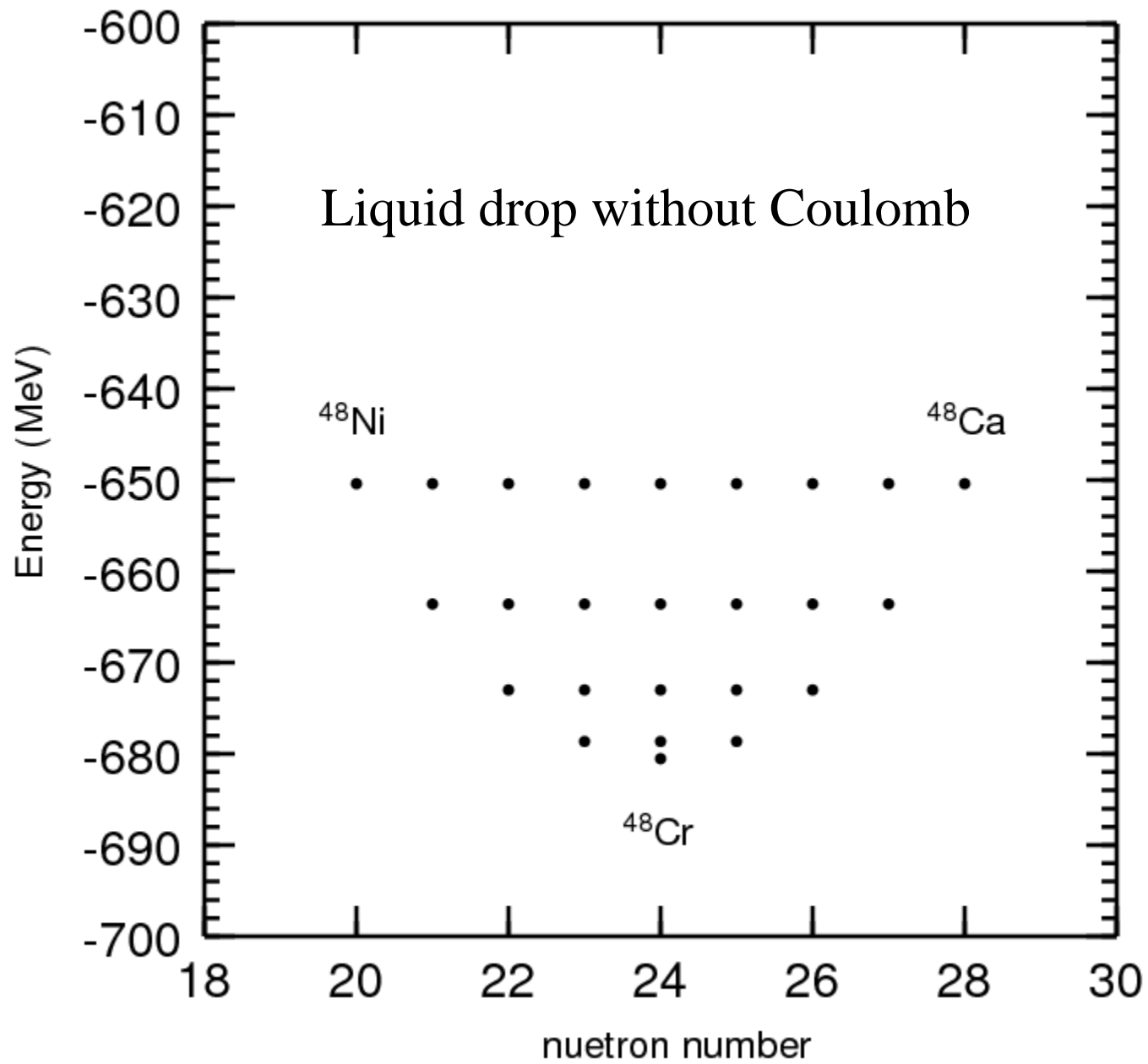
*Cyclotron Laboratory, Michigan State University, East Lansing, Michigan 48824*

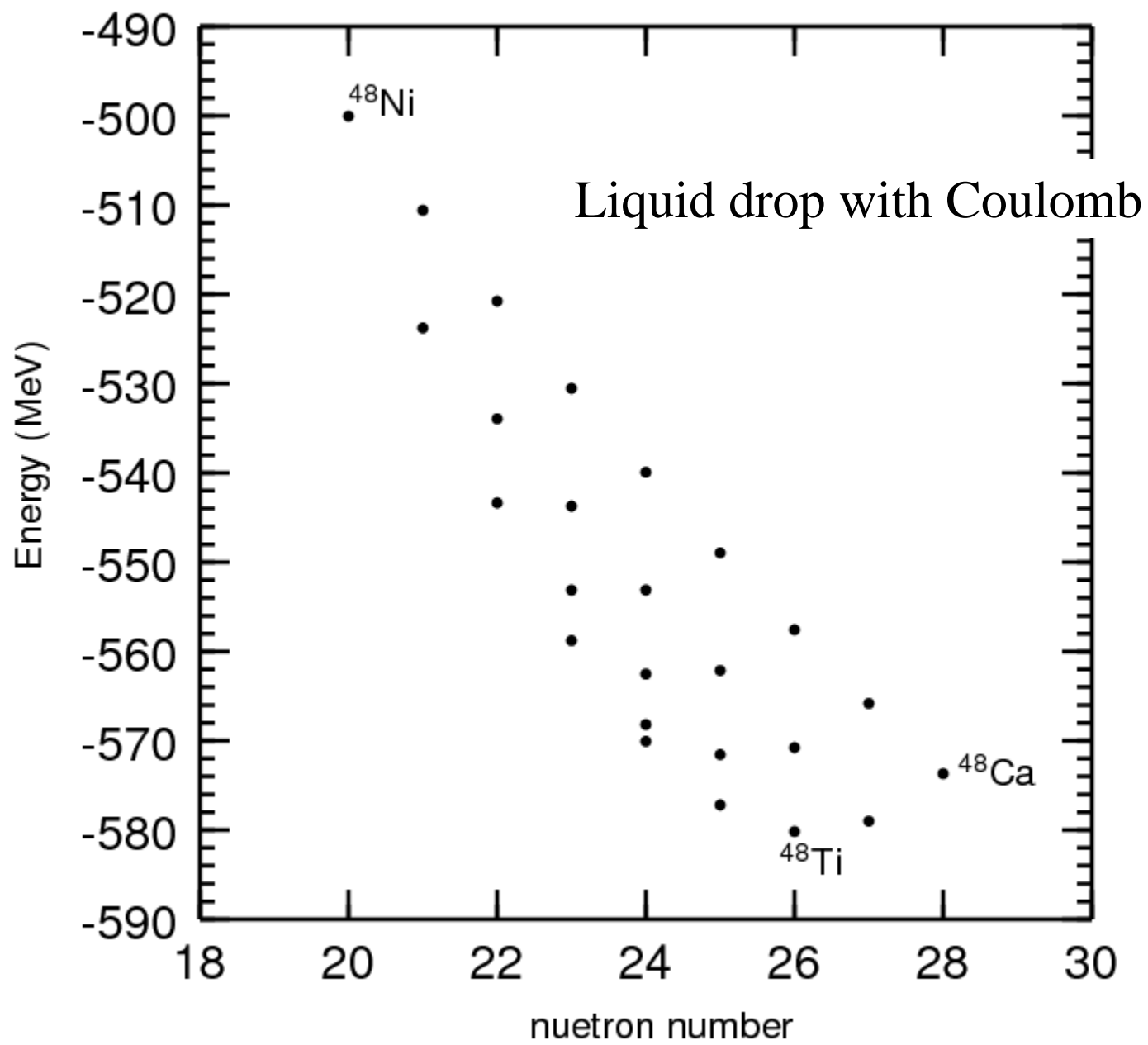
and

R. SHERR

*Joseph Henry Laboratories, Princeton University, Princeton, New Jersey 08540*



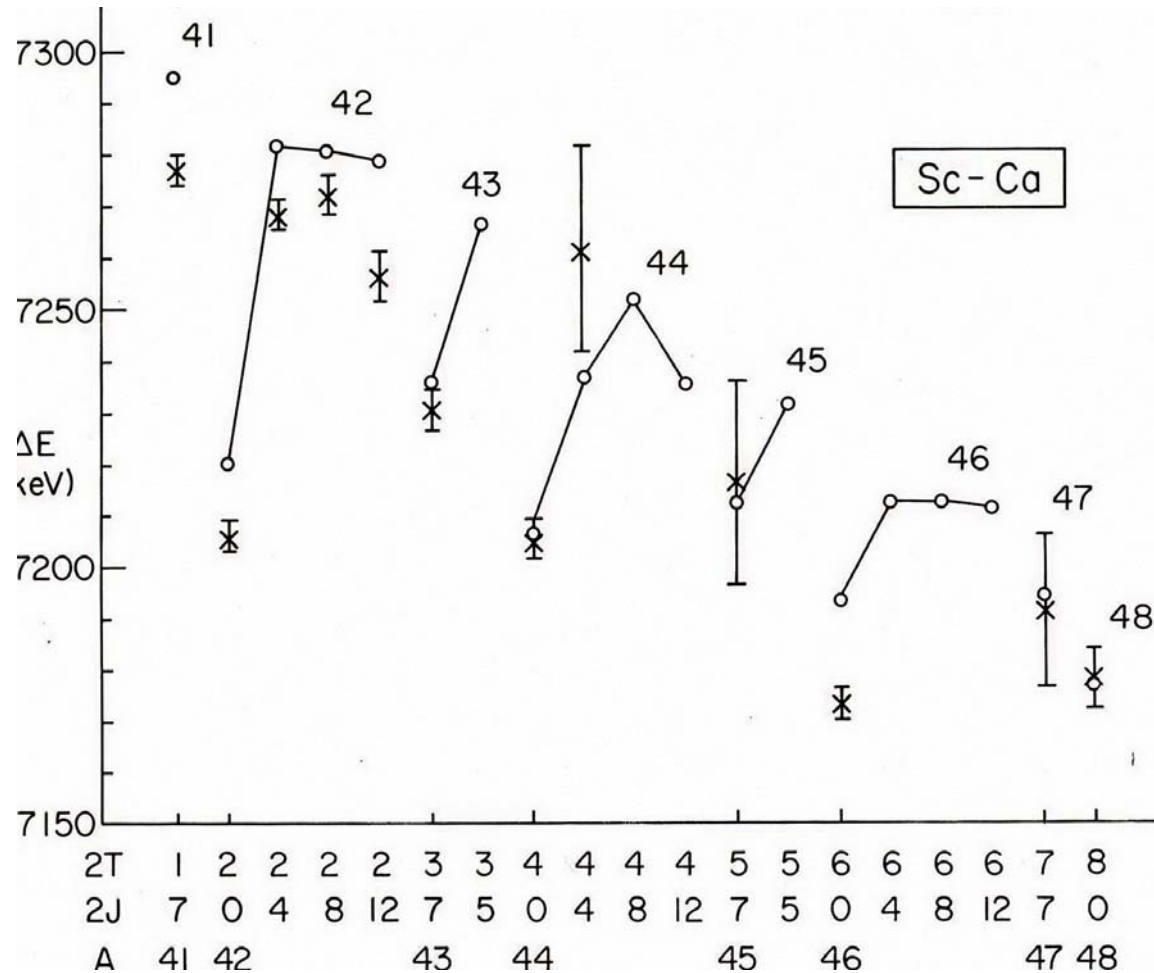


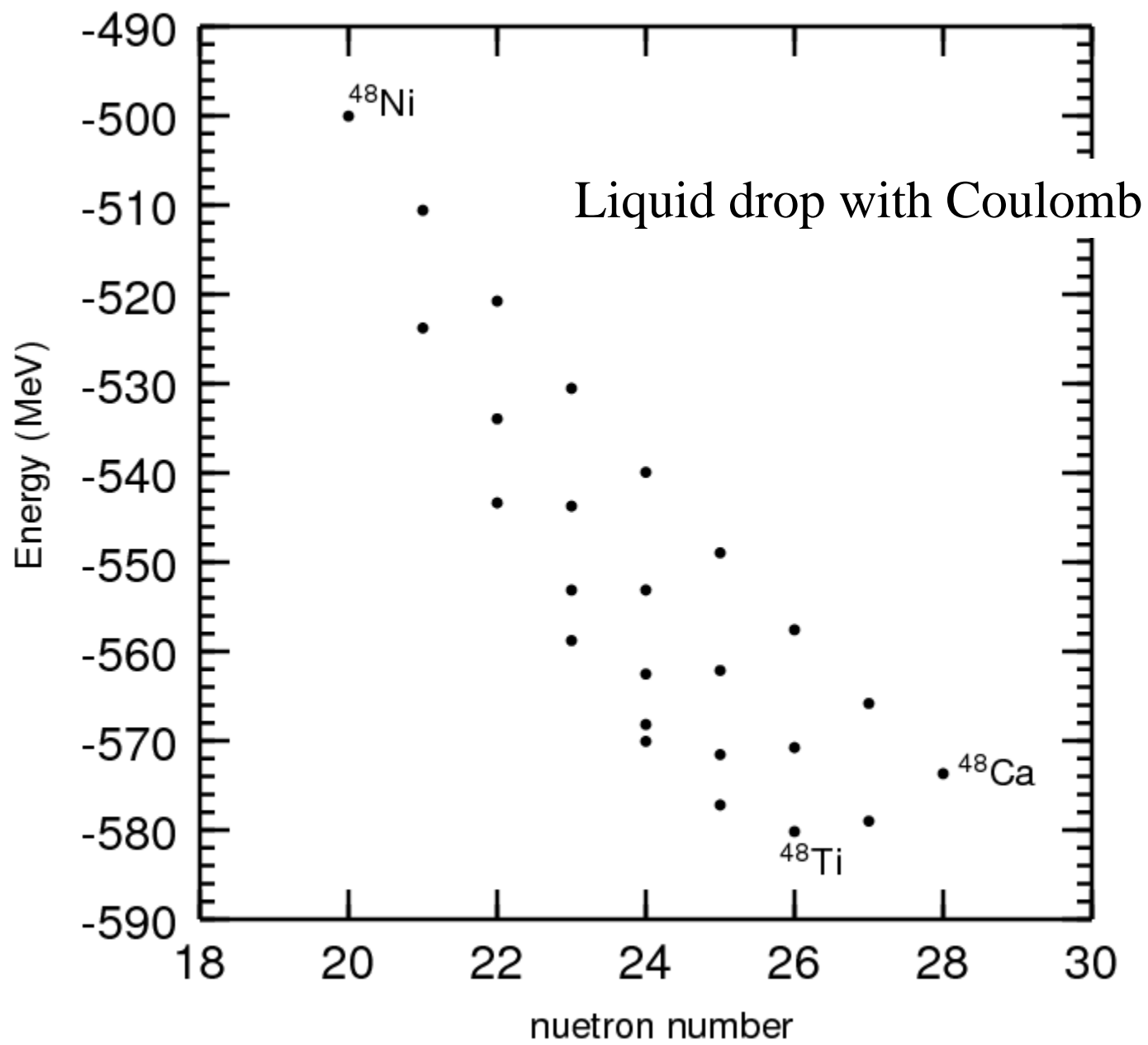




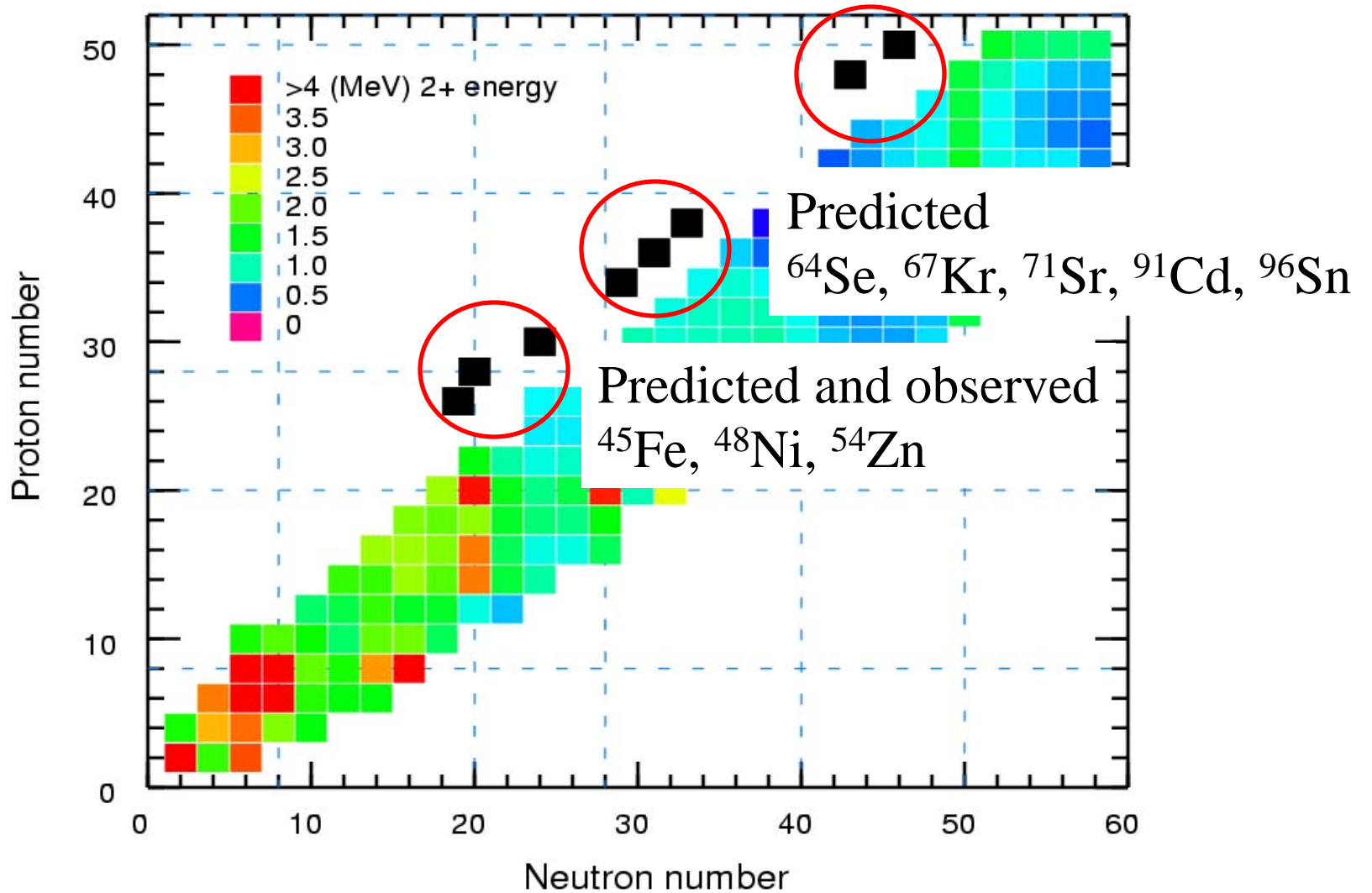
About 60 energies of isobaric analogue states for  $A=41-55$  measured to few keV accuracy

$(f_{7/2})^n$  models with 8 parameters fit all of them to 12 keV rms





# Di-proton decay



## Q-value results for $^{48}\text{Ni}$

$Q(\text{exp}) = 1.35 (2) \text{ MeV}$  Dossat et al. 2006 (one event)

$Q(\text{th}) = 1.36 (13)$  Brown 1991 (IMME)

3.1 (6) Audi-Wapstra extrapolation 2003

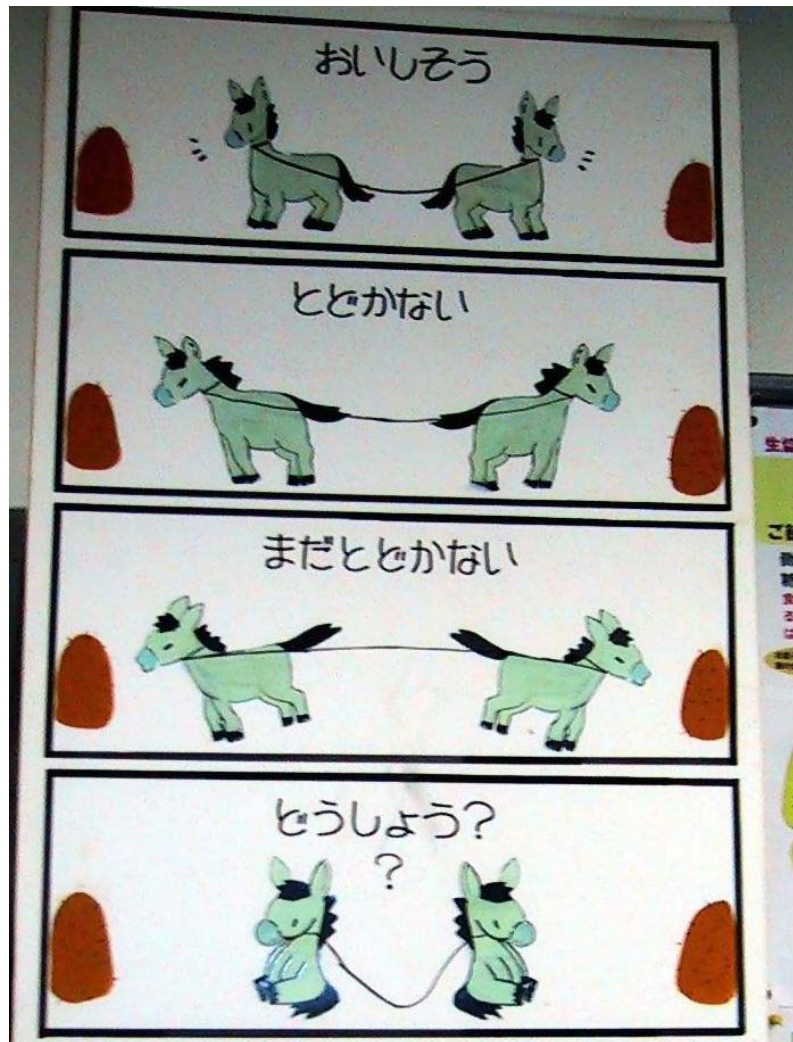
0.0-2.0 Nazarewicz et al 1996 (mean-field)



Di-proton decay = horses to hay

What is the cluster dynamics of di-proton decay ?

$^{48}\text{Ni}$  di-proton vs  $^{48}\text{Ca}(p,t)^{46}\text{Ca}$  ?



## Results for di-proton decay

Nucleus	exp $Q_{2p}$ (MeV)	exp (a) $T_{1/2}$ (ms)	theory (b) $T_{1/2}$ (ms)	S (c)
$^{45}\text{Fe}$	1.151(15)	2.4 - 3.9	14 - 22 (d)	0.272
$^{48}\text{Ni}$	1.35(2)	<21	4 - 11	0.188
$^{54}\text{Zn}$	1.48(2)	2.7 - 5.9	3 - 8	0.313

(a) B. Blank, J. Giovinazzo, M. Pfutzner.....

