

# Astrophysics Task Force

Caroline Nesaraja,  
Michael Smith  
ORNL Physics Division

Numerous USNDP institutions are pursuing projects that are very beneficial for studies in nuclear astrophysics

These activities include work on both nuclear reactions *and* nuclear structure

## Recent Activities in

Compilations & Evaluations

ANL, LANL, McMaster,  
ORNL

Development of Evaluation,  
Processing, & Dissemination Tools

ORNL

Nuclear Theory

LANL

# Compilations & Evaluations

## ANL (F. Kondev)

- **Measured** and **Evaluated** levels in  $^{186}\text{Re}$  that can modify the utilization of  $^{187}\text{Re}$  -  $^{187}\text{Os}$  abundances as a cosmochronometer to “date” the r-process

## LANL (G. M. Hale)

- **R-matrix calculations** of  $p(n,\gamma)$  and  $^{12}\text{C}(\alpha,\gamma)^{16}\text{O}$  reactions which are important in many astrophysics studies

# Compilations & Evaluations

## McMaster (A. Chen)

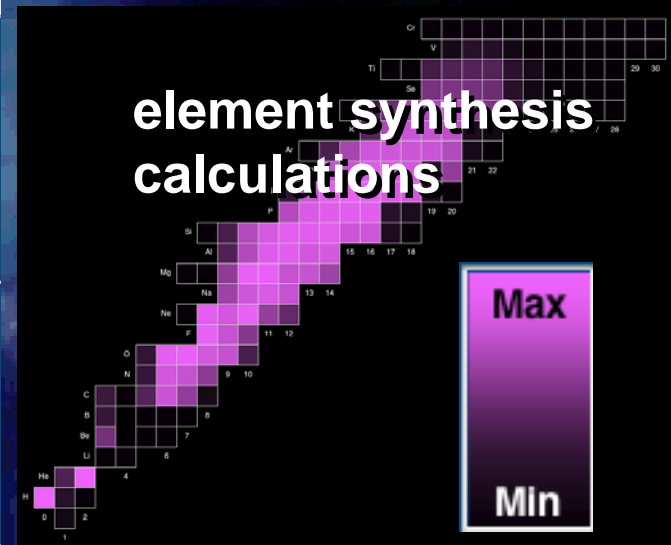
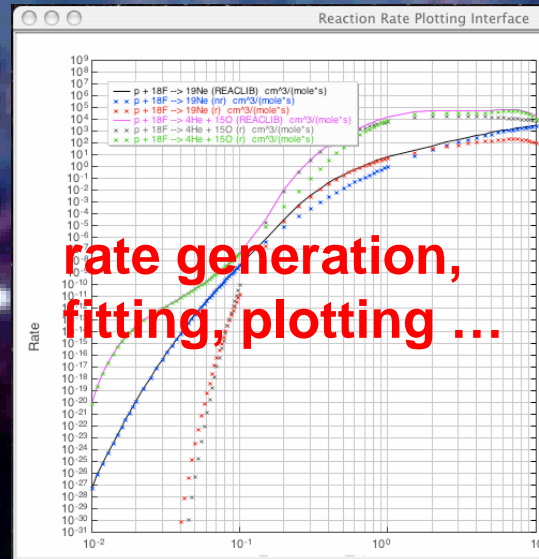
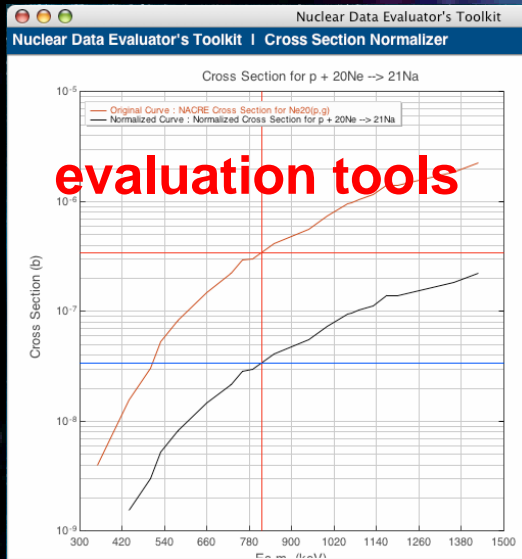
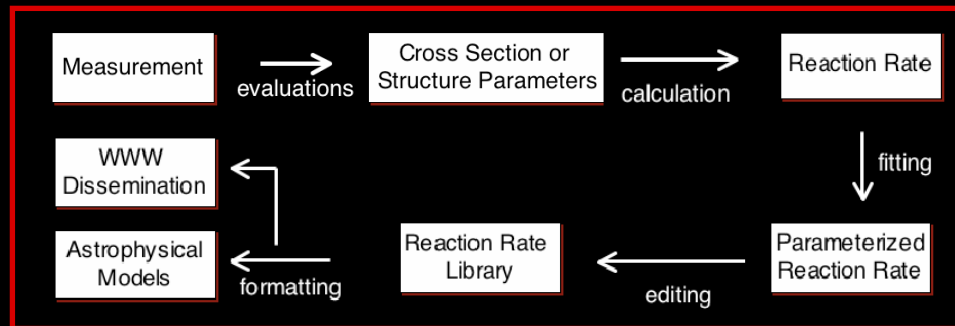
- Focus on reactions involving **radioactive nuclei** important for **stellar explosions** - coupled to ISAC measurements
- \* • Reaction evaluations in progress:  $^{13}\text{N}(p,\gamma)^{14}\text{O}$ ,  $^{25}\text{Al}(p,\gamma)^{26}\text{Si}$ ,  $^{18}\text{Ne}(\alpha,p)^{21}\text{Na}$ , &  $^{21}\text{Na}(p,\gamma)^{22}\text{Mg}$ ; others planned

## ORNL (C. Nesaraja)

- Focus on reactions involving **radioactive nuclei** important for **stellar explosions** - coupled to HRIBF measurements
- \* • Reaction evaluations in progress:  $^{18}\text{F}(p,\gamma)^{19}\text{Ne}$ ,  $^{18}\text{F}(p,\alpha)^{15}\text{O}$ ,  $^{30}\text{P}(p,\gamma)^{31}\text{S}$ ,  $^{33}\text{Cl}(p,\gamma)^{34}\text{Ar}$  ...
- \* All evaluations incorporating the **very latest experimental results**  
Results will be disseminated through **[nuastrodata.org](http://nuastrodata.org)** & NNDC

# Development of Evaluation, Processing, & Dissemination Tools

ORNL



Expansion of the **Computational Infrastructure for Nuclear Astrophysics** at [nucastrodata.org](http://nucastrodata.org) to include nuclear mass models, reaction rate commenting, reaction flux animation, export of movie files, improved performance, and many other features

# Nuclear Theory

LANL (T. Kawano)

- Calculation of  $^{95}\text{Zr}(n,\gamma)^{96}\text{Zr}$  cross section

Important for heavy element synthesis in the slow neutron capture process in red giant stars

- Parameterization of Nuclear Level Density Systematics (LANL - JAERI collaboration)

Important in calculations of reaction cross sections on unstable nuclei for use in r-process element synthesis

- Develop computer code for Direct Capture cross section calculations

Can help calculate rates on unmeasured reactions needed in astro simulations



# Nuclear Theory

LANL (P. Moller)

- Calculation of fission barriers for 3000 nuclei with  $A > 190$

Important for determining influence of fission on the r-process nuclei from the proton drip line to the neutron dripline

- Calculation of shape isomers (oblate/prolate, spherical/prolate...) for 7206 nuclei

Help understand structure of nuclei off stability for explosive burning in stars

- Calculation of  $\log(ft)$  values for electron capture on excited states in (neutron-rich and stable) nuclei

Help understand weak interactions that play a crucial role in the evolution and explosion of stars

In collaboration with MSU

# Summary

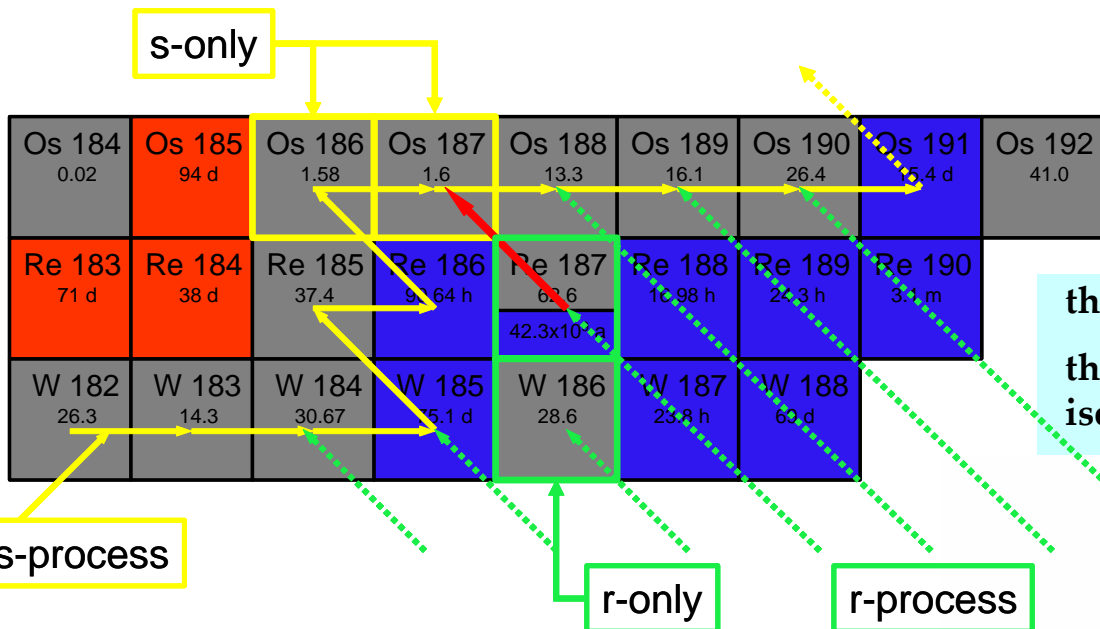
- Progress in understanding many **astrophysical phenomena requires improved nuclear data**
- Interesting, Important Astrophysics Projects involving
  - Structure & Reaction work
  - **Evaluations, Disseminations, Tool development, Nuclear Theory**
  - Multiple laboratories ANL, LANL, McMaster, ORNL
- **New computational infrastructure at [nucastrodata.org](http://nucastrodata.org)** now online to incorporate this information into astro models