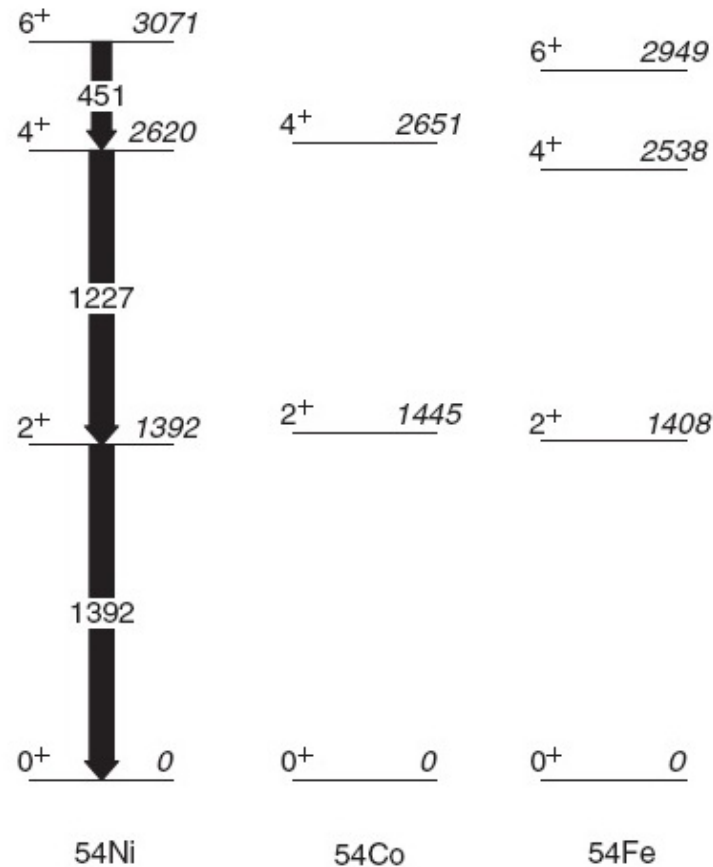
(b) B. A. Brown and F. C. Barker, Phys. Rev. C67, 041304(R) (2003).  
includes correlations (pairing) – three-body Coulomb asymptotics  
in R matrix with pp resonance as an intermediate state

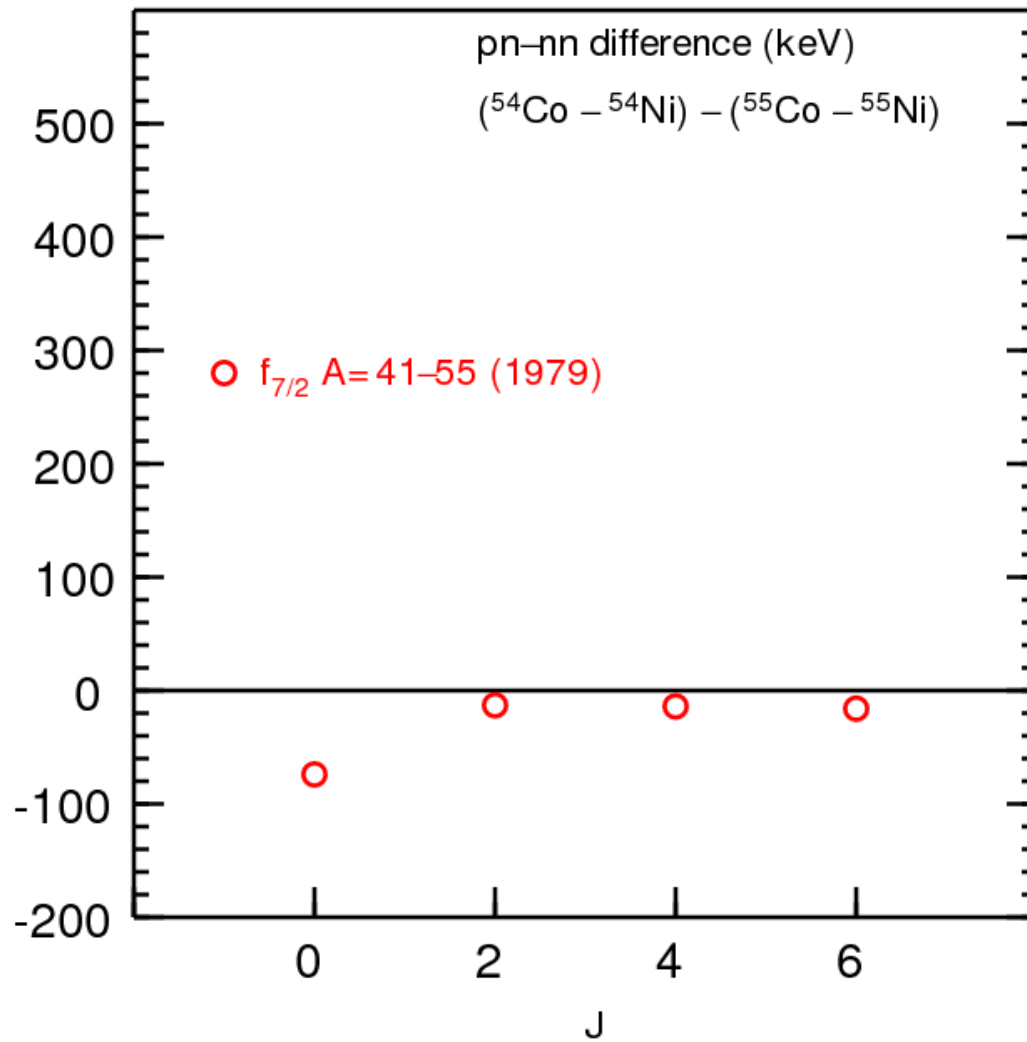
(c) Two-particle spectroscopic factor

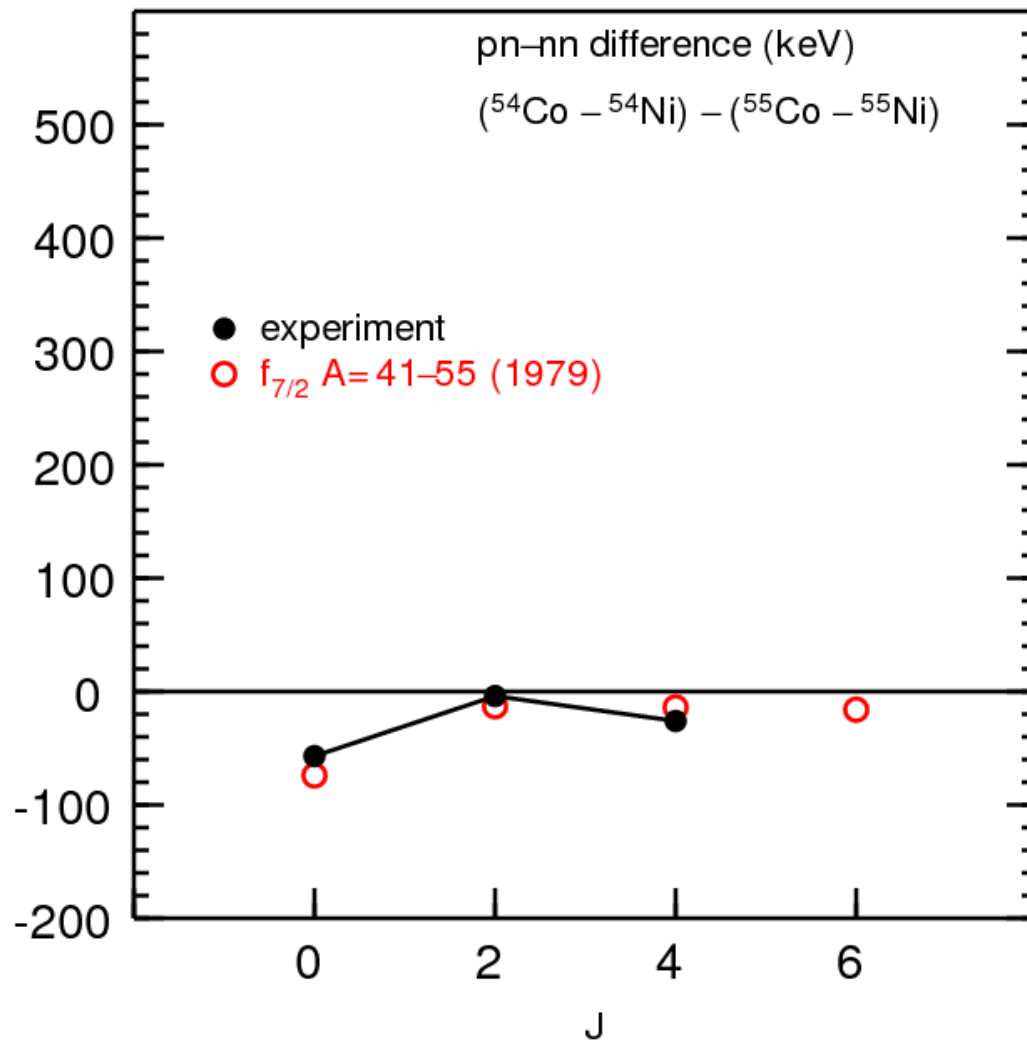
Grigorenko and Zukov, Phys. Rev. C, 68, 054005 (2003).  
single-particle model (no correlations)  
but includes full three-body decay with Coulomb

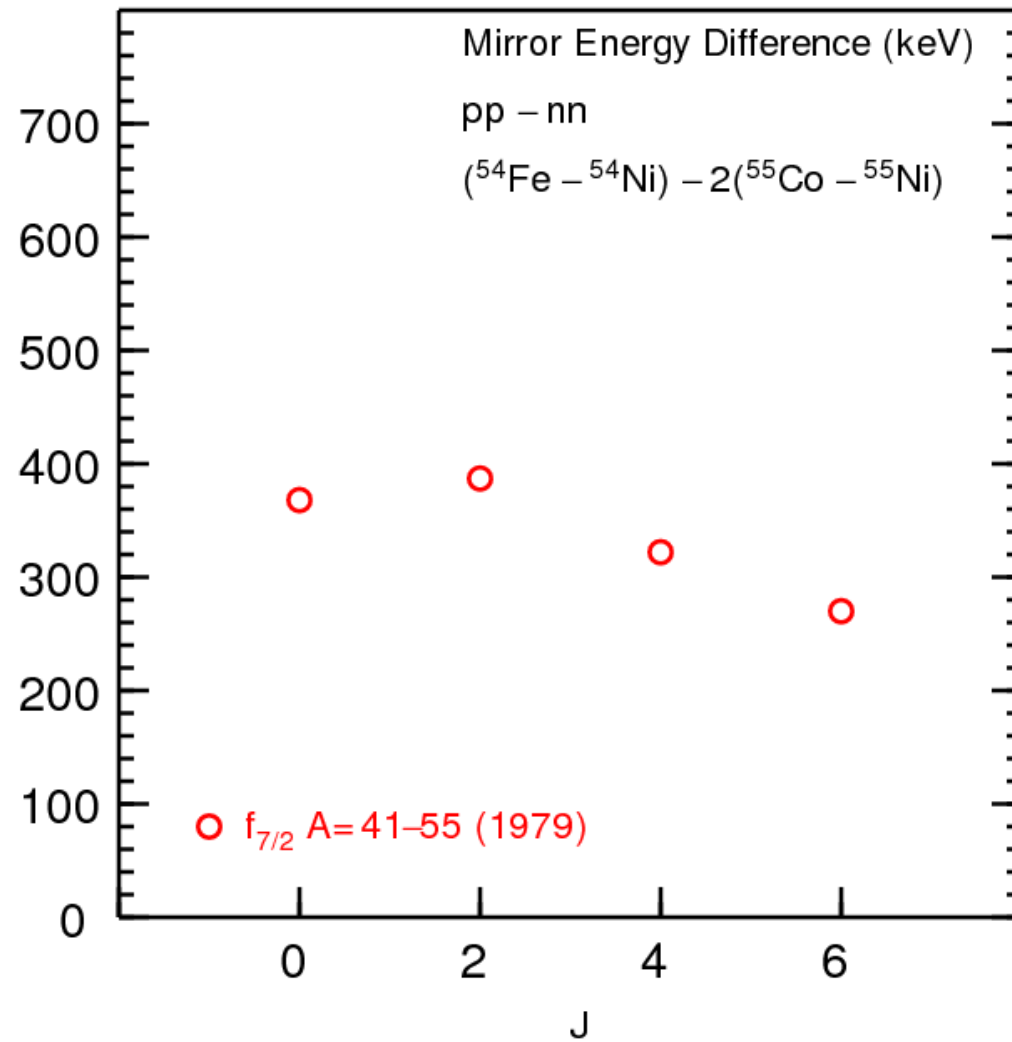
### Observation of $^{54}\text{Ni}$ : Cross-Conjugate Symmetry in $f_{7/2}$ Mirror Energy Differences

A. Gadea,<sup>1</sup> S. M. Lenzi,<sup>2</sup> S. Lunardi,<sup>2</sup> N. Mărginean,<sup>1,3</sup> A. P. Zuker,<sup>4</sup> G. de Angelis,<sup>1</sup> M. Axiotis,<sup>1</sup> T. Martínez,<sup>1</sup> D. R. Napoli,<sup>1</sup> E. Farnea,<sup>2</sup> R. Menegazzo,<sup>2</sup> P. Pavan,<sup>2</sup> C. A. Ur,<sup>2,3</sup> D. Bazzacco,<sup>2</sup> R. Venturelli,<sup>2</sup> P. Kleinheinz,<sup>5</sup> P. Bednarczyk,<sup>4,6</sup> D. Curien,<sup>4</sup> O. Dorvaux,<sup>4</sup> J. Nyberg,<sup>7</sup> H. Grawe,<sup>8</sup> M. Górska,<sup>8</sup> M. Palacz,<sup>9</sup> K. Lagergren,<sup>10</sup> L. Milechina,<sup>10</sup> J. Ekman,<sup>11</sup> D. Rudolph,<sup>12</sup> C. Andreoiu,<sup>12</sup> M. A. Bentley,<sup>13</sup> W. Gelletly,<sup>14</sup> B. Rubio,<sup>15</sup> A. Algora,<sup>15</sup> E. Nacher,<sup>15</sup> L. Caballero,<sup>15</sup> M. Trotta,<sup>16</sup> and M. Moszvíski<sup>17</sup>

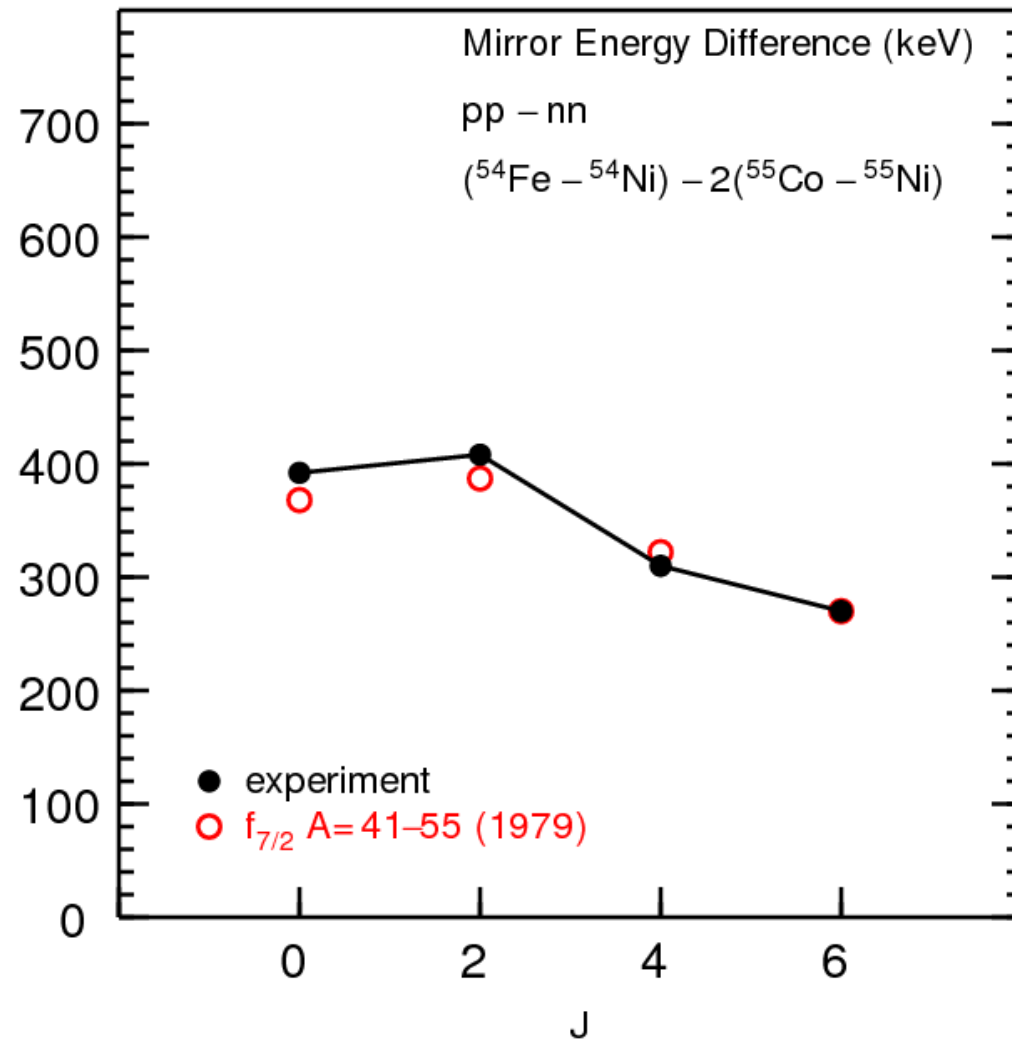




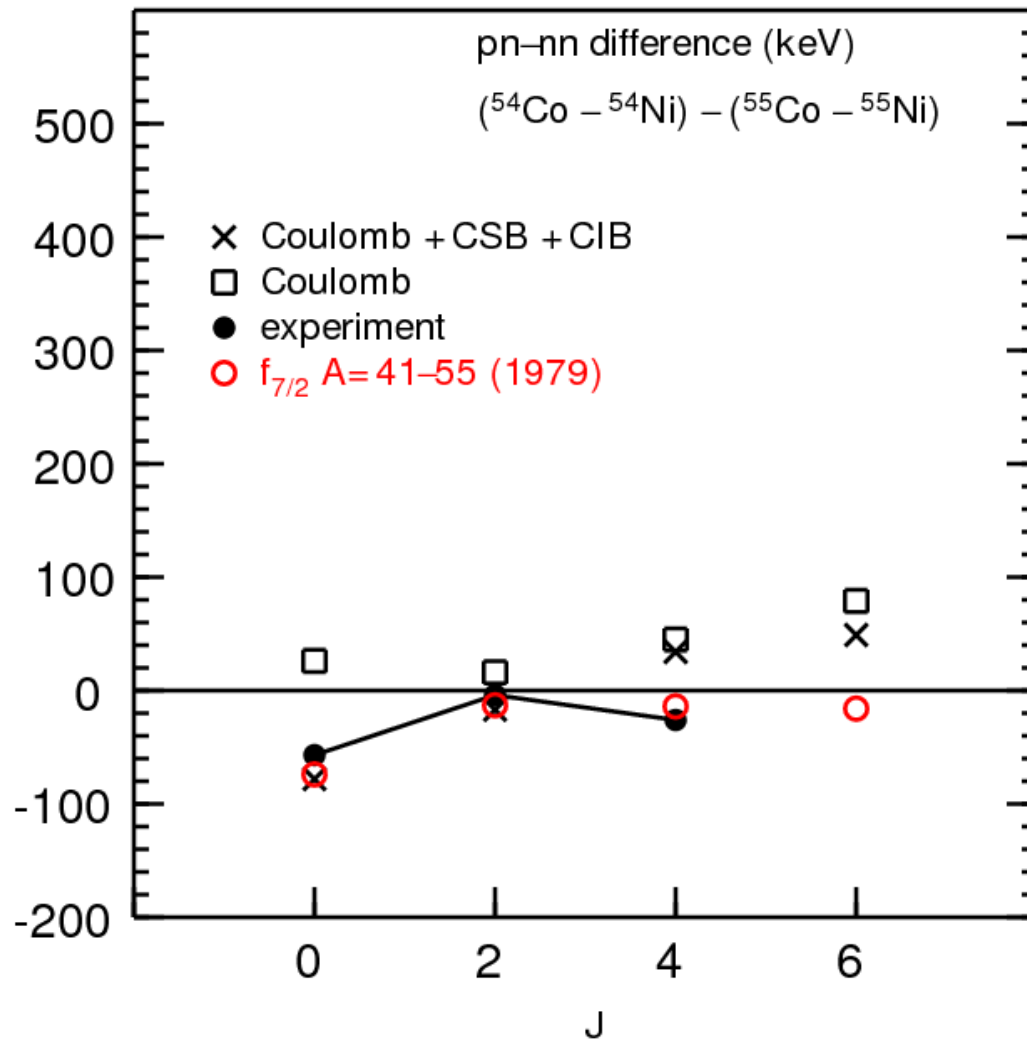




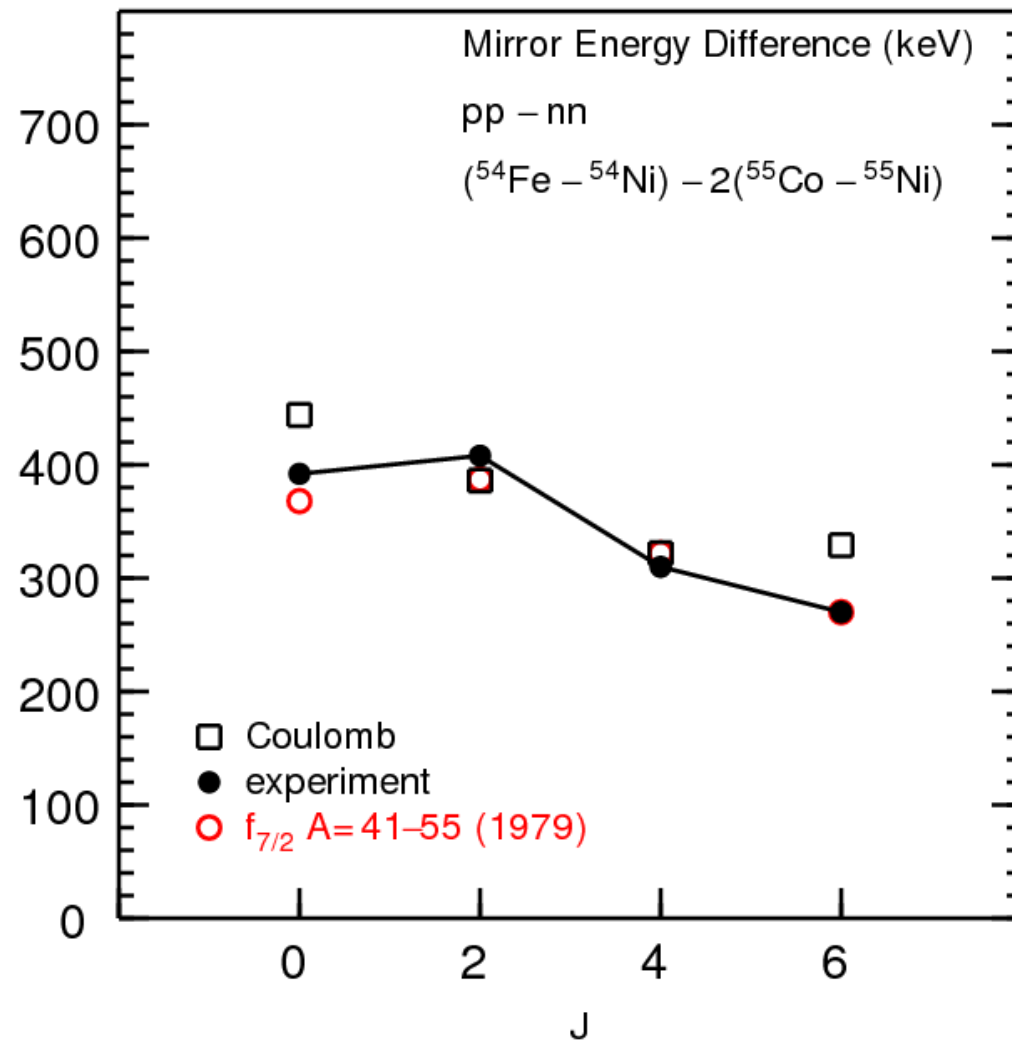




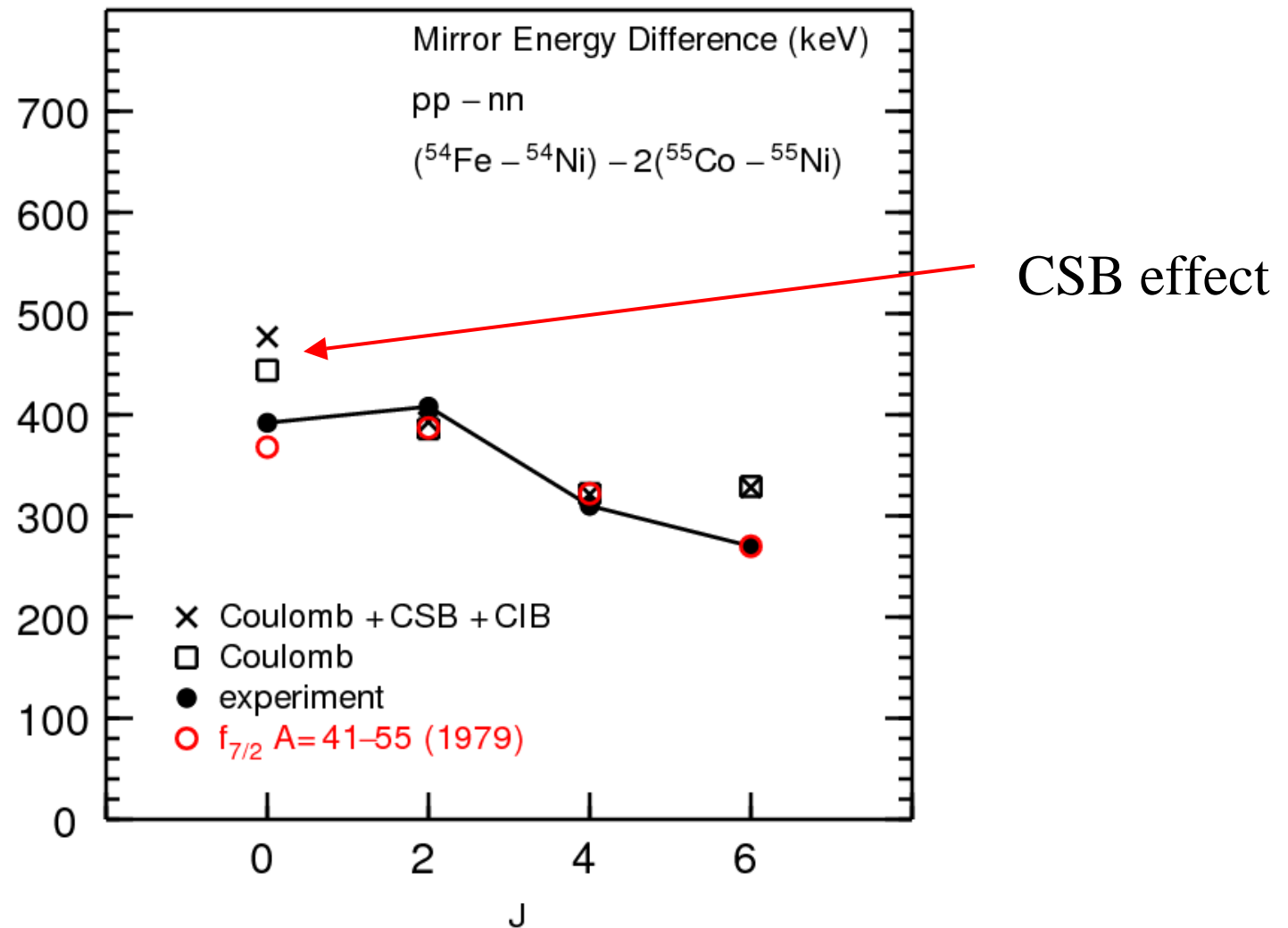
# N3LO V-lowk 6hw 2<sup>nd</sup> order – “full” pf shell (Angelo Signoracci)