# ANL: Studies of $^{186}\text{Re}$ of relevance for astrophysics

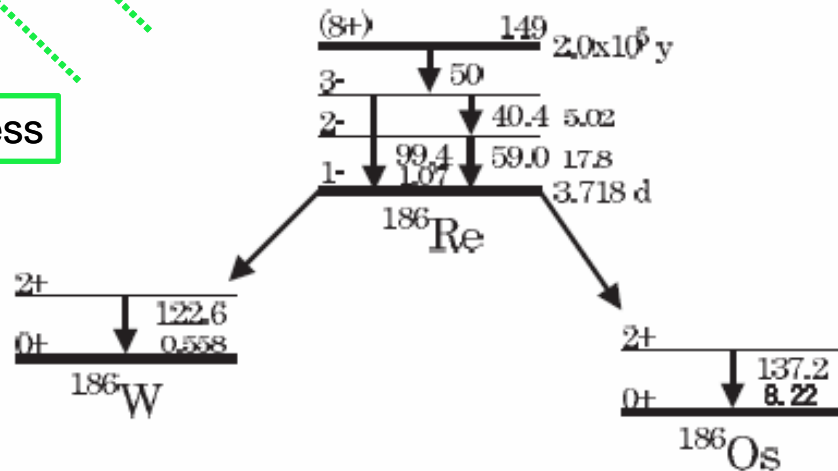
F. Kondev

$^{187}\text{Re}$ - $^{187}\text{Os}$  cosmochronometer can be used to date the r-process

D.D. Clayton, Ap.J. 139 (1964) 637.



the existence of long-lived isomeric state  
the production & destruction CS for the isomer are poorly known, but badly needed!



## Activation Technique

- ✓ difficult owing to the long  $T_{1/2}$
- ✓  $T_{1/2}$  - only one measurement without uncertainty

T. Hayakawa et al., Ap.J. 628 (2005) 533.

Using prompt  $\gamma$ -ray technique following n-capture, as recently demonstrated for  $^{239}\text{Pu}(n,2n)^{238}\text{Pu}$

L.A. Bernstein et al., Phys. Rev. C 65 (2003) 021601(R)

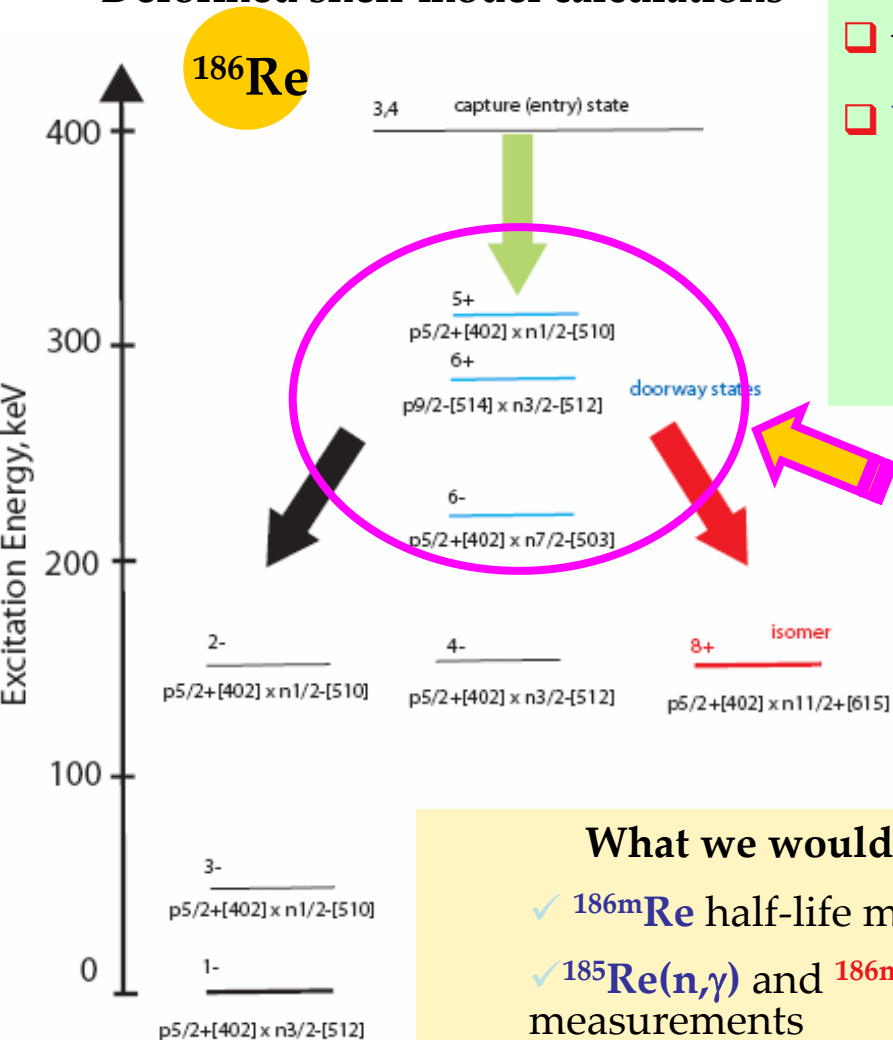
What is needed – detailed knowledge of the  $^{186}\text{Re}$  levels above the isomer!



# ANL: Studies of $^{186}\text{Re}$ of relevance for astrophysics

F. Kondev

## Deformed shell-model calculations



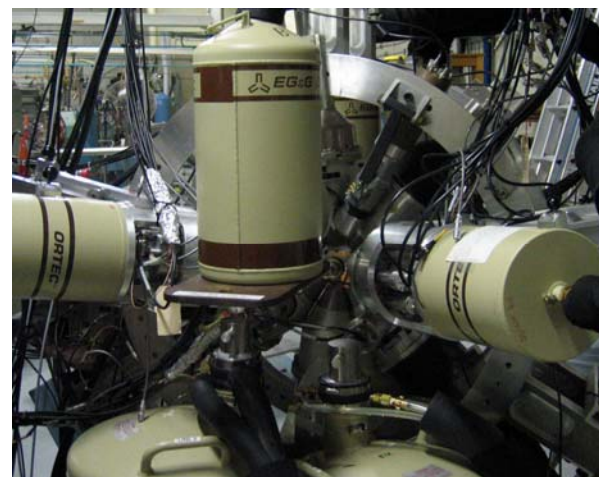
## New Measurements

- ❑ using  $^{186}\text{W}(d,2n)$  at ANU &  $\gamma$ -ray coin. technique
- ❑ DC beam & CAESAR array – 9 CSS Ge & 2 LEPS
  - ✓ excitation functions from 12 MeV to 18 MeV
  - ✓  $\gamma$ - $\gamma$  coin. at 12 MeV (near the barrier) – only a few channels are open – identification of  $^{186}\text{Re}$
  - ✓  $\gamma$ - $\gamma$  coin. at 14 MeV to enhance population of the isomer

“doorway” states above the isomer have been discovered and characterized!

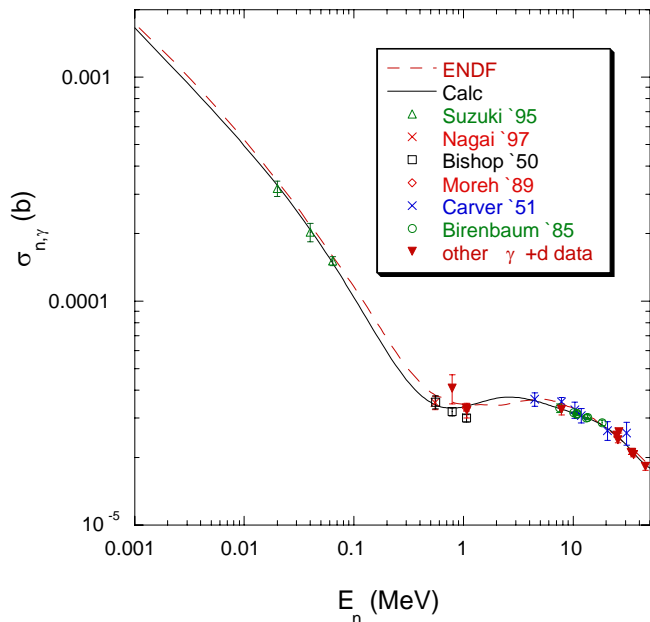
## What we would like to do next?