# N3LO V-lowk 6hw 2<sup>nd</sup> order – “full” pf shell (Angelo Signoracci)



# N3LO V-lowk 6hw 2<sup>nd</sup> order – “full” pf shell (Angelo Signoracci)



	N <sup>3</sup> LO <sup>a</sup>	Experiment <sup>b</sup>
		<sup>1</sup> S <sub>0</sub>
$a_{pp}^C$	-7.8188	-7.8196±0.0026
$r_{pp}^C$	2.795	2.790±0.014
$a_{pp}^N$	-17.083	
$r_{pp}^N$	2.876	
$a_{nn}^N$	-18.900	-18.9±0.4
$r_{nn}^N$	2.838	2.75±0.11
$a_{np}$	-23.732	-23.740±0.020
$r_{np}$	2.725	2.77±0.05
		<sup>3</sup> S <sub>1</sub>
$a_t$	5.417	5.419±0.007
$r_t$	1.752	1.753±0.008

From n+d

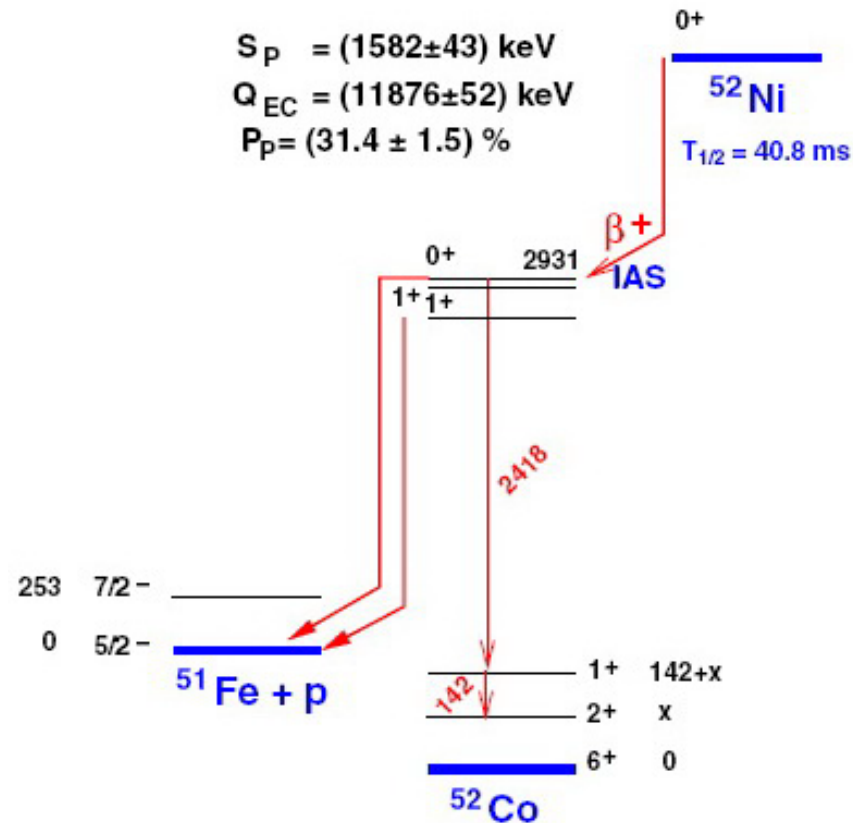
-18.7(6) PRL 83, 3788 (1999)

-16.3(40) PRL 85, 1190 (2000)



# Bertram Blank – isospin forbidden proton decay

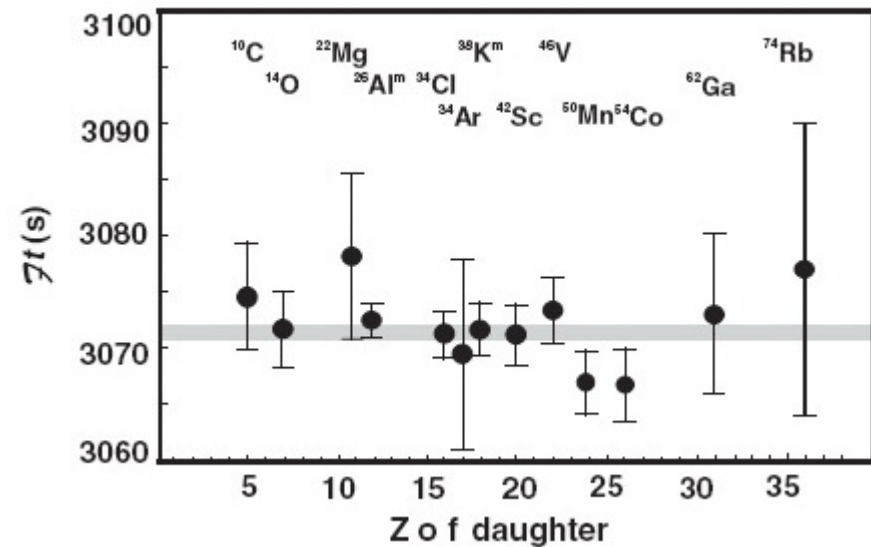
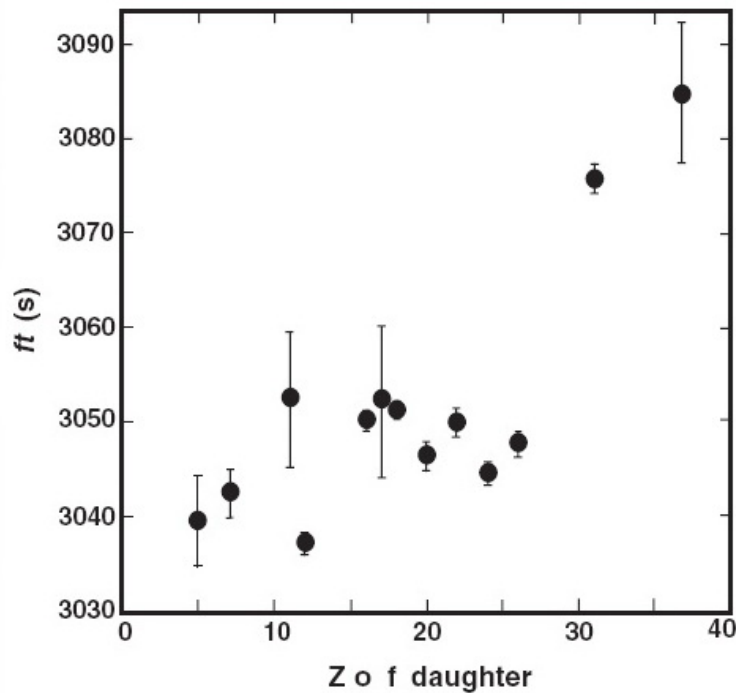
$^{52}\text{Co } 0^+ T=4$  to  $^{51}\text{Fe } 5/2^- T=1/2$



## Improved calculation of the isospin-symmetry-breaking corrections to superallowed Fermi $\beta$ decay

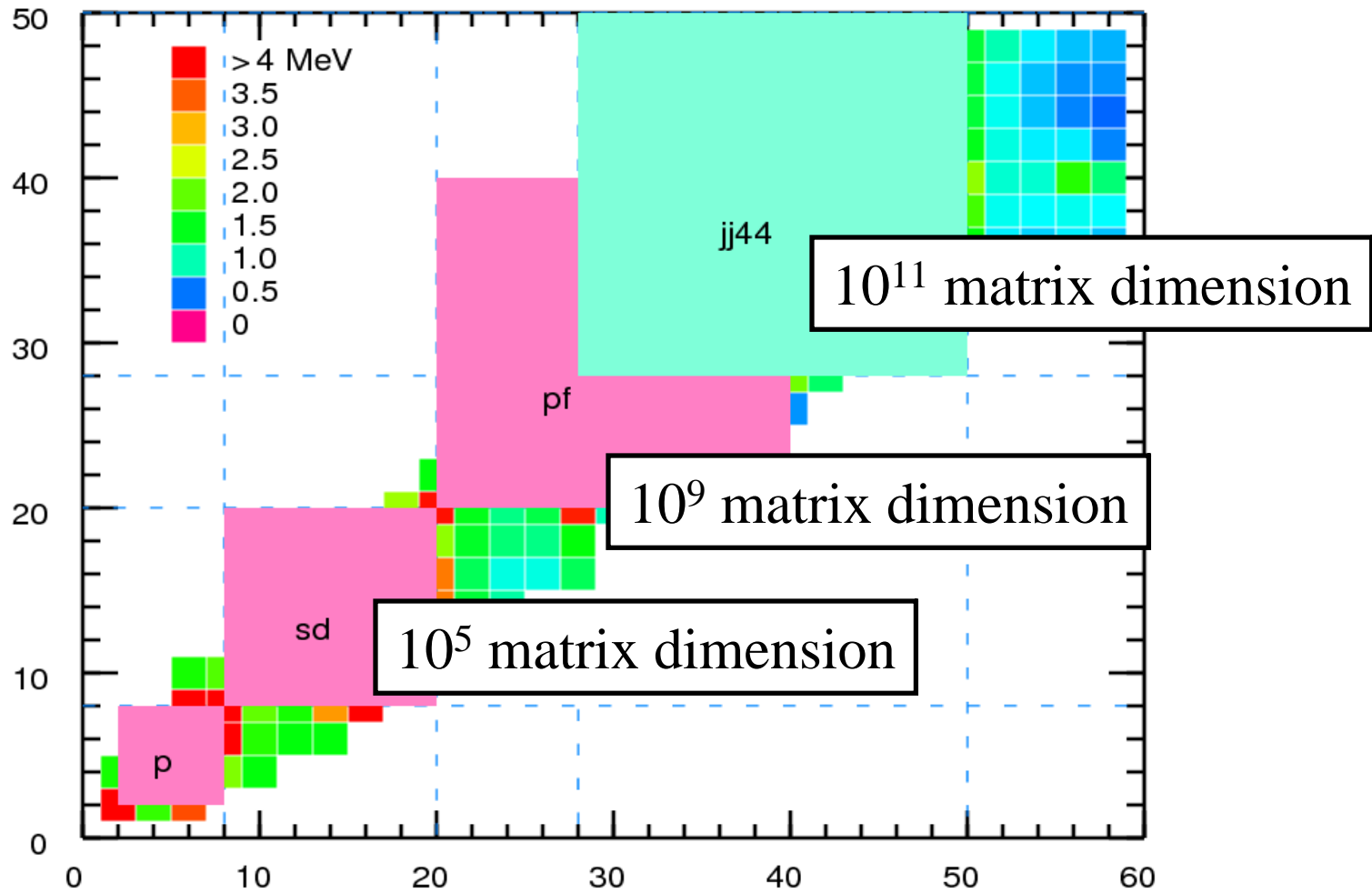
I. S. Towner\* and J. C. Hardy

Cyclotron Institute, Texas A&M University, College Station, Texas 77843, USA



$$|V_{ud}|^2 + |V_{us}|^2 + |V_{ub}|^2 = 1.0000 \pm 0.0011.$$



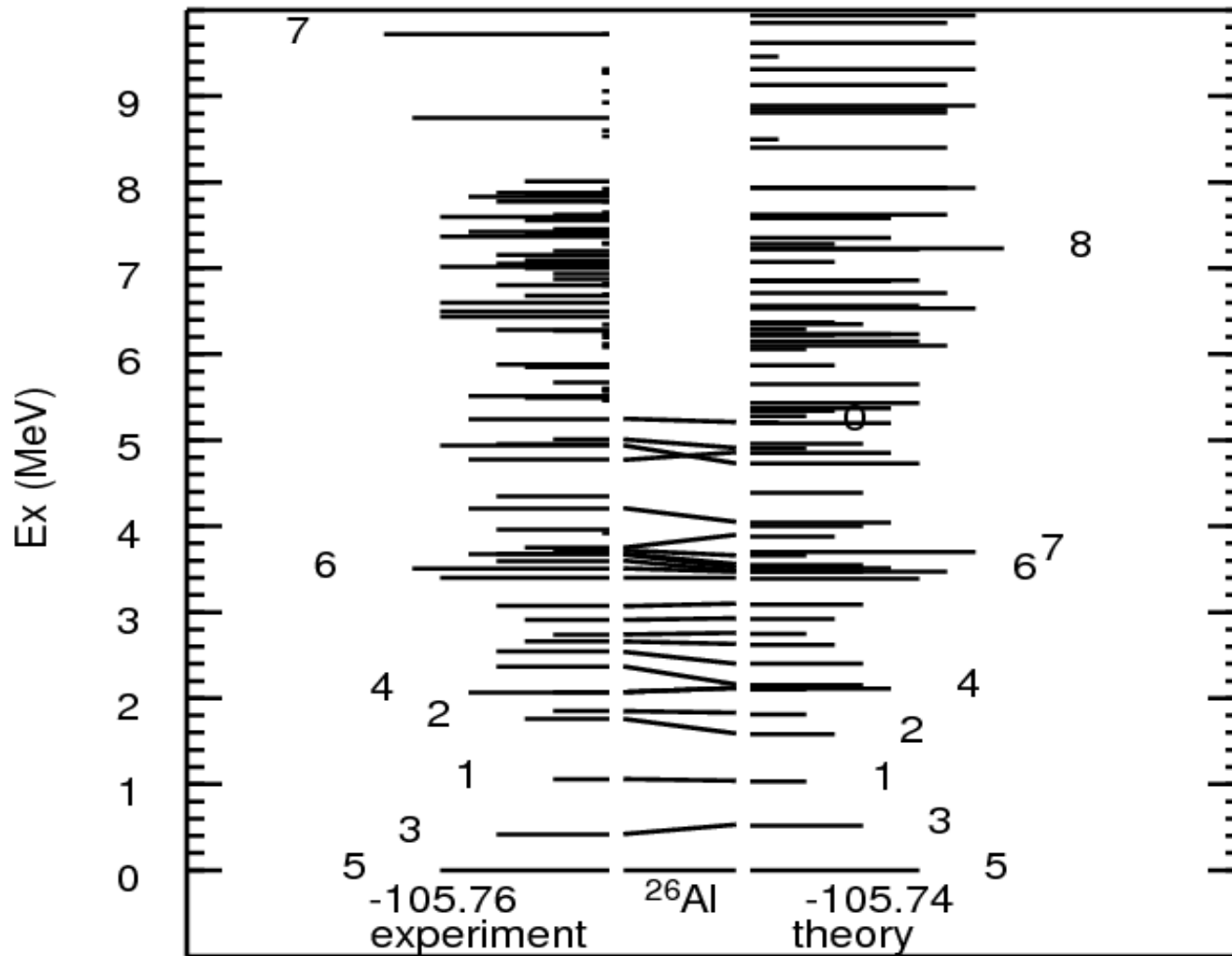


jj44 means  $f_{5/2}$ ,  $p_{3/2}$ ,  $p_{1/2}$ ,  $g_{9/2}$  orbits for protons and neutrons

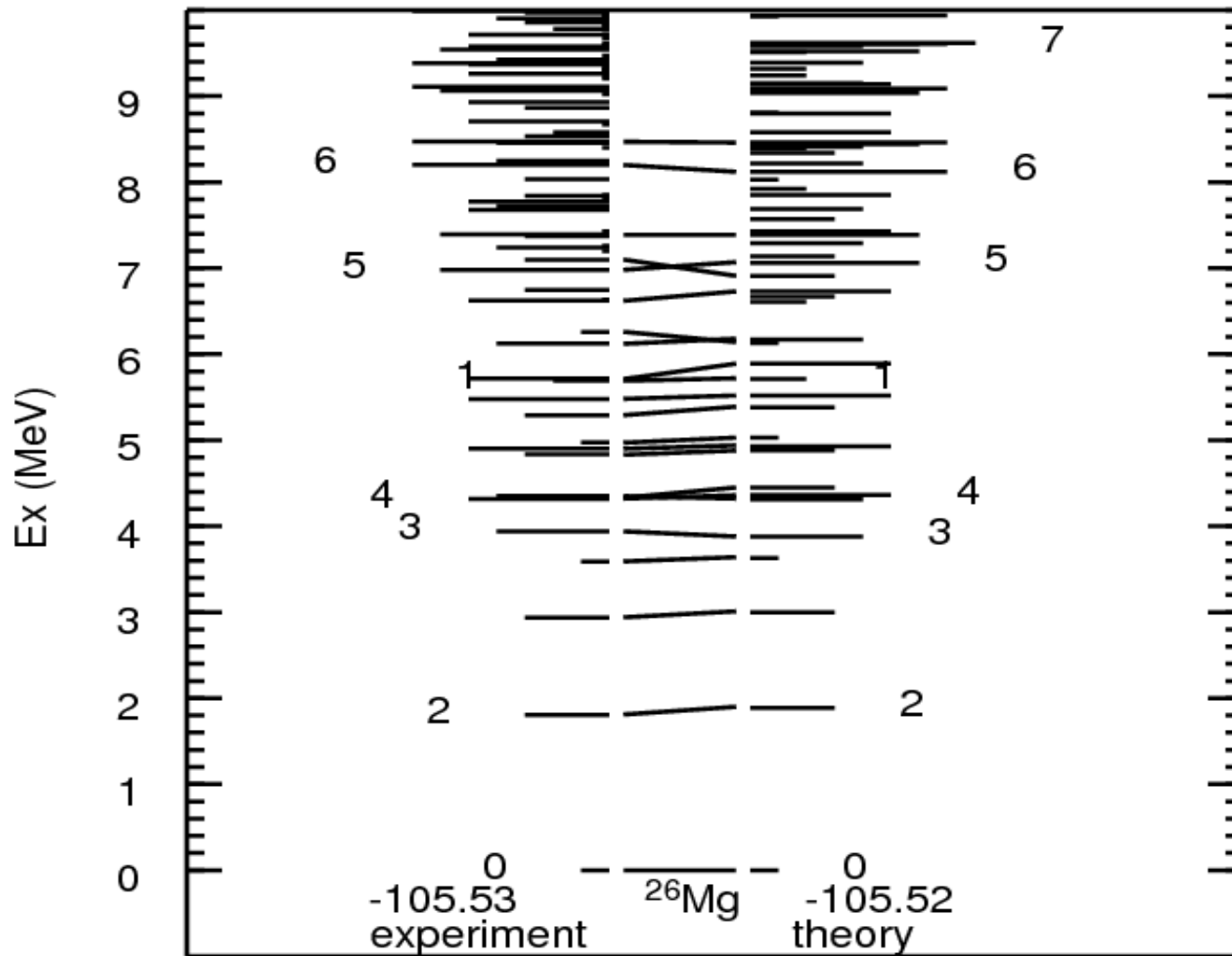
# A tour of the sd shell on the web

	Home	Bookmarks											
K											<a href="#">38, 0</a>	<a href="#">39, 1/2</a>	
Ar											<a href="#">36, 0</a>	<a href="#">37, 1/2</a>	<a href="#">38, 1</a>
Cl										<a href="#">34, 0</a>	<a href="#">35, 1/2</a>	<a href="#">36, 1</a>	<a href="#">37, 3/2</a>
S								<a href="#">32, 0</a>	<a href="#">33, 1/2</a>	<a href="#">34, 1</a>	<a href="#">35, 3/2</a>	<a href="#">36, 2</a>	
P							<a href="#">30, 0</a>	<a href="#">31, 1/2</a>	<a href="#">32, 1</a>	<a href="#">33, 3/2</a>	<a href="#">34, 2</a>	<a href="#">35, 5/2</a>	
Si						<a href="#">28, 0</a>	<a href="#">29, 1/2</a>	<a href="#">30, 1</a>	<a href="#">31, 3/2</a>	<a href="#">32, 2</a>	<a href="#">33, 5/2</a>	<a href="#">34, 3</a>	
Al					<a href="#">26, 0</a>	<a href="#">27, 1/2</a>	<a href="#">28, 1</a>	<a href="#">29, 3/2</a>	<a href="#">30, 2</a>	<a href="#">31, 5/2</a>	<a href="#">32, 3</a>	<a href="#">33, 7/2</a>	
Mg				<a href="#">24, 0</a>	<a href="#">25, 1/2</a>	<a href="#">26, 1</a>	<a href="#">27, 3/2</a>	<a href="#">28, 2</a>	<a href="#">29, 5/2</a>	<a href="#">30, 3</a>	<a href="#">31, 7/2</a>	<a href="#">32, 4</a>	
Na			<a href="#">22, 0</a>	<a href="#">23, 1/2</a>	<a href="#">24, 1</a>	<a href="#">25, 3/2</a>	<a href="#">26, 2</a>	<a href="#">27, 5/2</a>	<a href="#">28, 3</a>	<a href="#">29, 7/2</a>	<a href="#">30, 4</a>	<a href="#">31, 9/2</a>	
Ne		<a href="#">20, 0</a>	<a href="#">21, 1/2</a>	<a href="#">22, 1</a>	<a href="#">23, 3/2</a>	<a href="#">24, 2</a>	<a href="#">25, 5/2</a>	<a href="#">26, 3</a>	<a href="#">27, 7/2</a>	<a href="#">28, 4</a>	<a href="#">29, 9/2</a>	<a href="#">30, 5</a>	
F	<a href="#">18, 0</a>	<a href="#">19, 1/2</a>	<a href="#">20, 1</a>	<a href="#">21, 3/2</a>	<a href="#">22, 2</a>	<a href="#">23, 5/2</a>	<a href="#">24, 3</a>	<a href="#">25, 7/2</a>	<a href="#">26, 4</a>	<a href="#">27, 9/2</a>	<a href="#">28, 5</a>	<a href="#">29, 11/2</a>	
O	<a href="#">17, 1/2</a>	<a href="#">18, 1</a>	<a href="#">19, 3/2</a>	<a href="#">20, 2</a>	<a href="#">21, 5/2</a>	<a href="#">22, 3</a>	<a href="#">23, 7/2</a>	<a href="#">24, 4</a>	<a href="#">25, 9/2</a>	<a href="#">26, 5</a>	<a href="#">27, 11/2</a>	<a href="#">28, 6</a>	
	9	10	11	12	13	14	15	16	17	18	19	20	

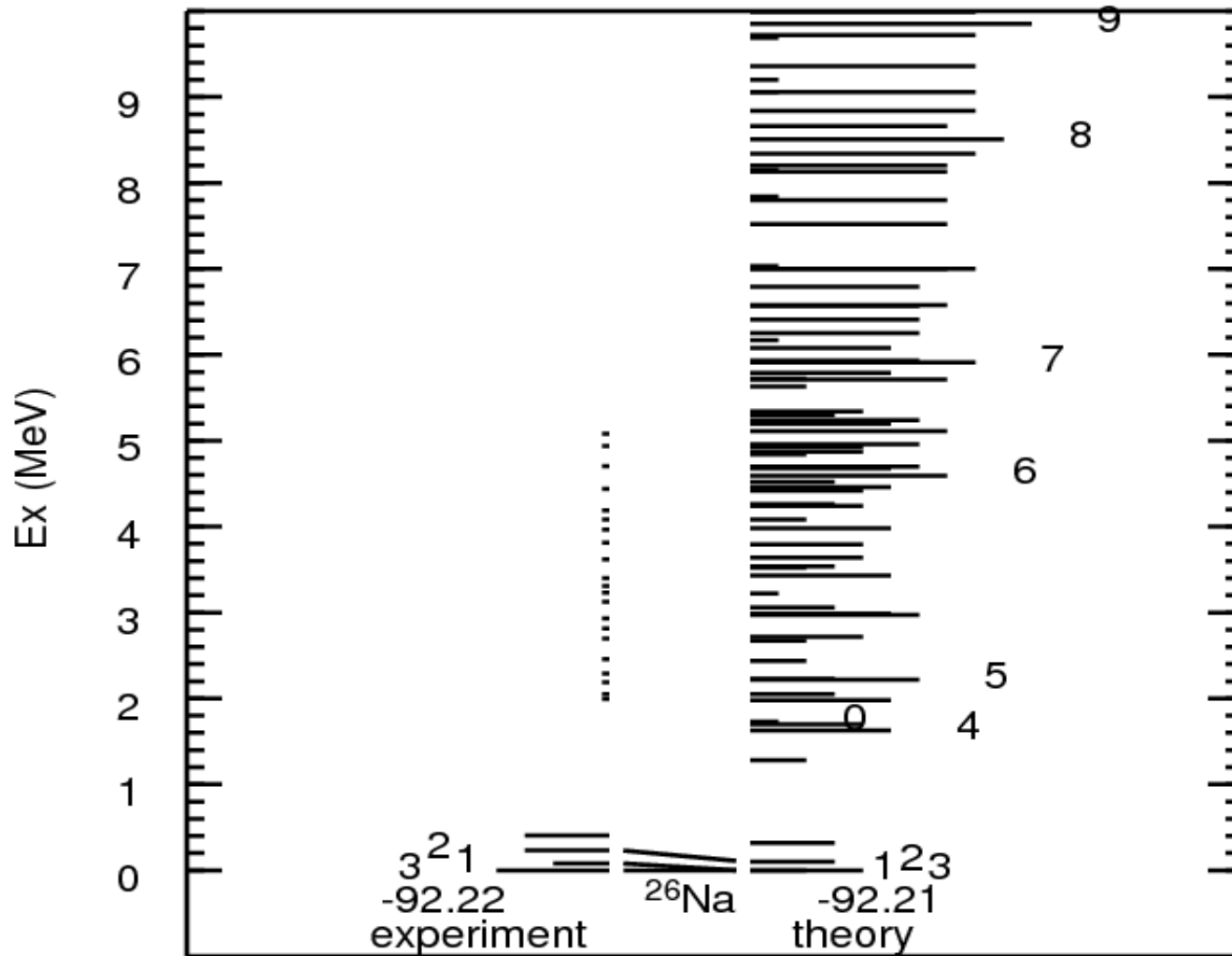
# Positive parity states for $^{26}\text{Al}$