- ✓  $^{186}\text{mRe}$  half-life measurements
- ✓  $^{185}\text{Re}(n,\gamma)$  and  $^{186}\text{mRe}(n,\gamma)$  cross-section measurements

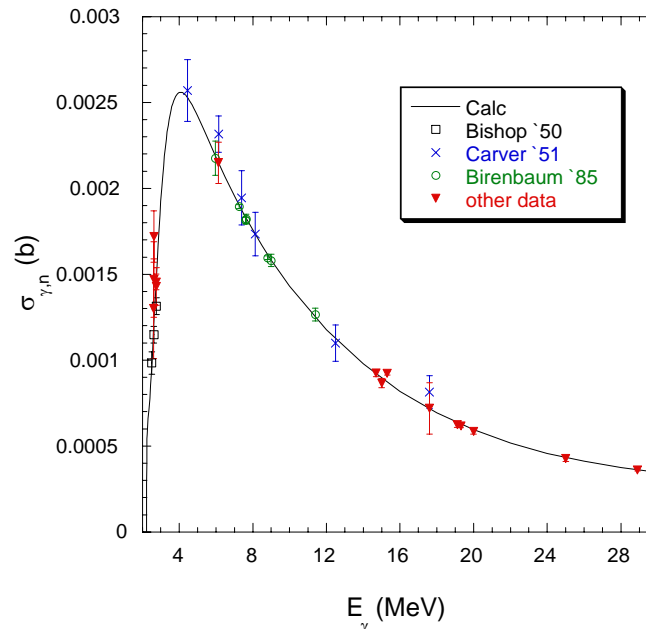


# LANL: n-p Capture Evaluation

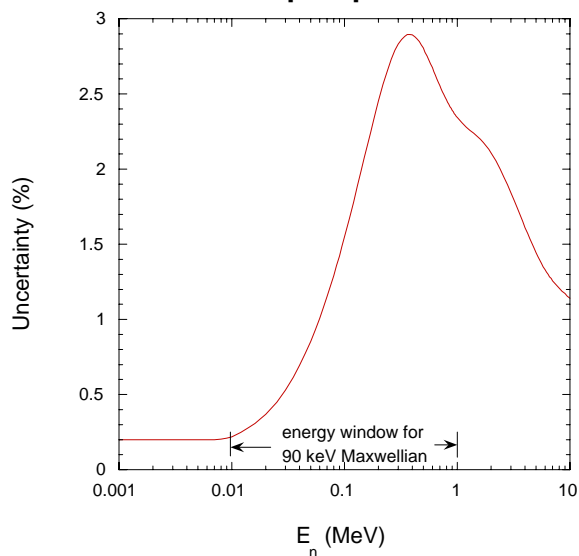
## $p(n,\gamma)d$ Cross Section



## $d(\gamma,n)p$ Cross Section



## n-p Capture

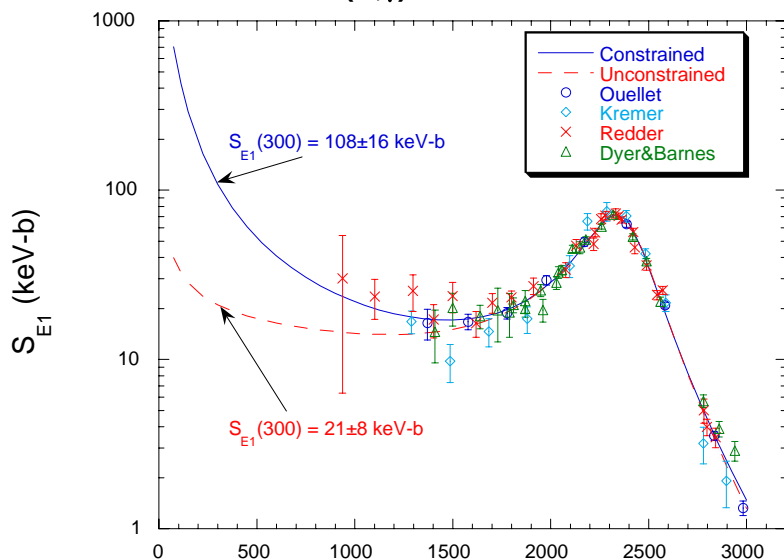


Data are shown above both as  $p(n,\gamma)$  and  $d(\gamma,n)$ .

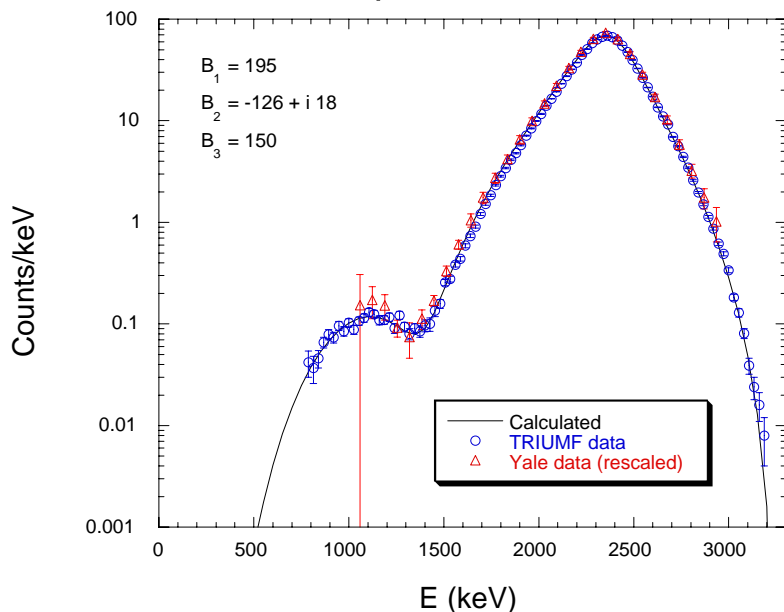
Calculated uncertainty for  $p(n,\gamma)$  is shown at the left.

# LANL: Update of $^{12}\text{C}(\alpha,\gamma)^{16}\text{O}$ Analysis

$^{12}\text{C}(\alpha,\gamma)^{16}\text{O}$  E1 S-factor



$^{16}\text{N}(\beta^-)^{16}\text{O} \rightarrow \alpha + ^{12}\text{C}$



- Uses new photon-channel prescription
- Includes elastic scattering data of Plaga, E1 capture data
- **Unconstrained:**  $\gamma_{1\gamma}$  fixed in -45 keV level by  $\Gamma_{1\gamma} = 0.05485$  eV.
- **Constrained:**  $|\gamma_{1\alpha}|$  fixed in -45 keV level by  $C^2 = (4.33 \pm 0.84) \times 10^{28} \text{ fm}^{-1}$  measured by Brune et al. [PRL **83**, 4025 (1999)]
- 
- Spectrum from the  $\beta$ -delayed  $\alpha$ -decay of  $^{16}\text{N}$  looks reasonable; height of first maximum determines the combination  $B_1|\gamma_{1\alpha}|$ .

## McMaster: Evaluations for Astrophysics

- Personnel: A. Chen (McMaster faculty)  
J. Pearson (postdoc, part-time)
- Reactions evaluated:  $^{25}\text{Al}(p,\gamma)^{26}\text{Si}$  and  $^{13}\text{N}(p,\gamma)^{14}\text{O}$ 
  - Additional ongoing evaluations:  $^{21}\text{Na}(p,\gamma)^{22}\text{Mg}$  and  $^{18}\text{Ne}(\alpha,p)^{21}\text{Na}$
  - Planned evaluations:  $^{26}\text{Al}(p,\gamma)^{27}\text{Si}$  (new data in 2005 from TRIUMF-ISAC)  
 $^{40}\text{Ca}(\alpha,\gamma)^{44}\text{Ti}$  (new data in 2005 from TRIUMF-ISAC)
  - New reaction rates will be disseminated through the new ORNL computational infrastructure at [www.nucastrodata.org](http://www.nucastrodata.org).