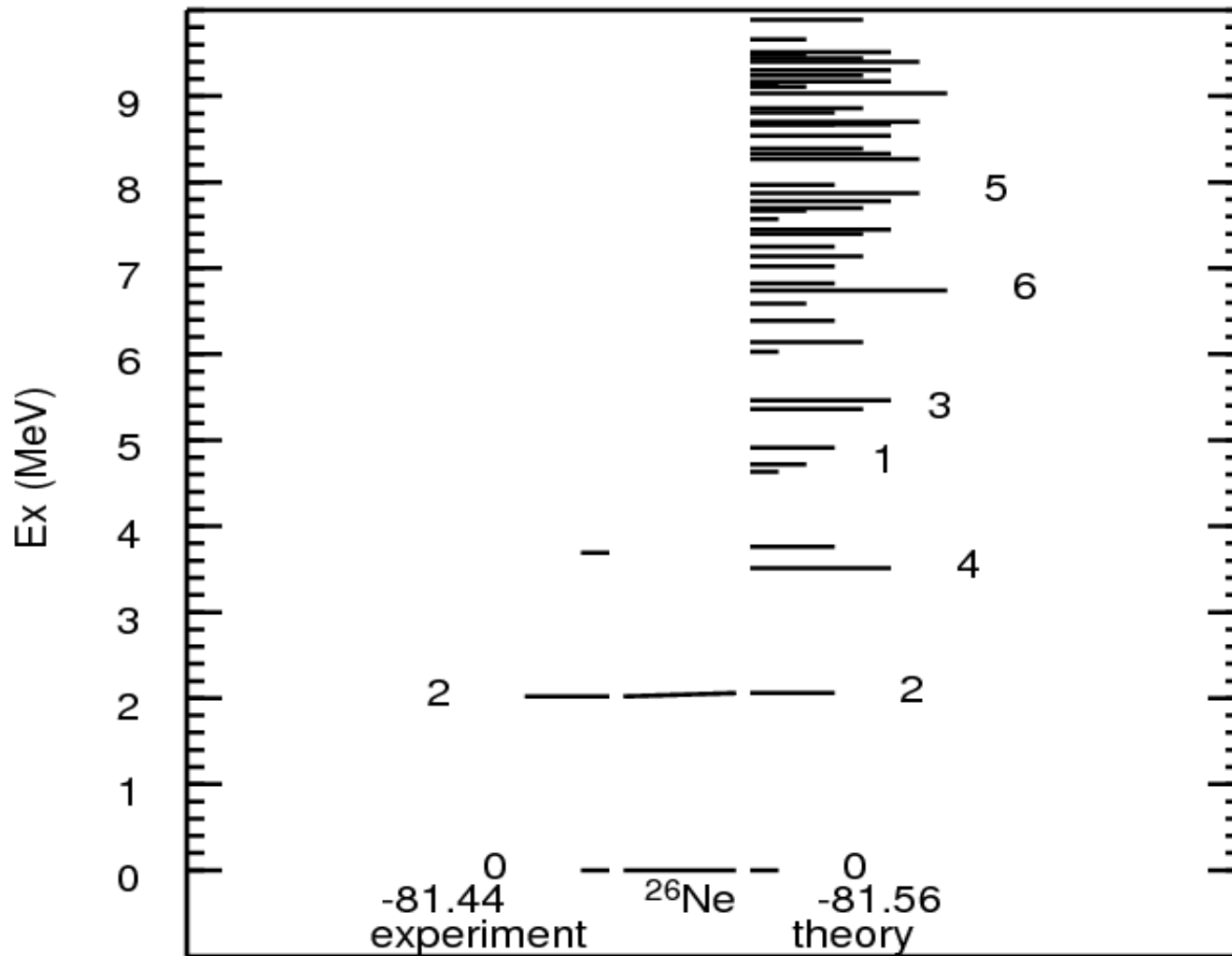
# Positive parity states for $^{26}\text{Mg}$



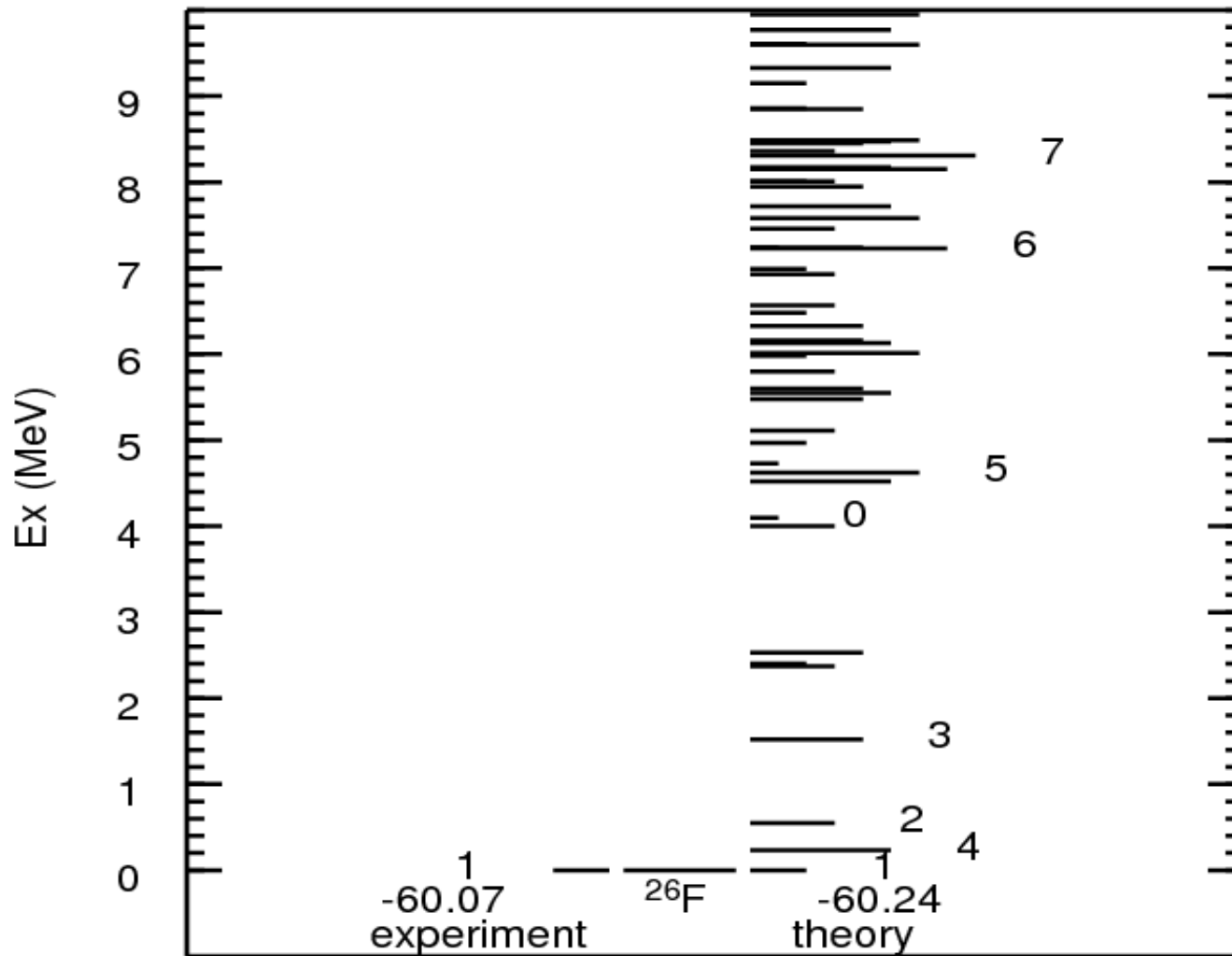
# Positive parity states for $^{26}\text{Na}$



# Positive parity states for $^{26}\text{Ne}$

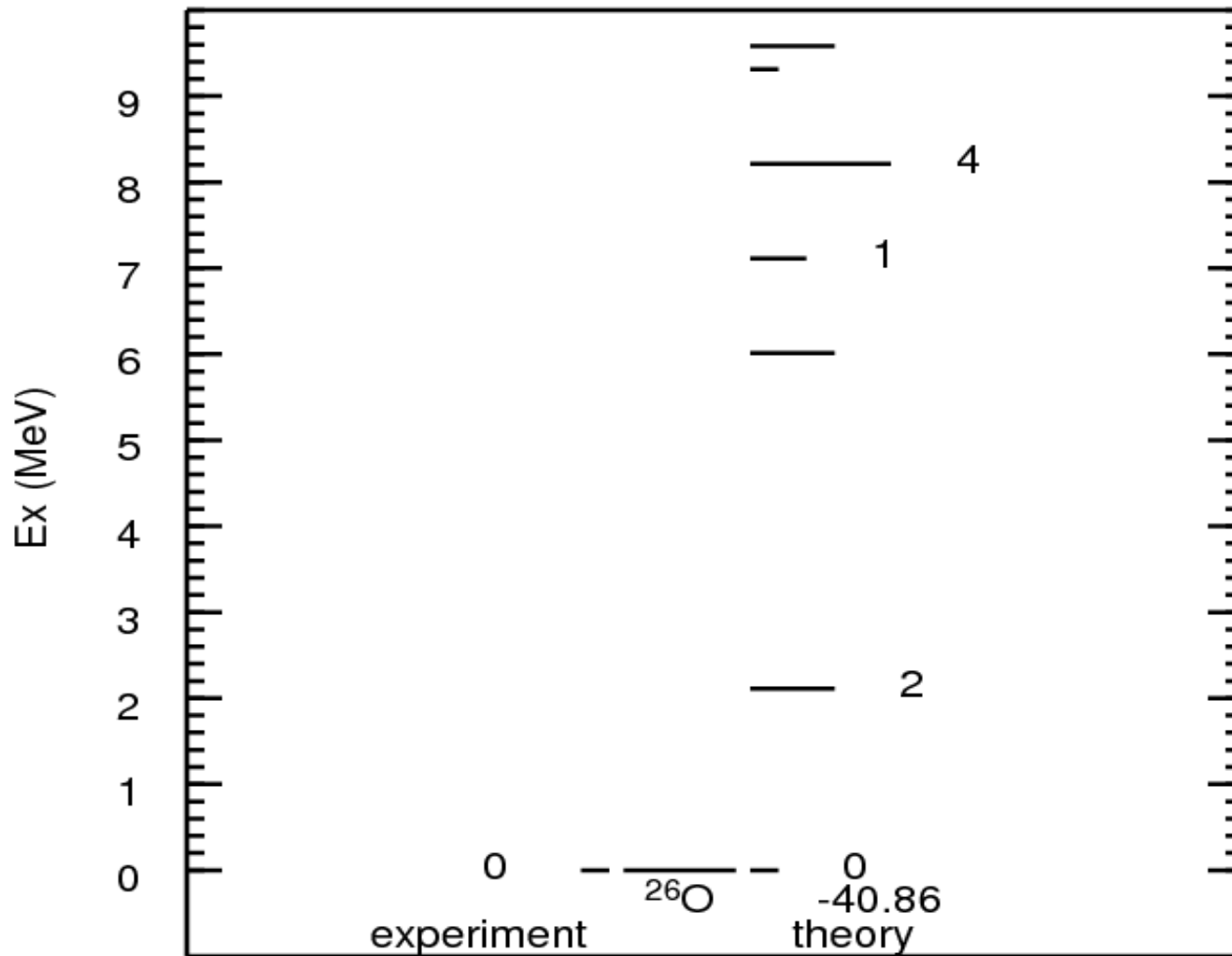


# Positive parity states for $^{26}\text{F}$





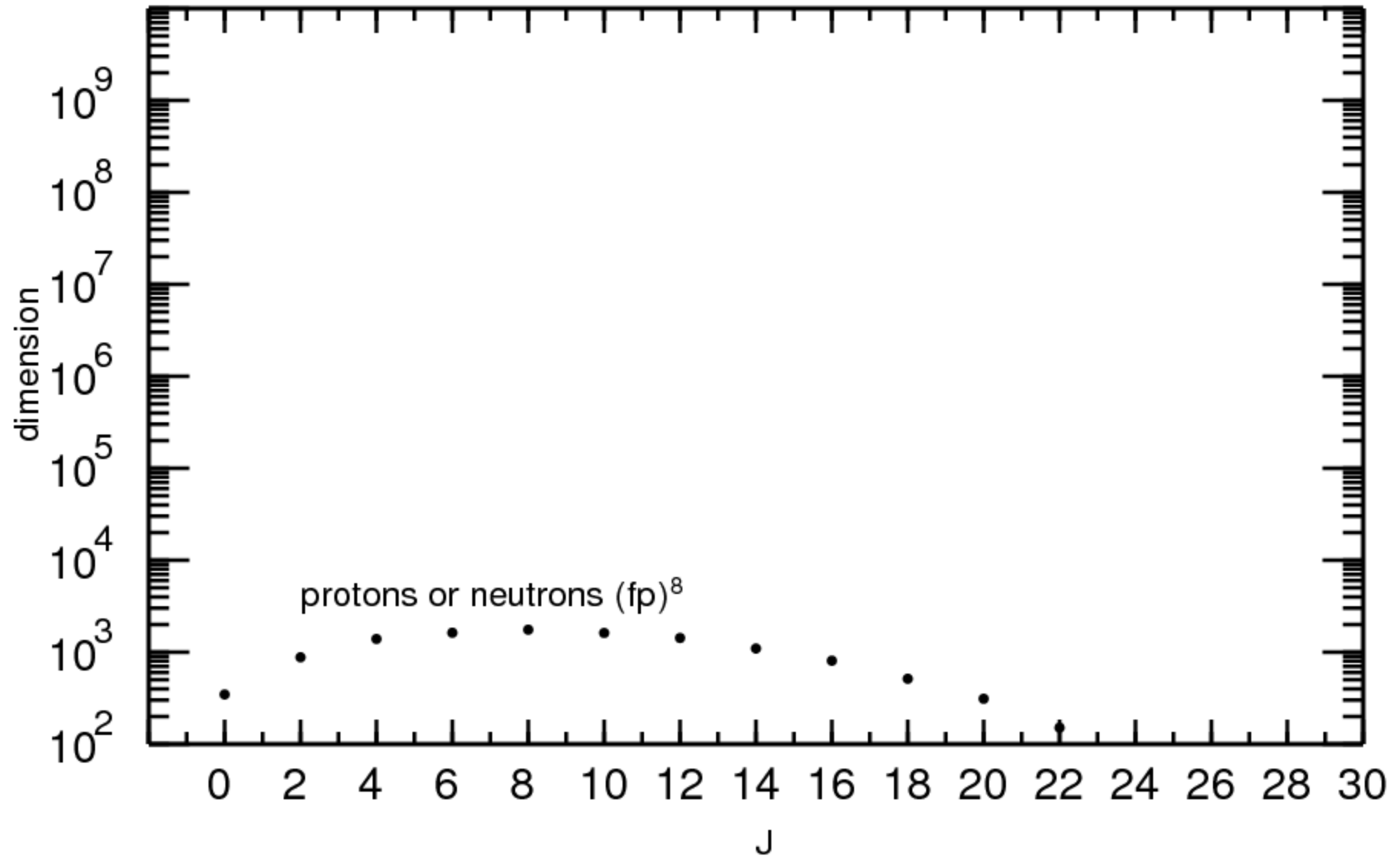
# Positive parity states for $^{26}\text{O}$



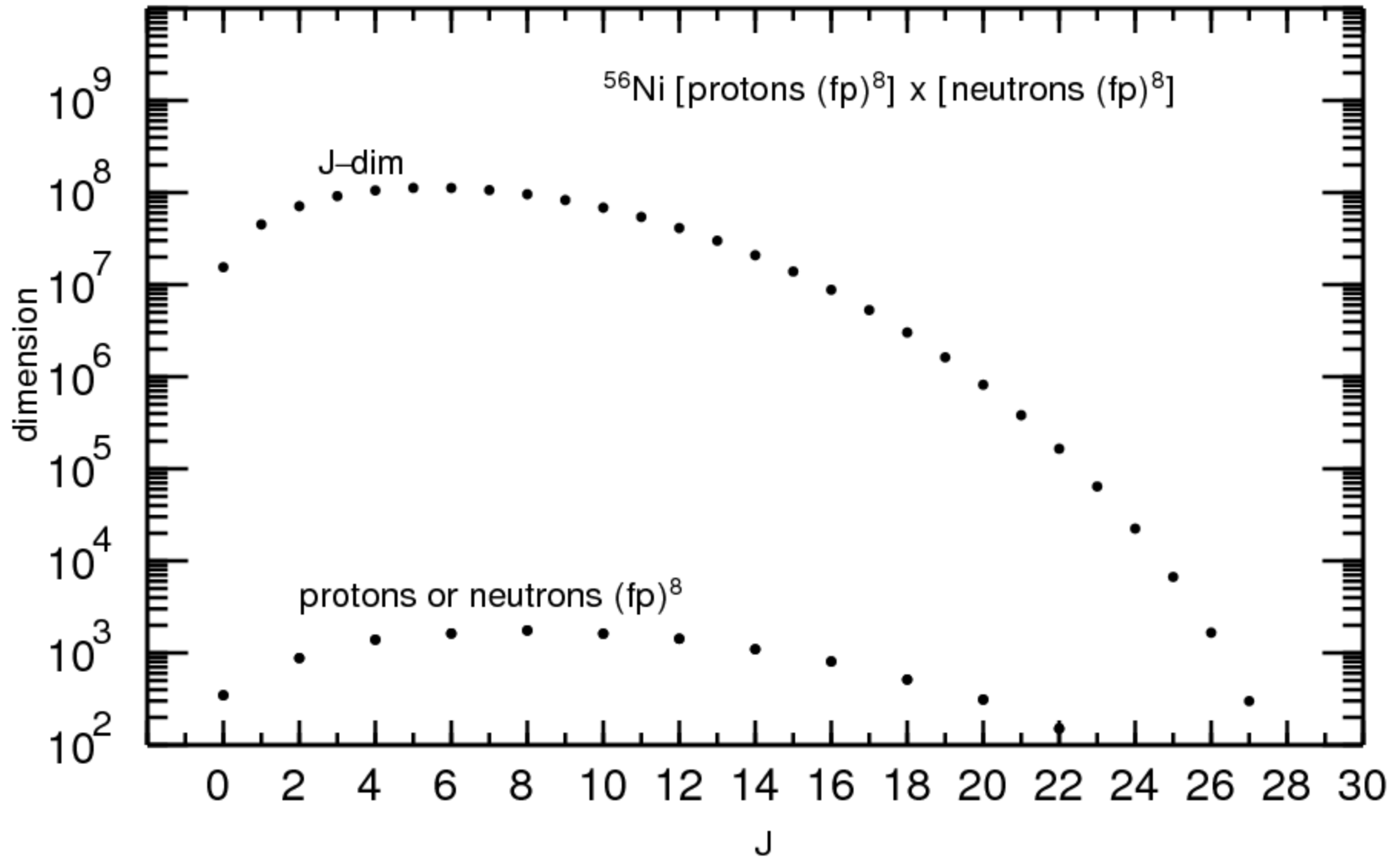
## How to count basis dimensions

- Protons and neutrons – all of those allowed by the triangle conditions  $[(J_p)] \times [(J_n)] \supset [J_{pn}]$   $D_{pn} = D_p D_n$
- Number of states for a given M-value – the sum of the J dimensions from  $J_{\max}$  down to  $J = M$
- J-scheme – basis has good J (or JT)
- M-scheme – basis does not have good J – only M is fixed but the eigenstates will have good J since H is rotationally invariant.

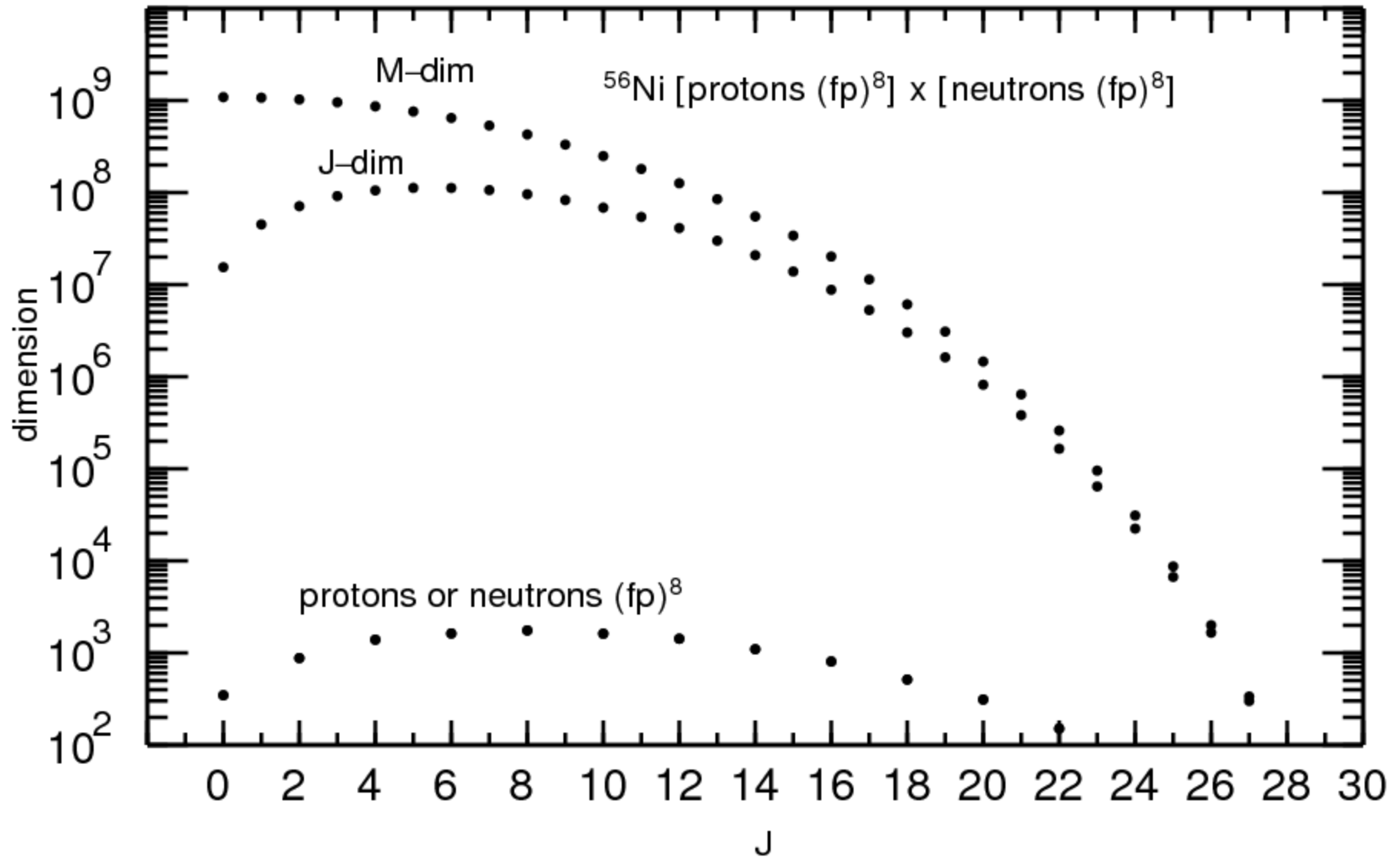
# Example for $^{48}\text{Ca}$ in the pf shell



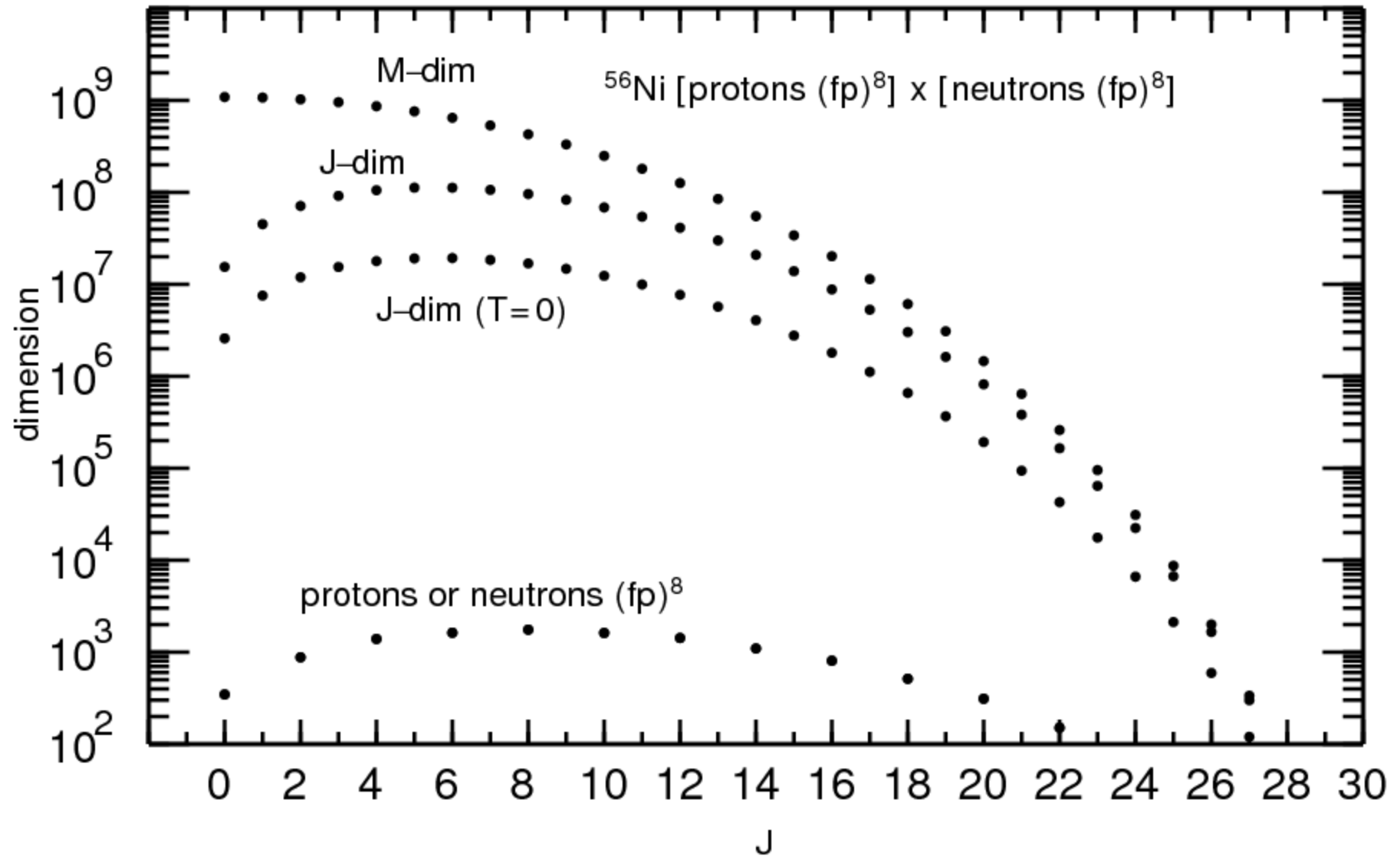
# Example for $^{56}\text{Ni}$ in the pf shell



# Example for $^{56}\text{Ni}$ in the pf shell



# Example for $^{56}\text{Ni}$ in the pf shell



# Codes

- M-scheme (matrix not stored) ( $\sim 10^{10}$  M-states)
  - Antoine (Caurier)
  - Redstick (Ormand and Johnson)
  - CMUShell (Horoi)
  - Mshell (Mizuzaki)
  - MFDn (Vary et al)
- JT-projected M-scheme (matrix stored) ( $\sim 10^5$  JT-states)
  - Oxbash (Brown and Rae) (now replaced by NuShell@MSU)
  - NuShell (Rae)
  - NuShell@MSU (Brown and Rae) (NuShell with Oxbash style input and output)
- J-scheme (matrix not stored) ( $\sim 10^8$  J-states)
  - Nathan (Caurier)
  - EICODE (Toivanen)
  - NuShellx (Rae)
  - NuShellx@MSU (Brown and Rae) (NuShellx with Oxbash style input and output)

available on the web




Bill Rae (Garsington) has made big advances

Oxbash -> Nushell -> Nushellx

Nushellx uses the [ Jp Jn ] J coupling to eliminate storage of the matrix.

Similar to Nathan (Caurier et al) and Eicode (Toivanen) but faster.

NuShellx@MSU uses these codes as a core for nuclear structure applications.



**Nuclear Shell Model Codes**  
Home of NuShell, NuShellX and SunShell

[Home](#)  
[About Us](#)  
[Download](#)

Home Page of NuShell, NuShellX and SunShell.

NuShell is possibly one of the easiest shell model codes to use!

---

**NuShellX - NuShell's Big Brother**

**The faster, easy choice for large scale shell model calculations !**

Effective interactions – what are the two-body matrix elements?

For sd shell USD, USDA, USDB interactions obtained from a fit to data  
use singular-value-decomposition method to obtain values  
for 20-30 of the most important linear combinations of TBME  
from about 600 energies

For pf same procedure for about 600 energies

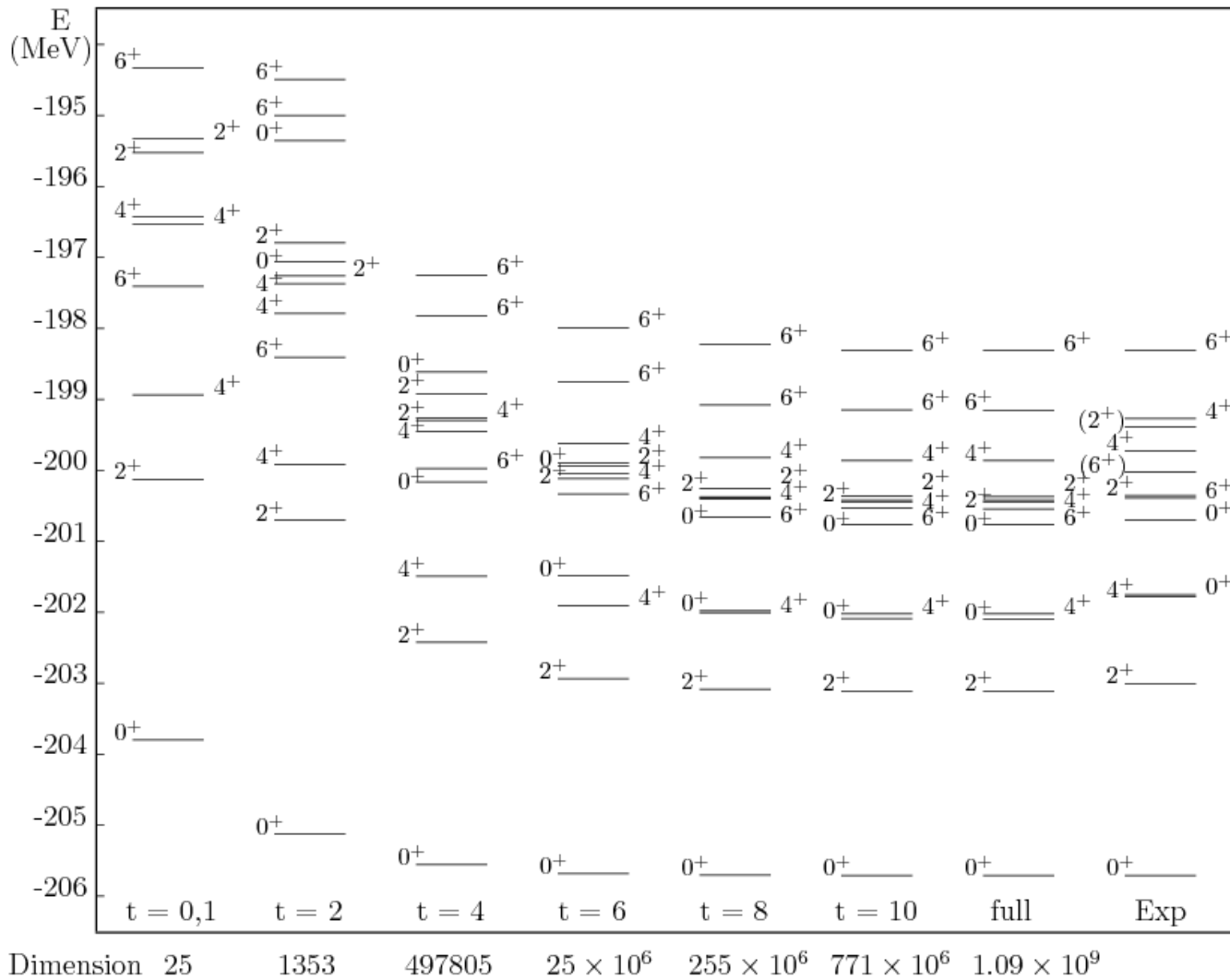
M. Honma, T. Otsuka, B. A. Brown and T. Mizusaki,  
Phys. Rev. C65, 061301 (2002) - GPFX1, GPFX1A

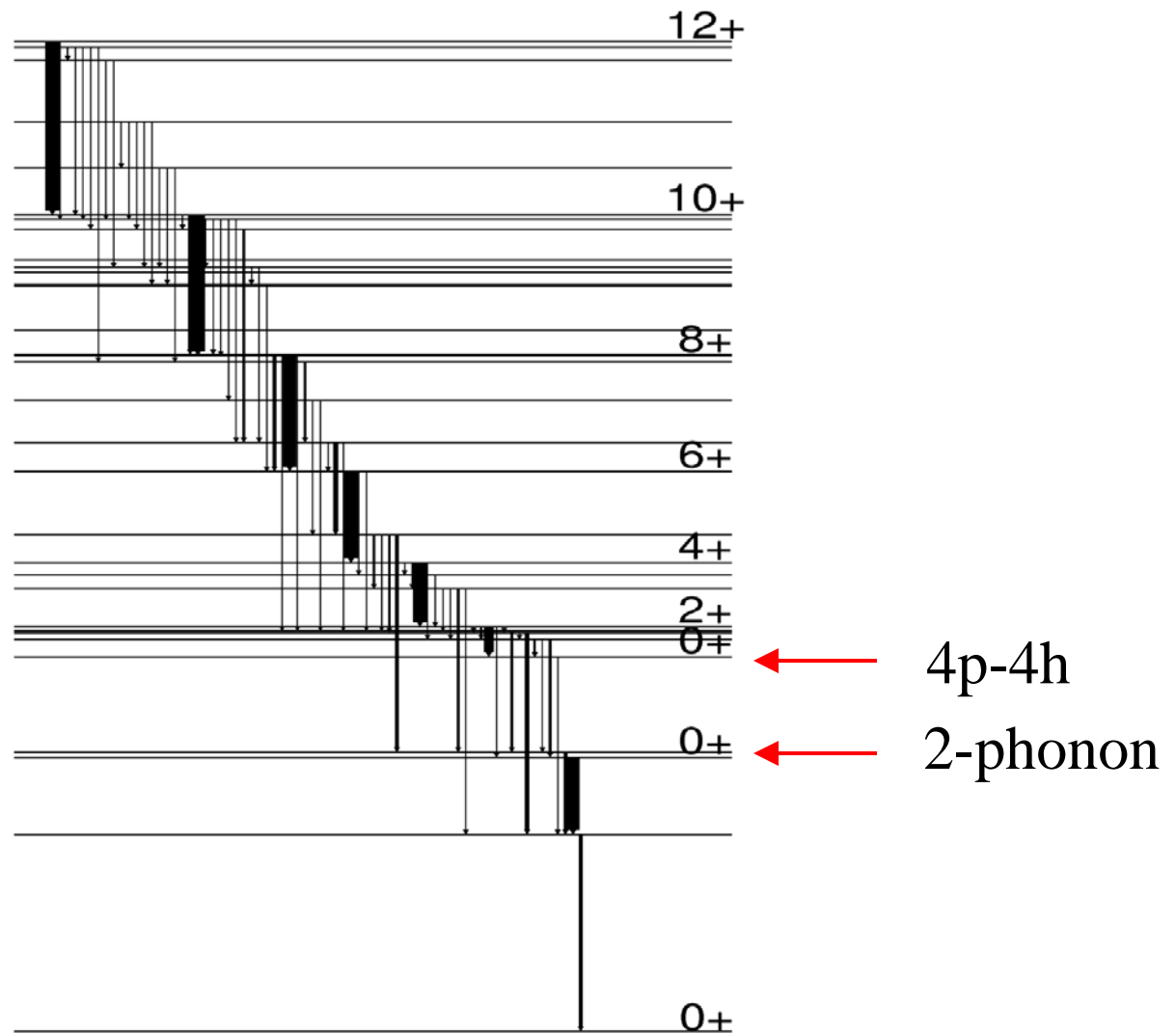
For heavier nuclei this method becomes unfeasible – we need better  
methods for understanding the nuclear medium and model space  
dependence of the NN and NNN interactions  
(tomorrow) talk at Stony Brook on  $^{132}\text{Sn}$  and  $^{208}\text{Pb}$  regions



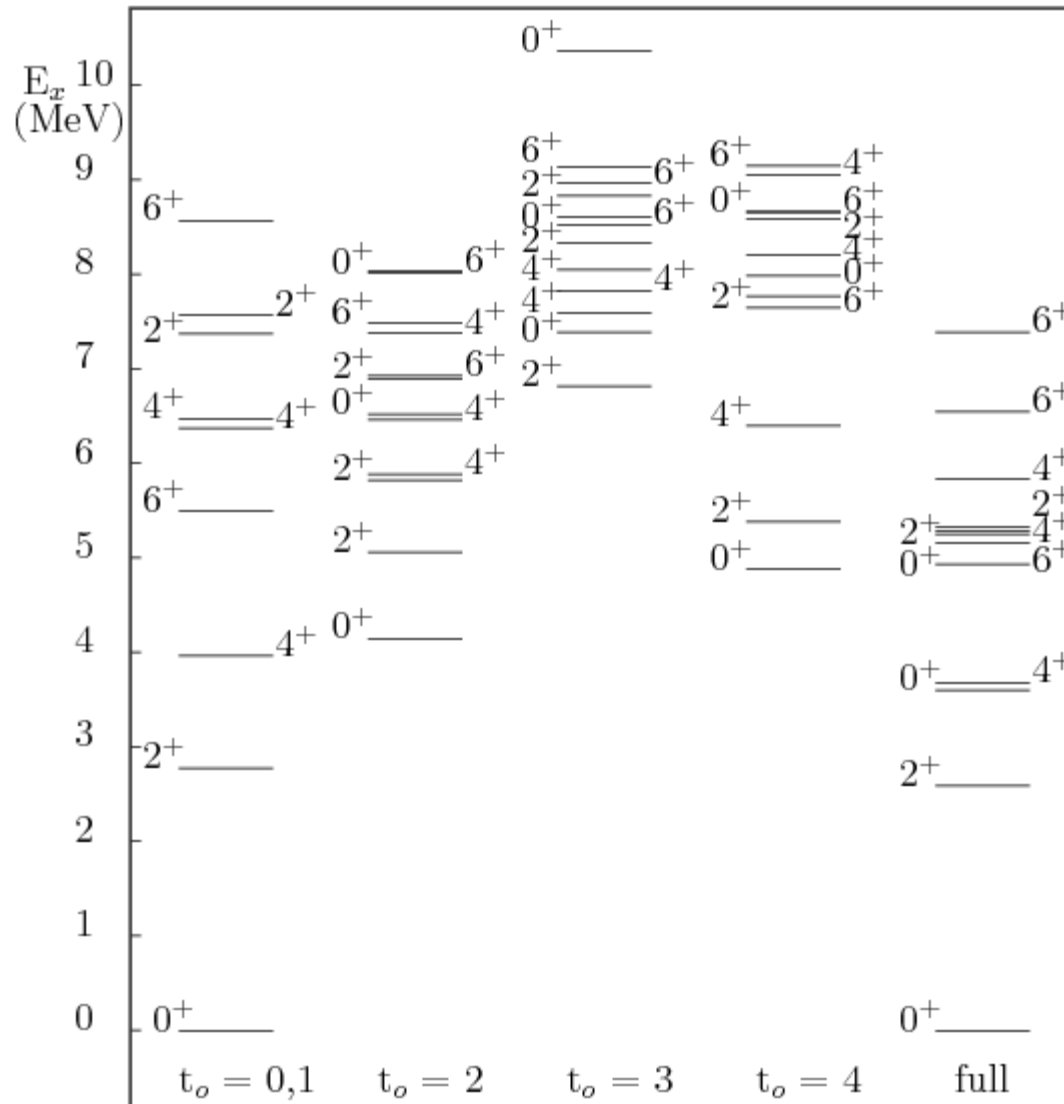
# Full pf space for $^{56}\text{Ni}$ with GXPF1A Hamiltonian (order of one day computing time)

M. Horoi, B. A. Brown, T. Otsuka, M. Honma and T. Mizusaki, Phys. Rev. C 73, 061305(R) (2006).





# Pure configurations



Requires an effective shell gap 0.9 MeV smaller than full fp



PHYSICAL REVIEW C

VOLUME 10, NUMBER 5

NOVEMBER 1974

## Levels of ${}^{56}\text{Ni}^\dagger$

H. Nann\* and W. Benenson

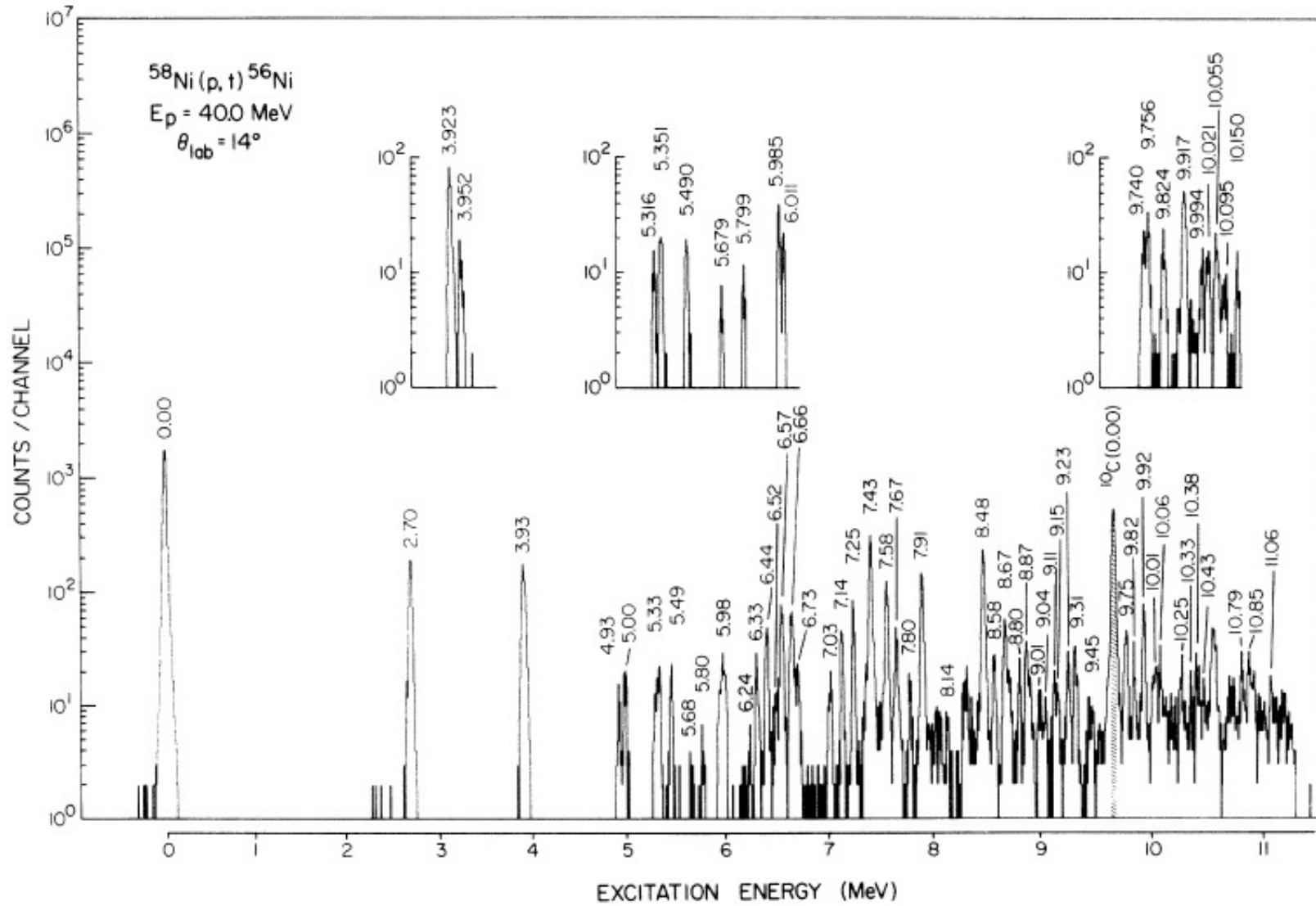
*Cyclotron Laboratory and Department of Physics, Michigan State University, East Lansing, Michigan 48824*

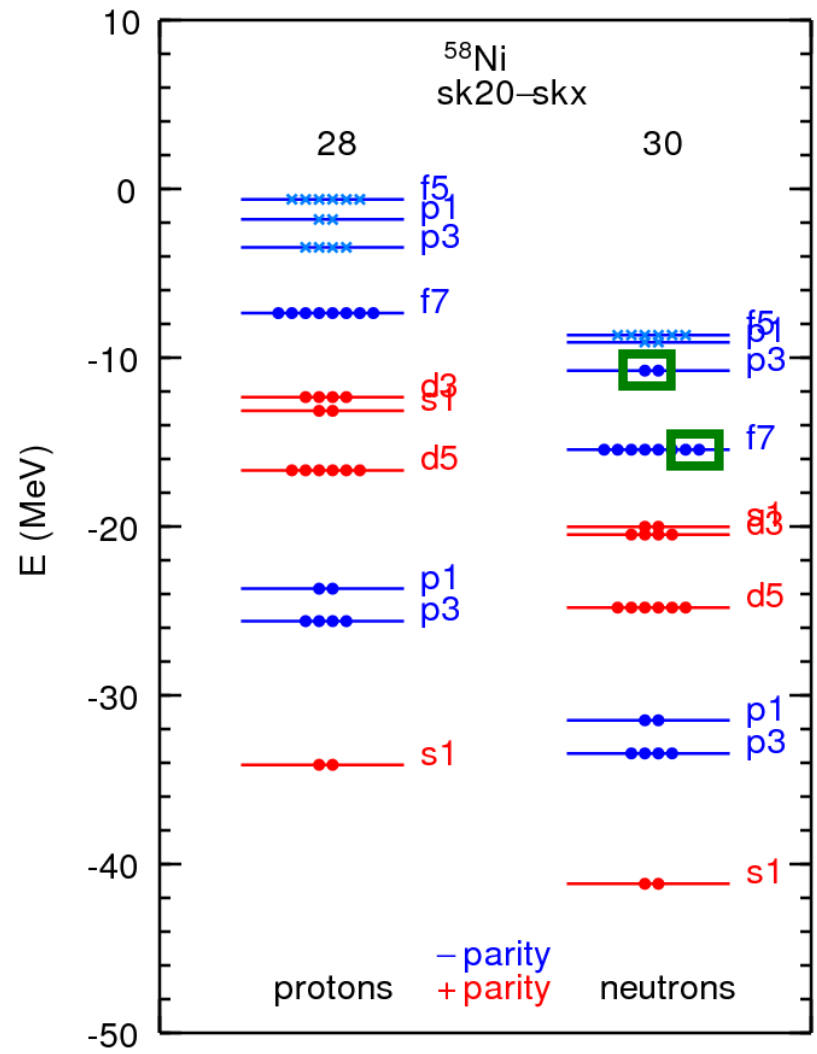
(Received 5 August 1974)

The  ${}^{58}\text{Ni}(p,t){}^{56}\text{Ni}$  reaction was studied at 40 and 45 MeV beam energy. An energy resolution of 10–25 keV permitted observation of 60 levels with excitation energy up to 10.5 MeV. Spin and parity are assigned to levels which were excited with characteristic angular distributions. These include  $0^+$  states at 3.95, 5.00, 6.44, 7.91, 9.92, 9.99, and 10.02 MeV.