## McMaster: Evaluation of the $^{25}\text{Al}(p,\gamma)^{26}\text{Si}$ reaction rate

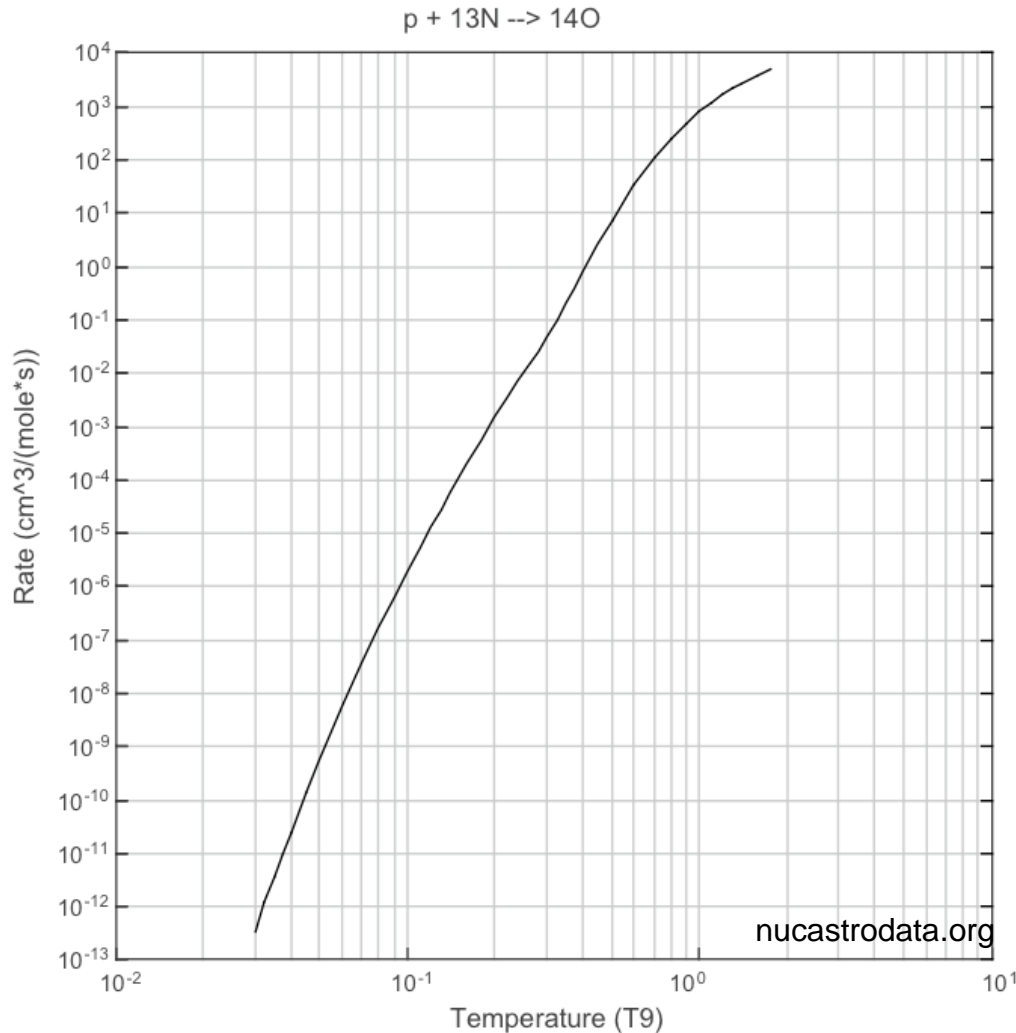
A. Chen

- Critical rate in the production of galactic  $^{26}\text{Al}$  in nova explosions.
- Incorporated recent data from various transfer reaction studies (e.g., (p,t) at ORNL, ( $^3\text{He}$ , $^6\text{He}$ ) at Yale University, and ( $^3\text{He}$ ,n) at Ohio University).
- **Evaluation will continue** as new results on this reaction become available from TRIUMF-ISAC and other laboratories, including a planned direct measurement of the cross section at ISAC.