# 60 levels 10 keV resolution



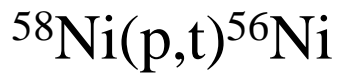




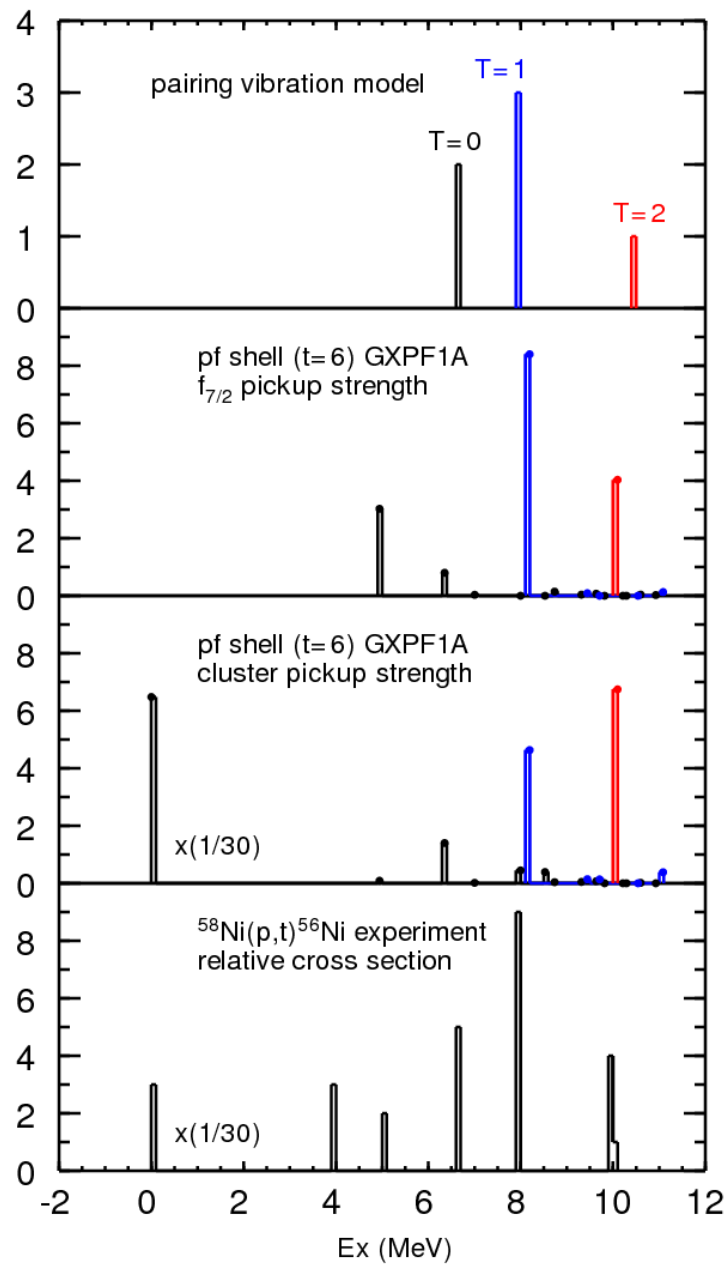


Pairing vibrations expect three  $0^+$  levels with  $T=0,1,2$  strength 2:3:1 and spacing that goes as  $T(T+1)$

- <sup>13</sup>A. Bohr, in *International Symposium on Nuclear Structure, Dubna, 1968* (IAEA, Vienna, 1968), p. 179.
- <sup>14</sup>O. Nathan, in *International Symposium on Nuclear Structure, Dubna, 1968* (see Ref. 13), p. 191.

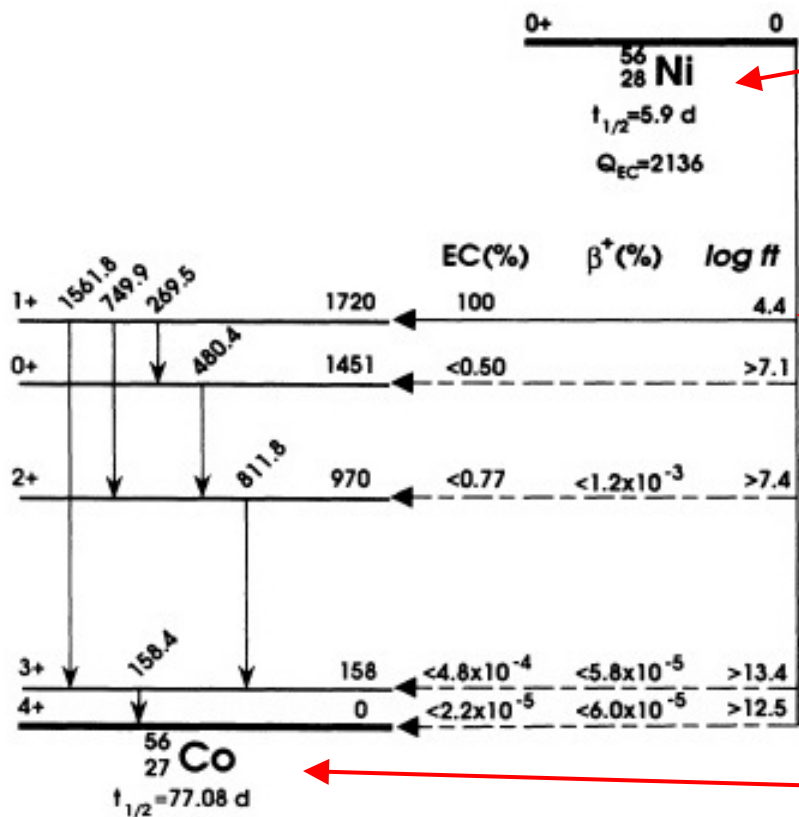


Relative strength  
For  $0^+$  states in  $^{56}\text{Ni}$



### Reinvestigation of <sup>56</sup>Ni decay

Bhaskar Sur, Eric B. Norman, K. T. Lesko, Edgardo Browne, and Ruth-Mary Larimer  
 Nuclear Science Division, Lawrence Berkeley Laboratory, 1 Cyclotron Road, Berkeley, California 94720



End product of Si burning in stars.  
 Without electrons half-life of <sup>56</sup>Ni increases from 5.9 days to 10<sup>5</sup> years.

B(GT) = 0.155 compared to (f<sub>7/2</sub>)<sup>-1</sup> (f<sub>5/2</sub>) value of 48x2/7 = 13.71

B(F) < 0.0005

After explosion light output from <sup>56</sup>Co decay with half-life of 77 days.

STELLAR WEAK INTERACTION RATES<sup>1</sup> FOR INTERMEDIATE-MASS NUCLEI. II.  $A=21$  TO  $A=60$

GEORGE M. FULLER<sup>2</sup> AND WILLIAM A. FOWLER  
W. K. Kellogg Radiation Laboratory, California Institute of Technology

AND

MICHAEL J. NEWMAN  
Applied Theoretical Physics Division, Los Alamos National Laboratory, University of California, Los Alamos  
*Received 1981 June 12; accepted 1981 August 3*

THE ASTROPHYSICAL JOURNAL, **252**:715–740, 1982 January 15

© 1982. The American Astronomical Society. All rights reserved. Printed in U.S.A.

---

ELECTRON CAPTURE AND  $\beta$ -DECAY IN PRESUPERNOVA STARS

M. B. AUFDERHEIDE

Physics Department, State University of New York at Stony Brook; and Department of Physics, Brookhaven National Laboratory

G. E. BROWN, T. T. S. KUO, and D. B. STOUT

Physics Department, State University of New York at Stony Brook

AND

P. VOGEL

Physics Department, California Institute of Technology

*Received 1989 September 25; accepted 1990 April 2*

THE ASTROPHYSICAL JOURNAL, **362**:241–250, 1990

---

RATE TABLES FOR THE WEAK PROCESSES OF  $pf$ -SHELL NUCLEI  
IN STELLAR ENVIRONMENTS

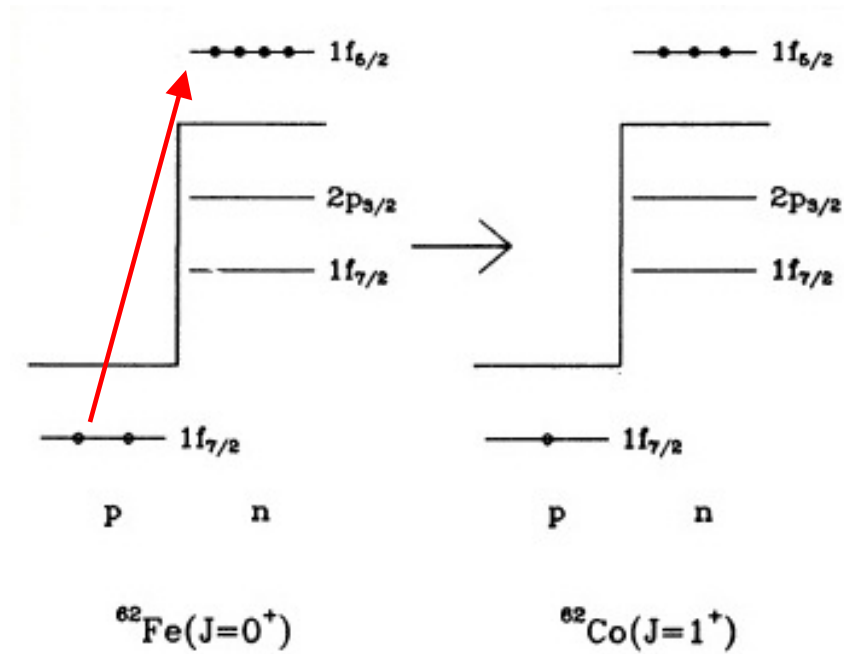
K. LANGANKE<sup>1</sup> and G. MARTÍNEZ-PINEDO<sup>2</sup>

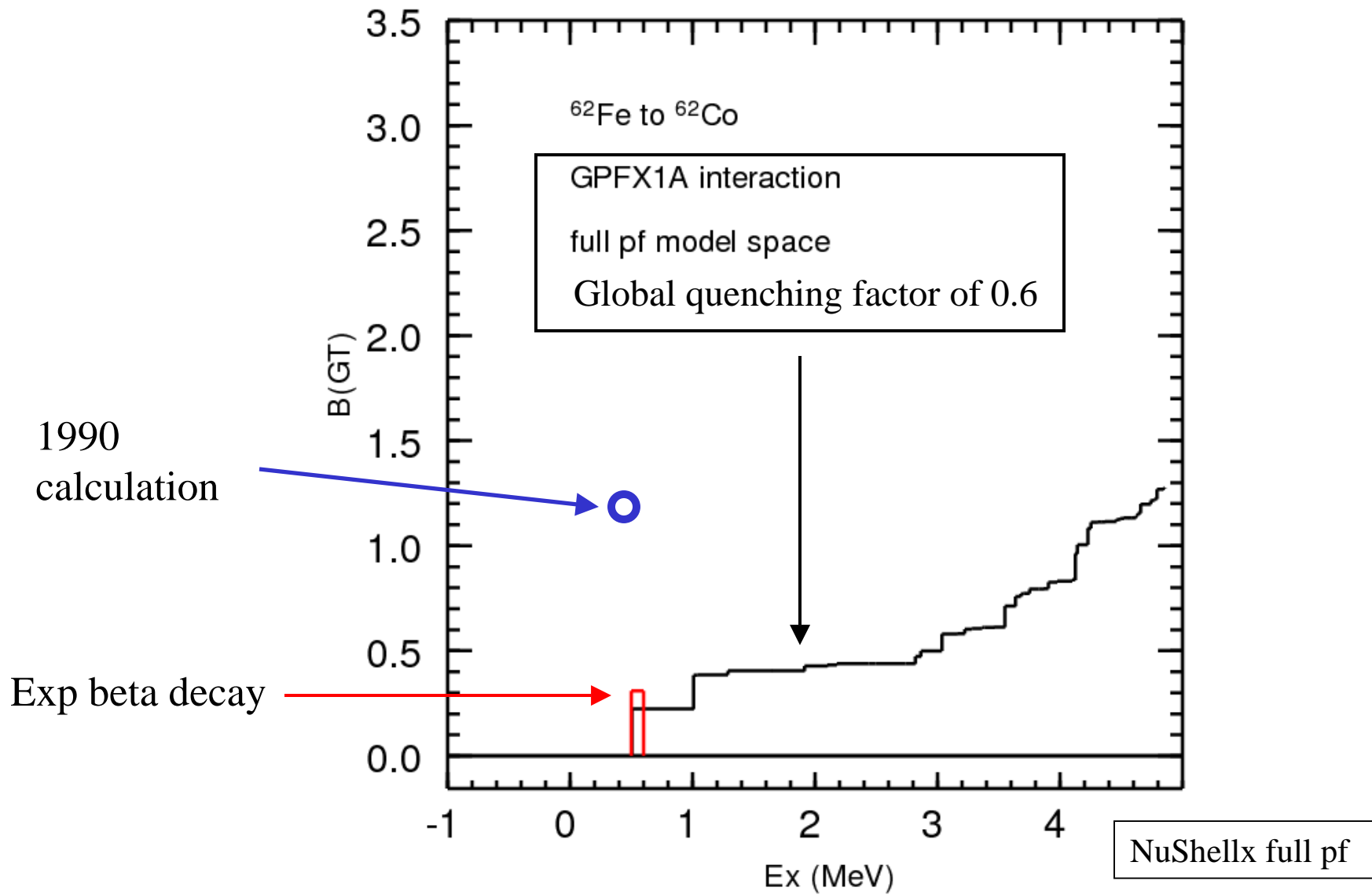
Institut for Fysik og Astronomi, Århus Universitet, DK-8000 Århus C, Denmark

Atomic Data and Nuclear Data Tables **79**, 1–46 (2001)

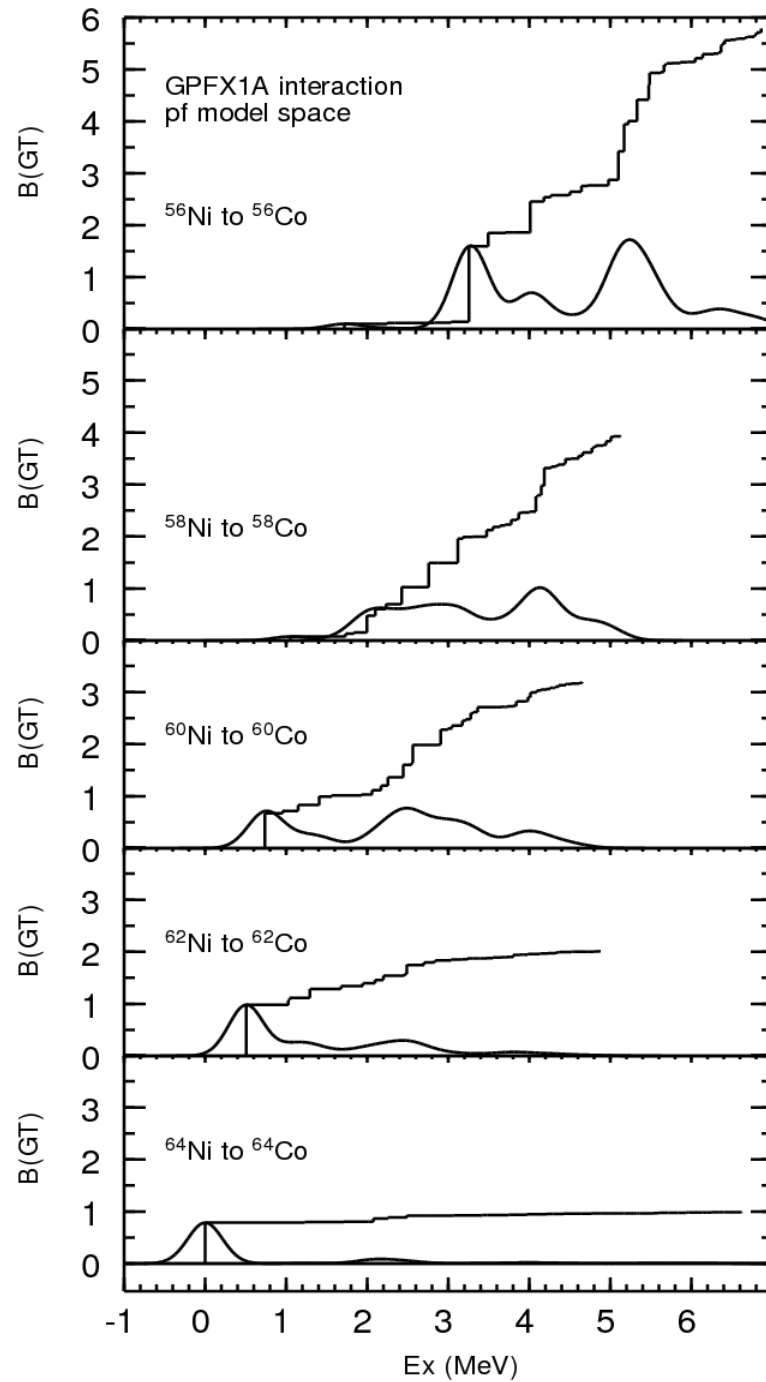


# $^{62}\text{Fe}$ to $^{62}\text{Co}$ model from Aufderheide et al. 1990





# B(GT) for electron capture for Ni to Co

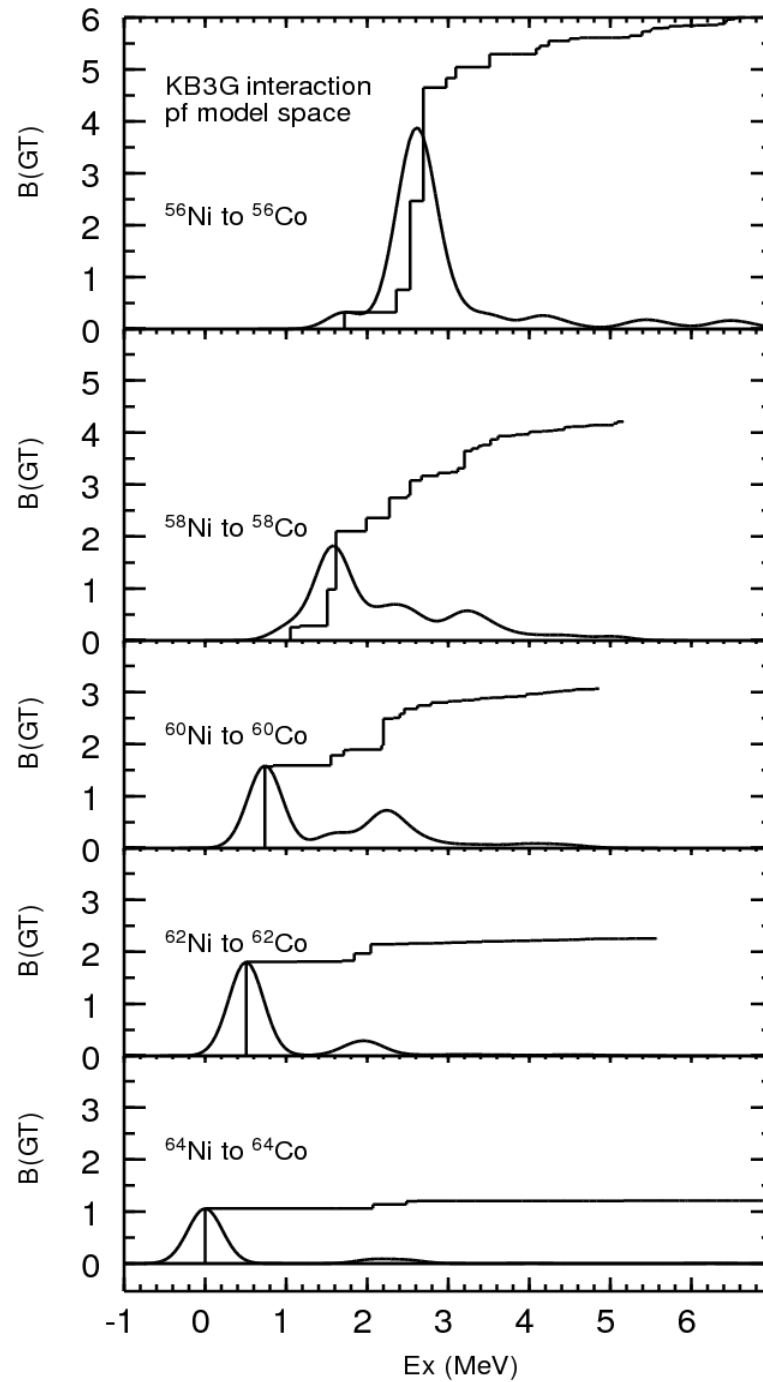


NuShellx t=6

7, USNDP, BNL, Nov 6, 2008



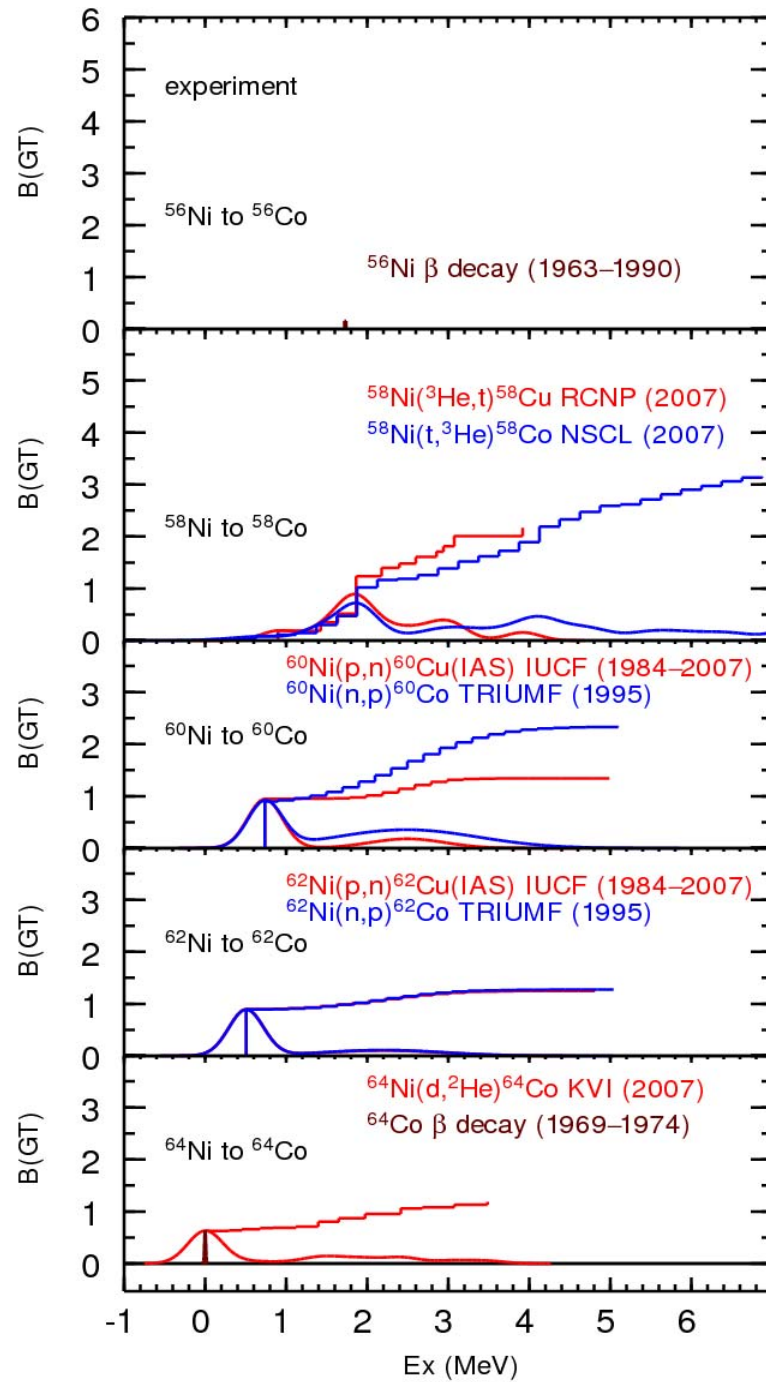
# B(GT) for electron capture for Ni to Co

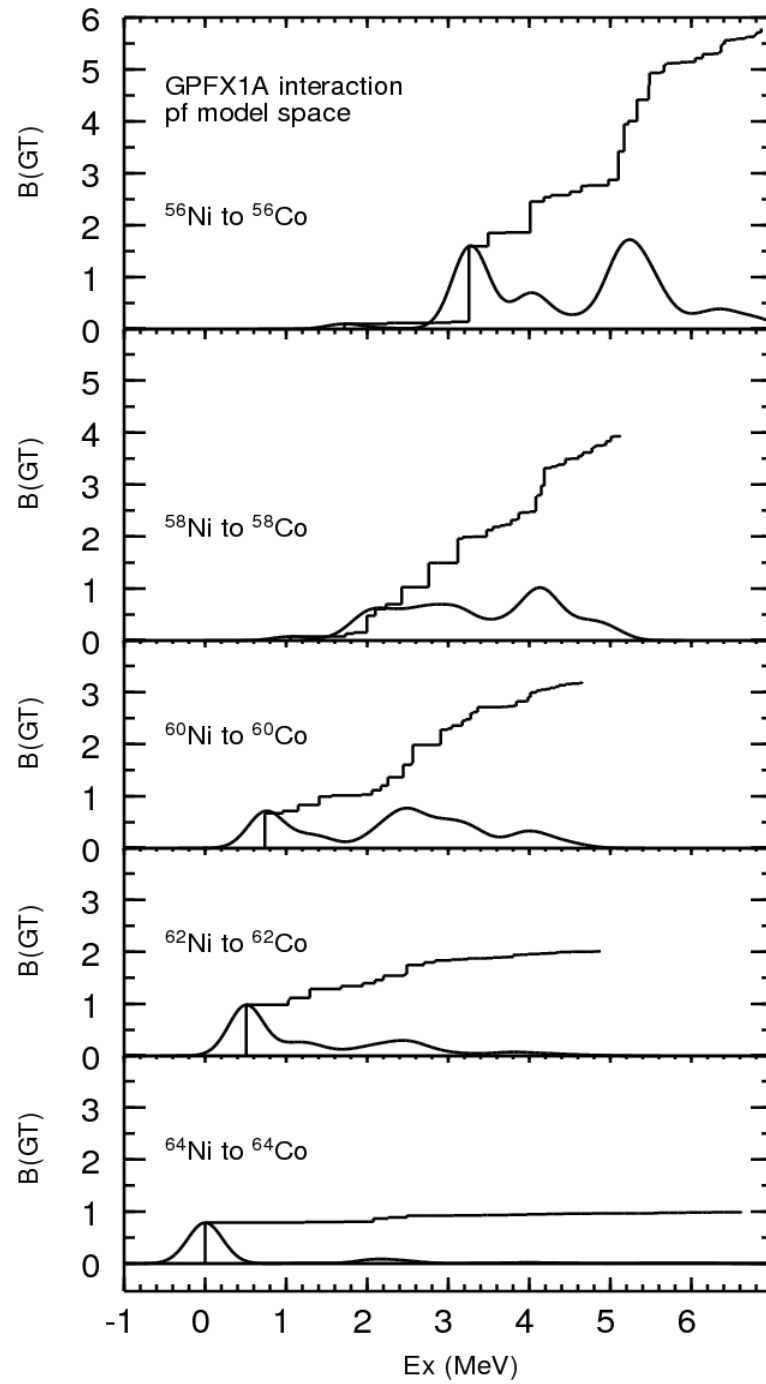


NuShellx t=6

in, USNDP, BNL, Nov 6, 2008



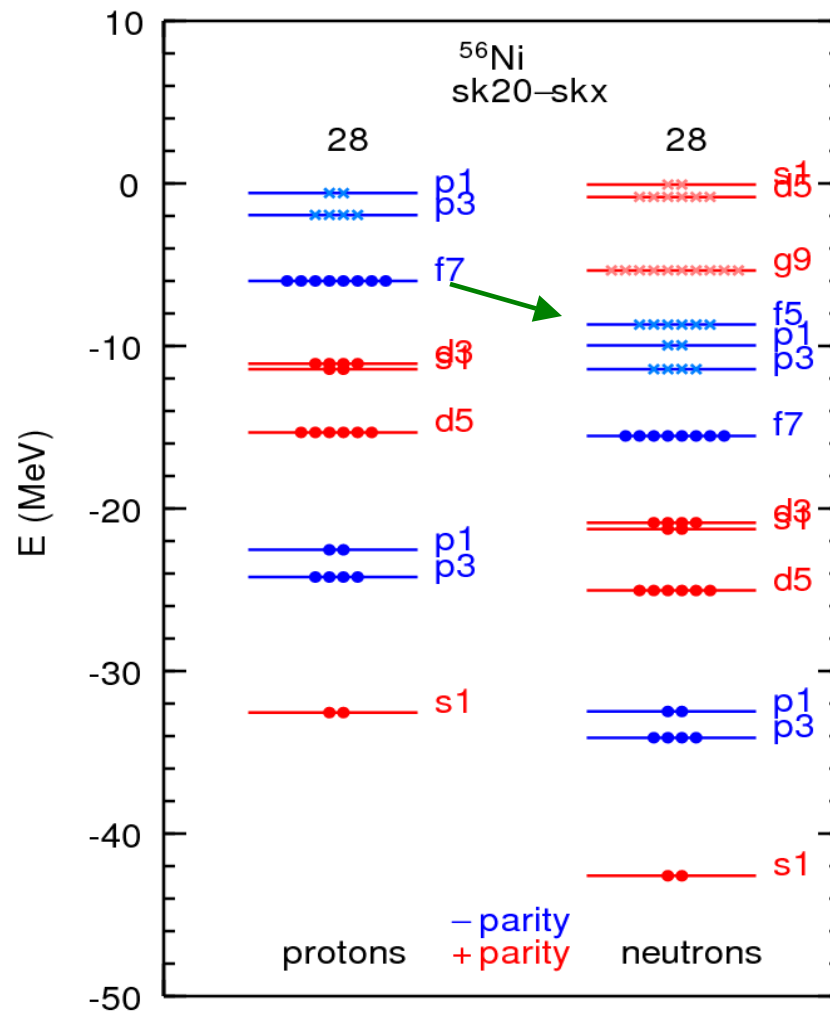


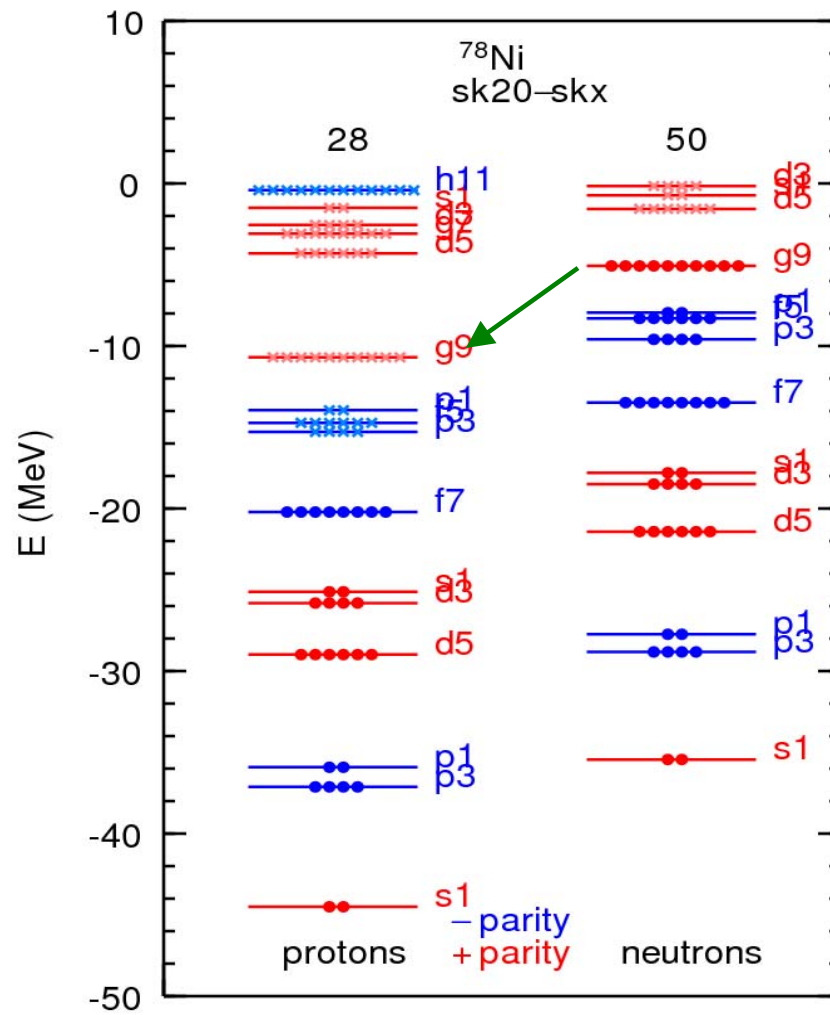


NuShellx t=6

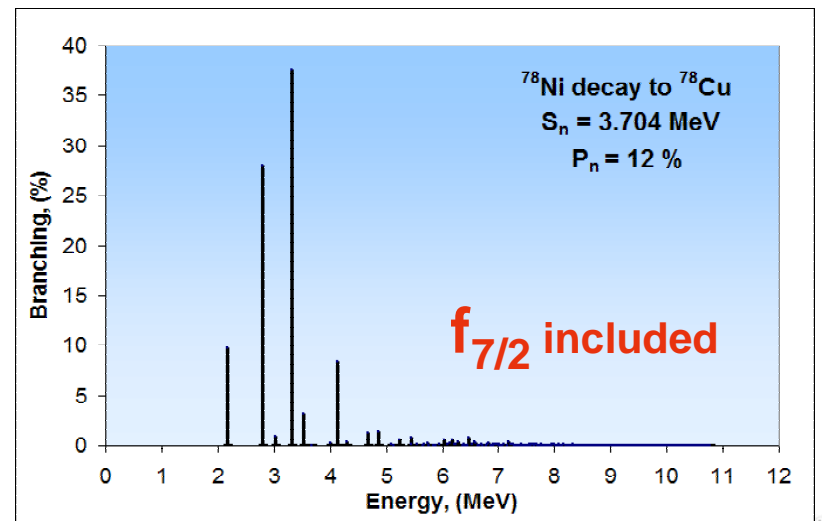
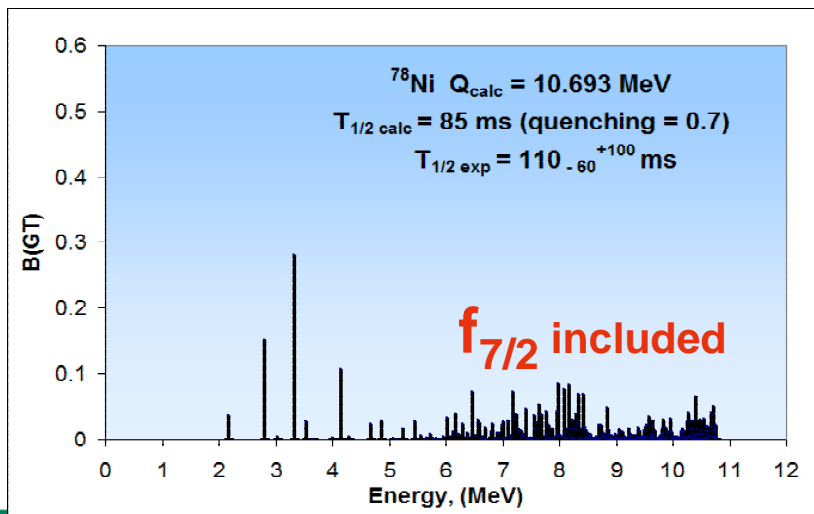
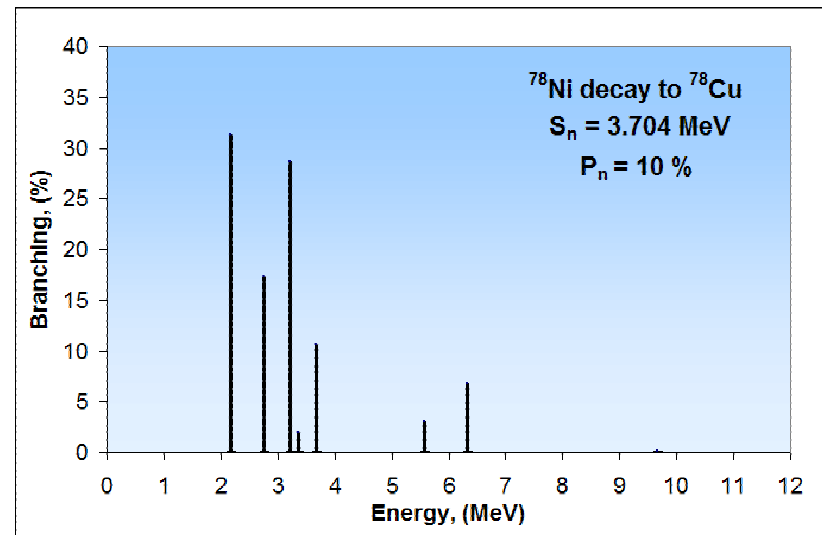
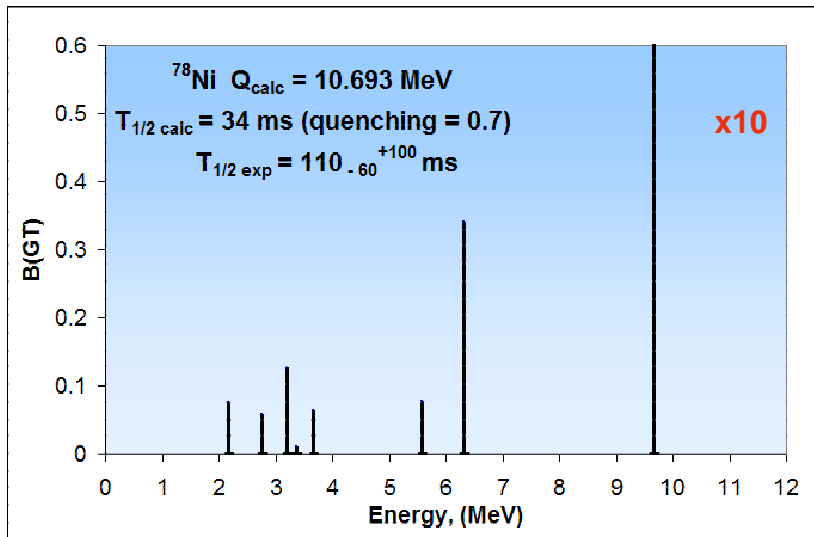
7, USNDP, BNL, Nov 6, 2008



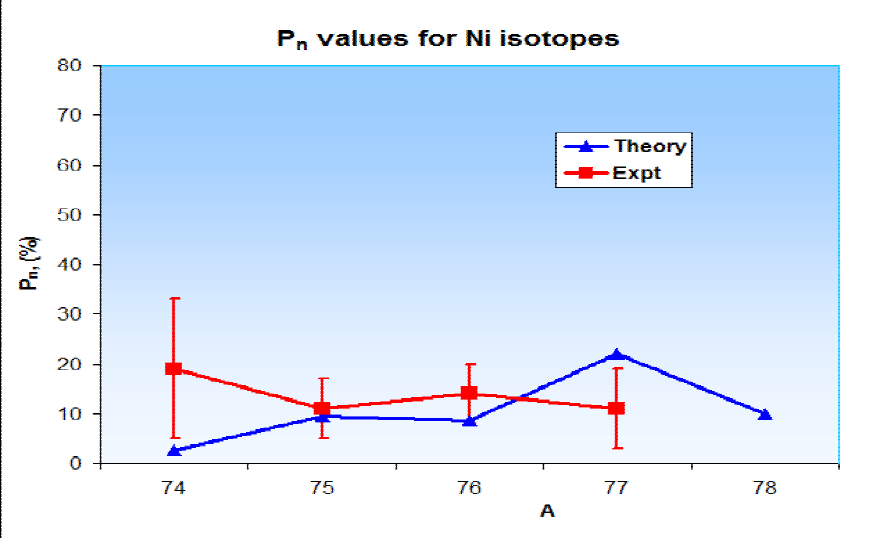
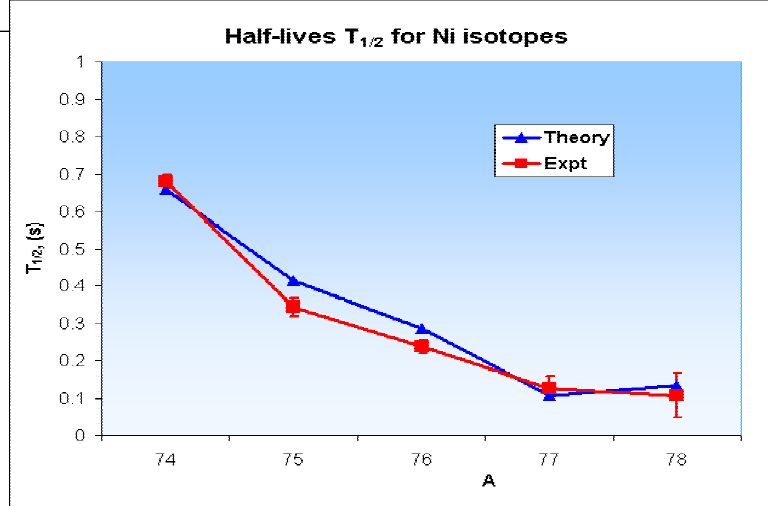
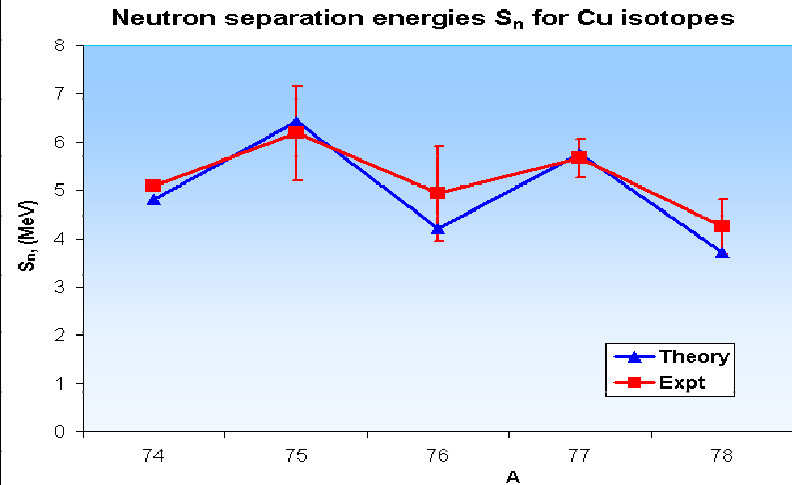
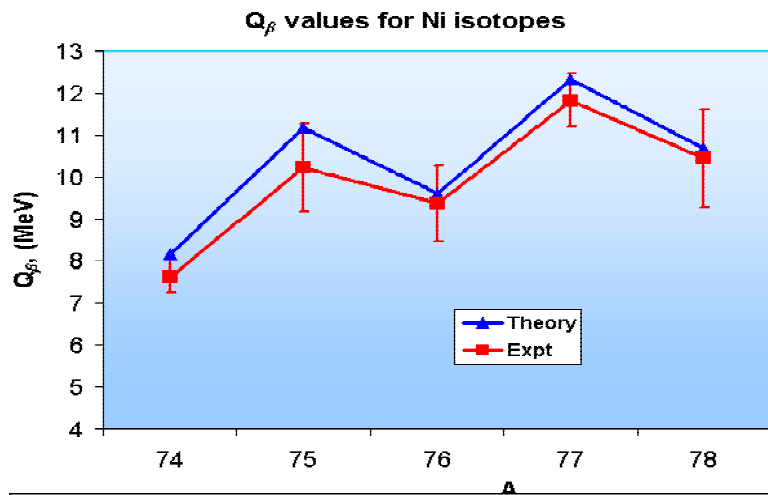




# $^{78}\text{Ni}$ : beta-decay (Lisetsky)

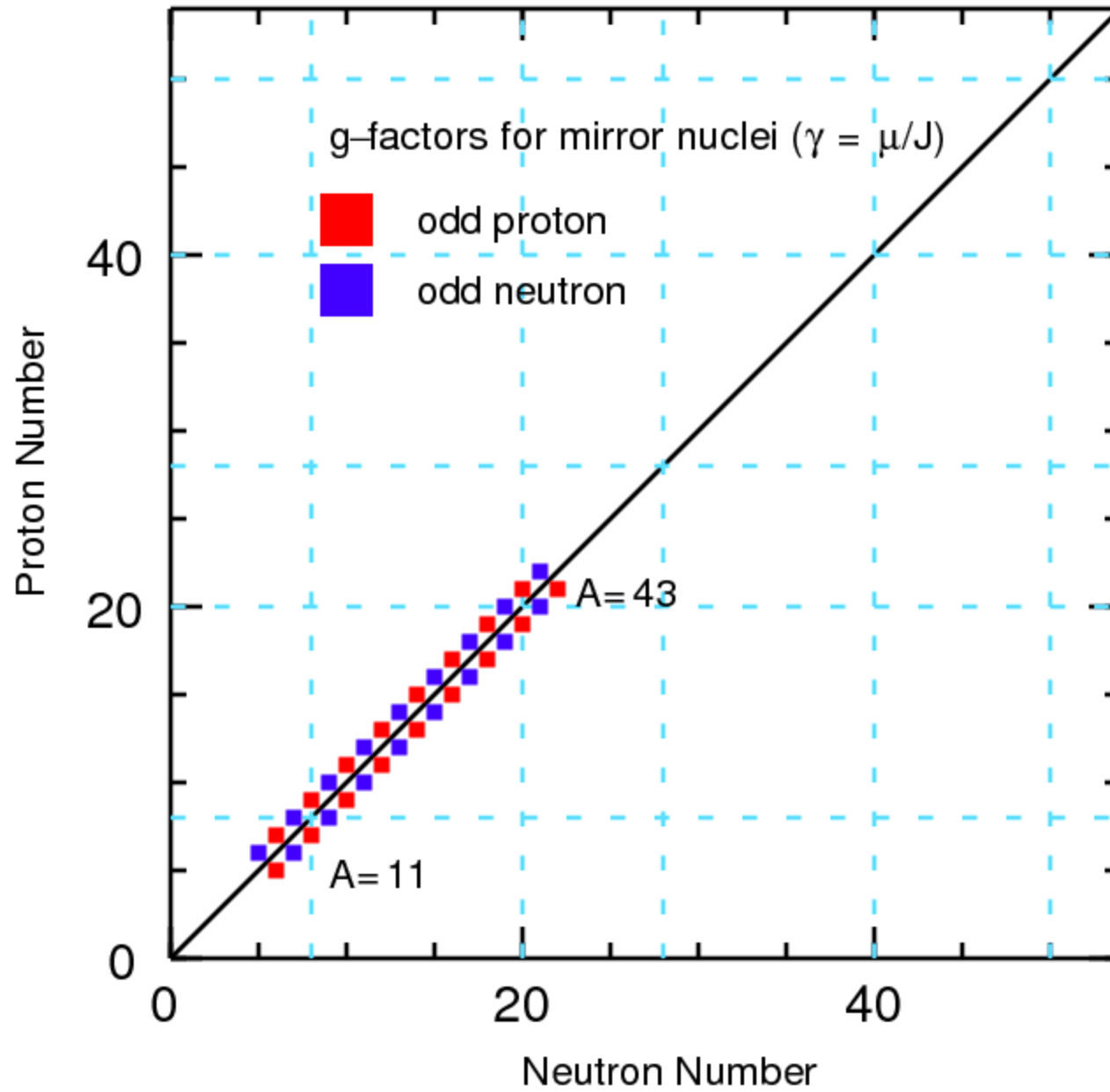


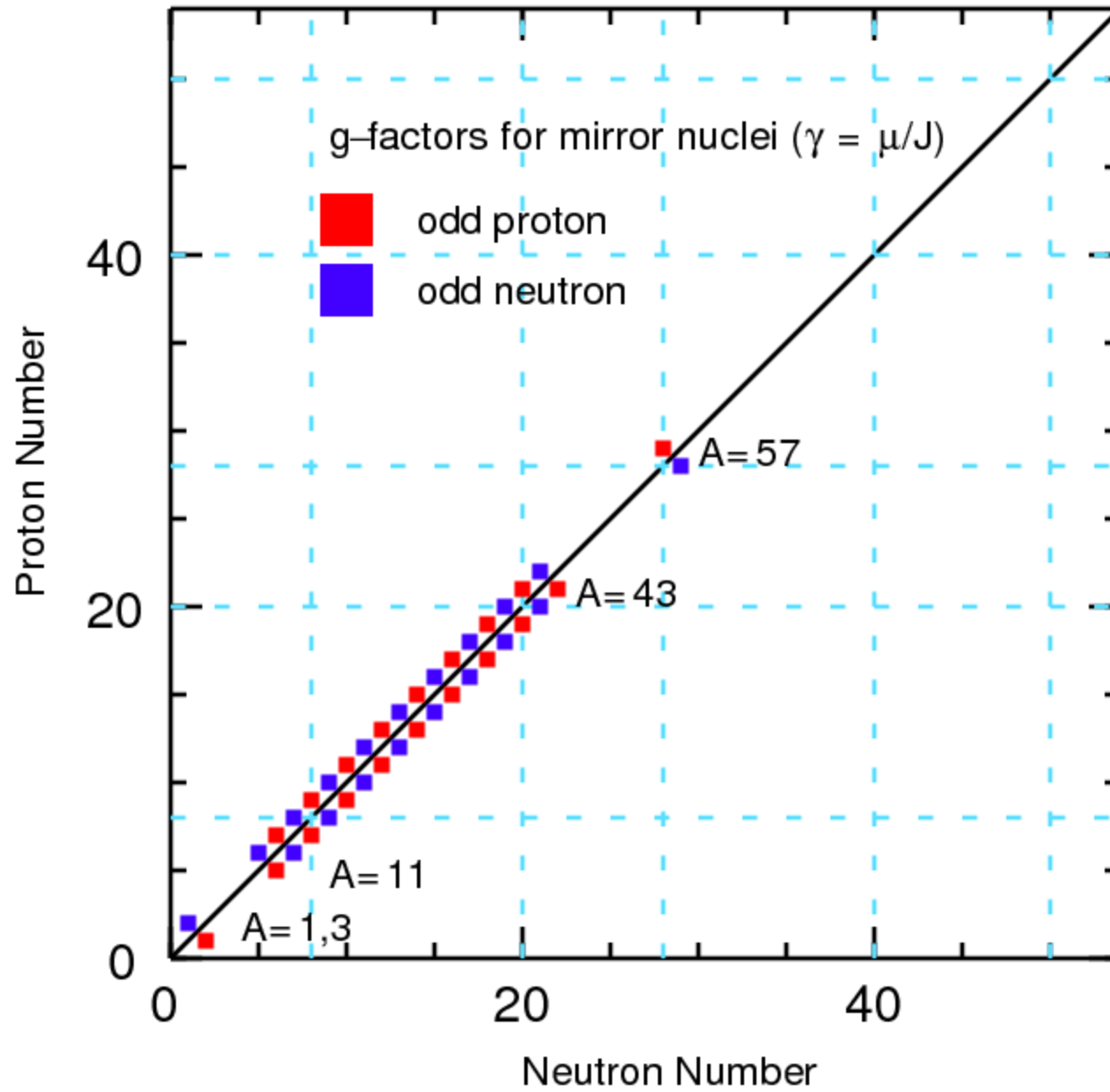
# Ni: Beta-decay results (Lisetsky)



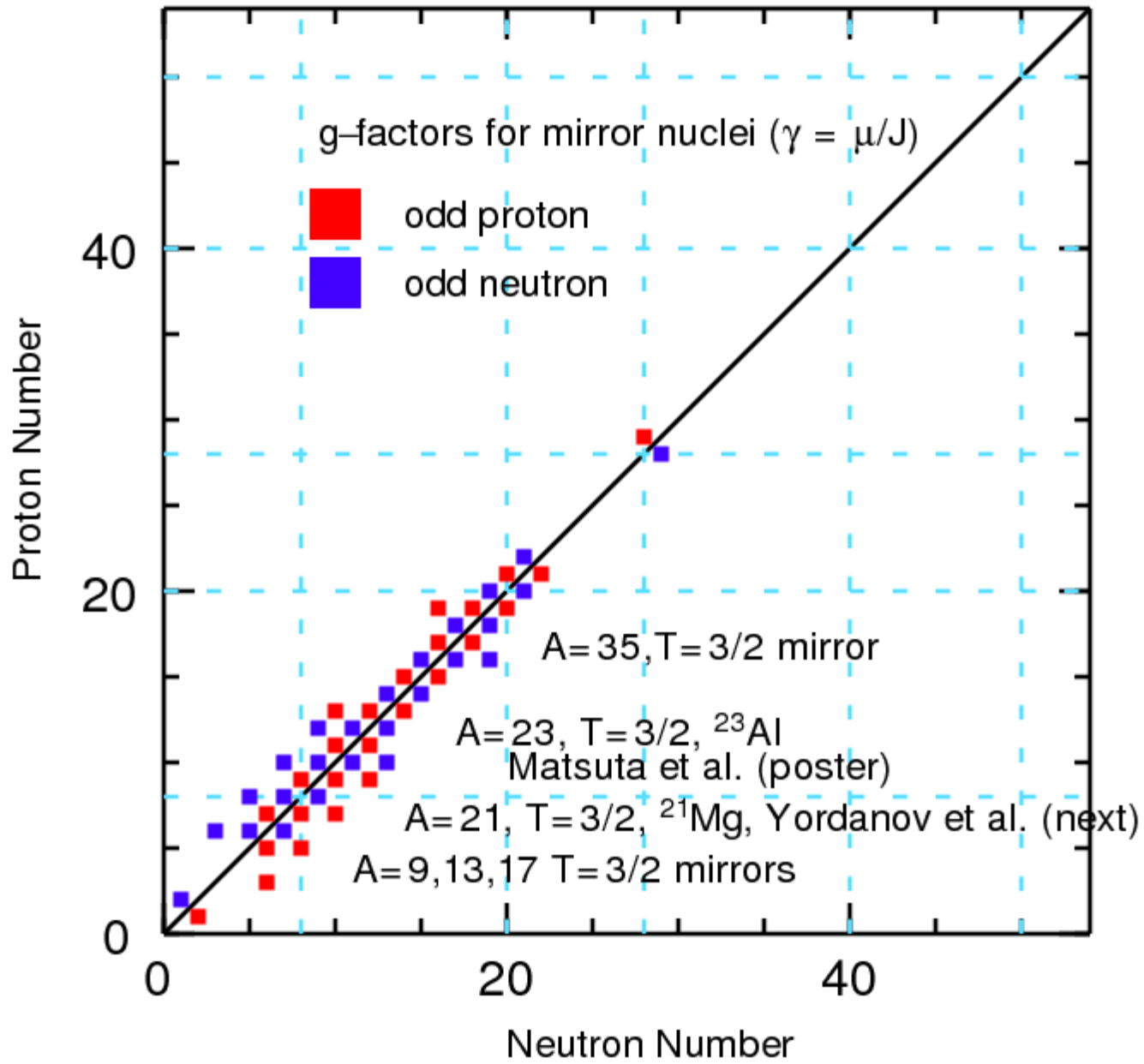
P. T. Hosmer et al., PRL 94, 112501 (2005)

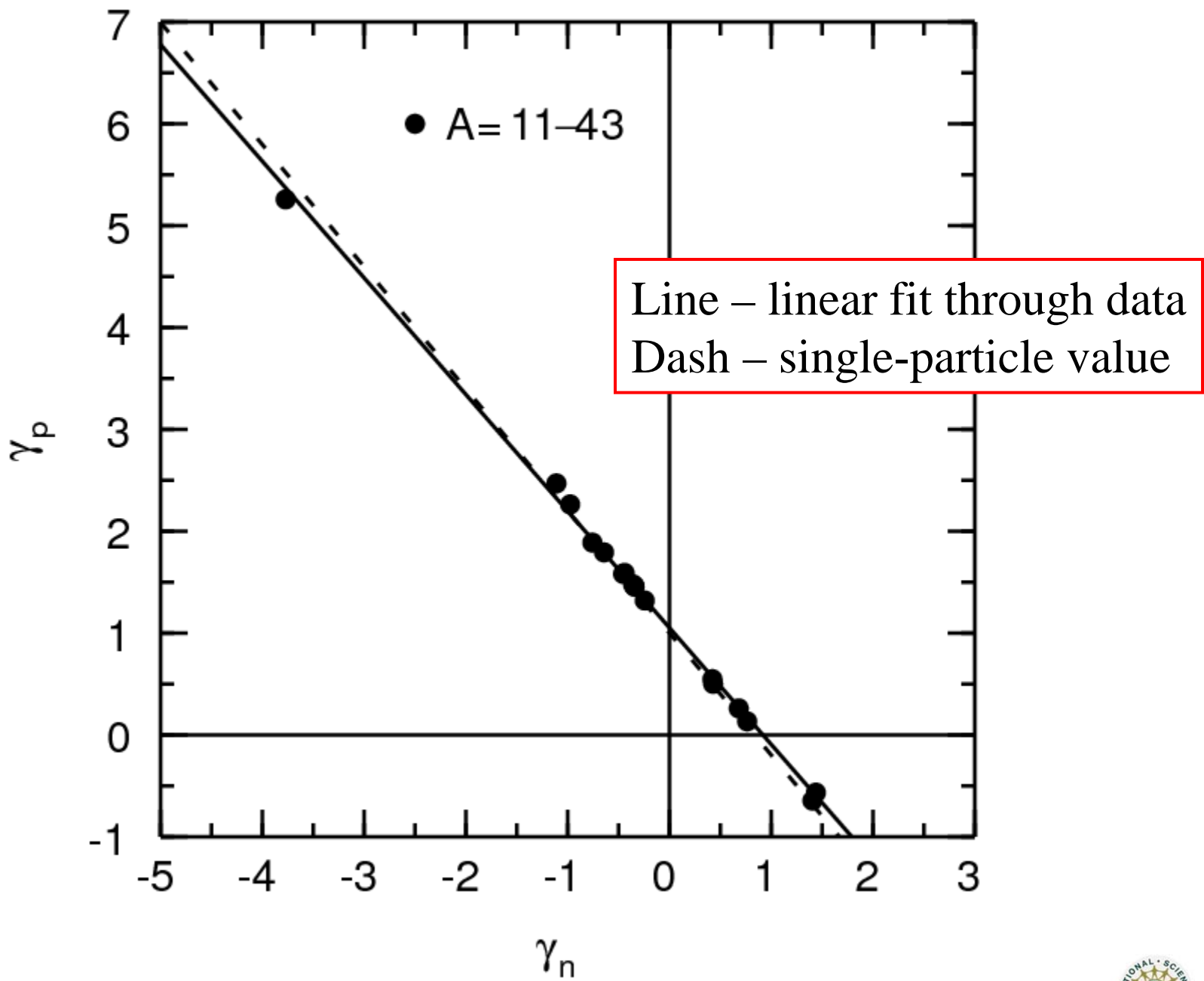


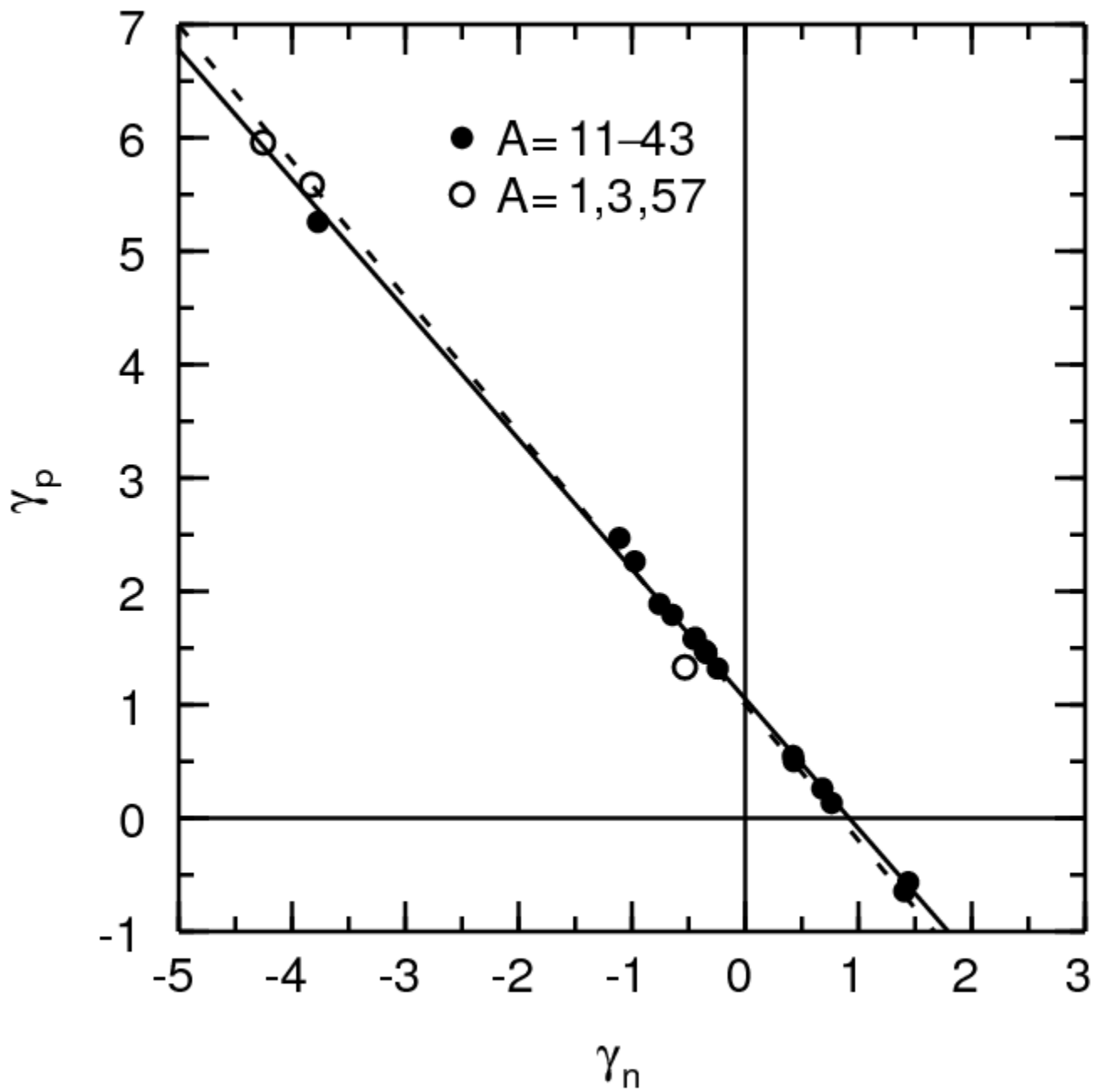


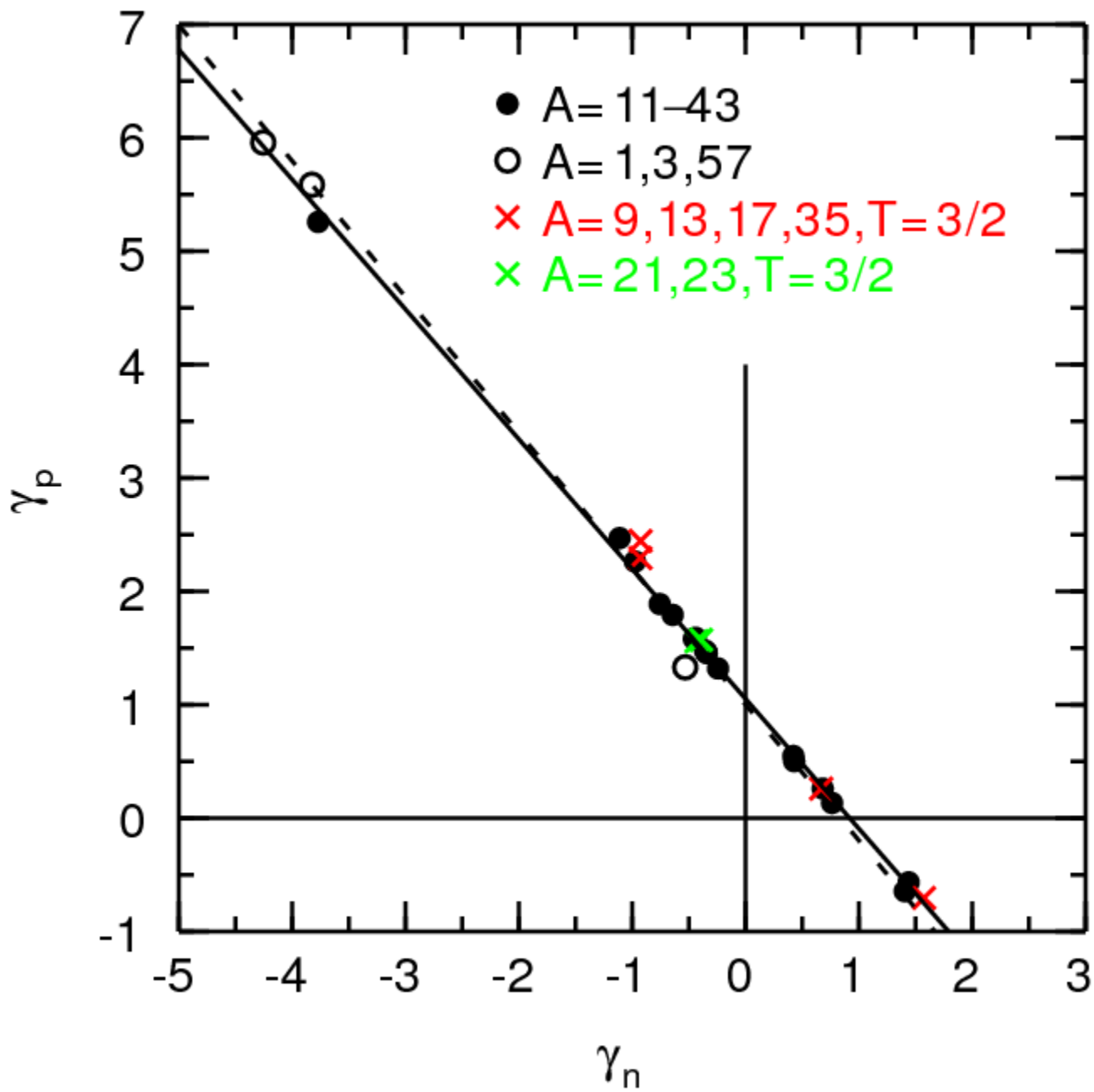












Phys. Rev. C 77, 064311 (2008)

## Correlations between magnetic moments and beta decay

S. M. Perez<sup>1,2</sup>, W. A. Richter<sup>3</sup>, B. A. Brown<sup>4</sup> and M. Horoi<sup>5</sup>

<sup>1</sup> *Department of Physics, University of Cape Town, Private Bag, Rondebosch 7700, South Africa*

<sup>2</sup> *iThemba LABS, P. O. Box 722, Somerset West 7129, South Africa*

<sup>3</sup> *Department of Physics, University of the Western Cape, Private Bag X17, Bellville 7535, South Africa*

<sup>4</sup> *Department of Physics and Astronomy, and National Superconducting Cyclotron Laboratory, Michigan State University, East Lansing, Michigan 48824-1321, USA and*

<sup>5</sup> *Department of Physics, Central Michigan University, Mount Pleasant, MI 48859, USA*

---

VOLUME 50, NUMBER 25

PHYSICAL REVIEW LETTERS

20 JUNE 1983

---

### New Look at Magnetic Moments and Beta Decays of Mirror Nuclei

B. Buck

*Department of Theoretical Physics, Oxford University, Oxford OX1 3RH, United Kingdom*

and

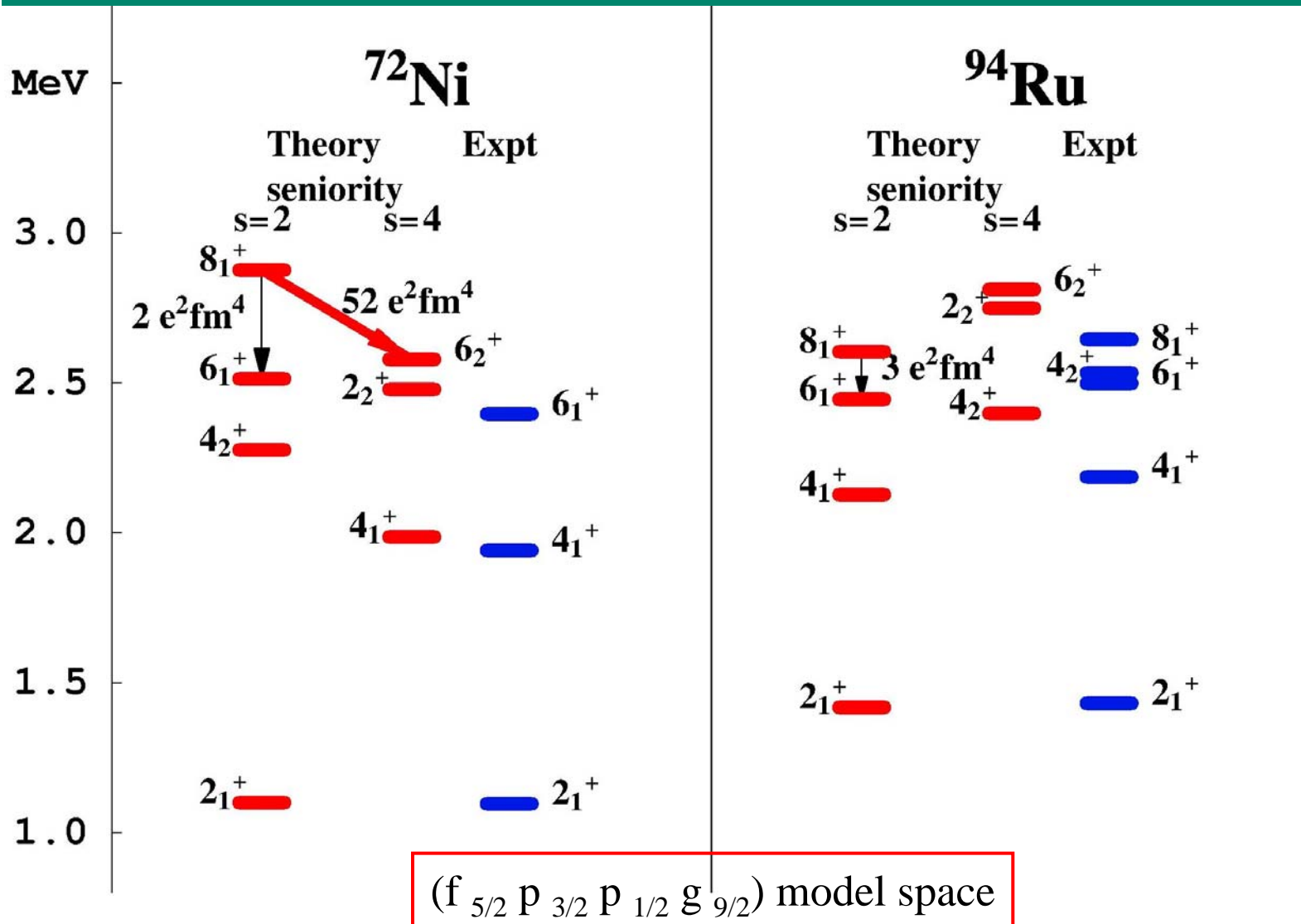
S. M. Perez



B. Alex Brown, USNDP, BNL, Nov 6, 2008

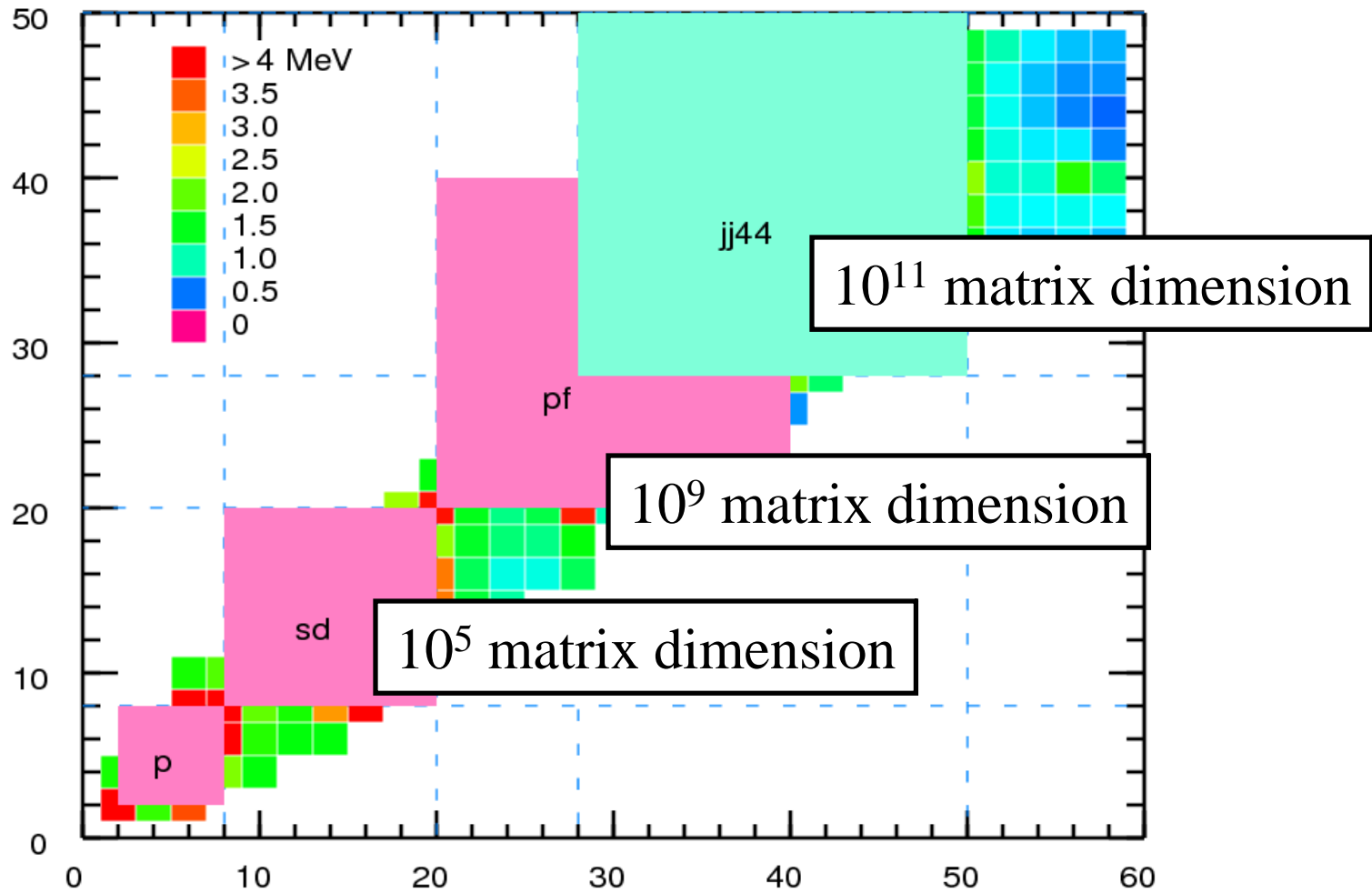


# Faded seniority isomerism near $^{78}\text{Ni}$



A. Lisetskiy et al., PRC 70, 044312 (2004)

B. Alex Brown, USNDP, BNL, Nov 6, 2008



jj44 means  $f_{5/2}$ ,  $p_{3/2}$ ,  $p_{1/2}$ ,  $g_{9/2}$  orbits for protons and neutrons