# McMaster: Evaluation of the $^{13}\text{N}(p,\gamma)^{14}\text{O}$ reaction rate

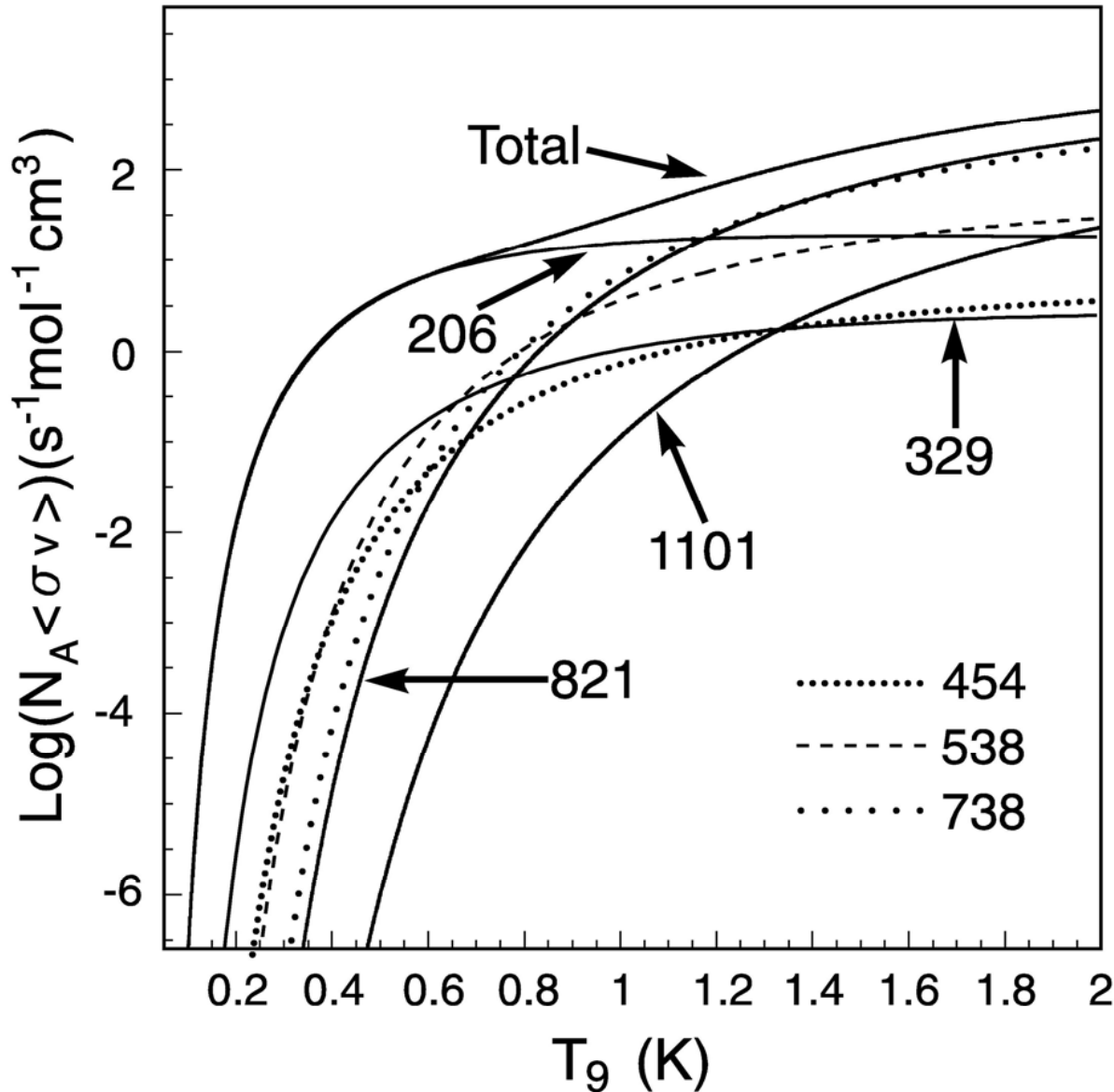
A. Chen



- Important in the breakout from the CNO cycles to the Hot-CNO cycles in novae and X-ray bursts.
- Incorporated all data on this reaction to date.
- New direct measurement of the reaction rate is planned at TRIUMF-ISAC.

# McMaster: Evaluation of the $^{21}\text{Na}(p,\gamma)^{22}\text{Mg}$ reaction rate

A. Chen



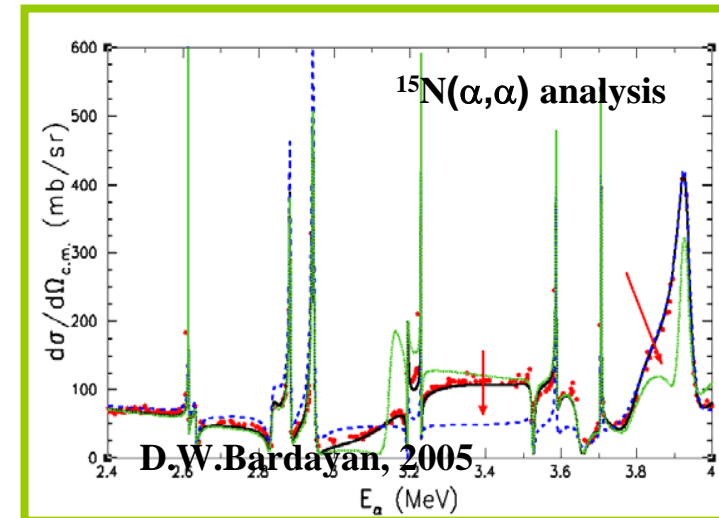
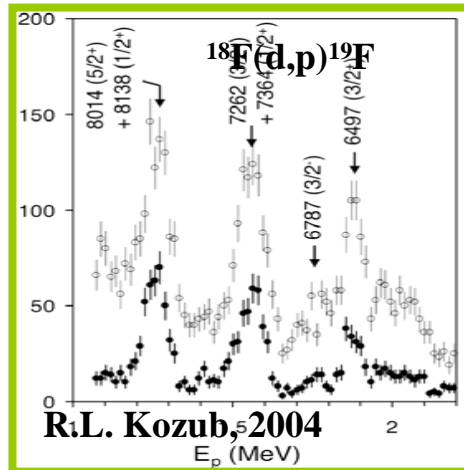
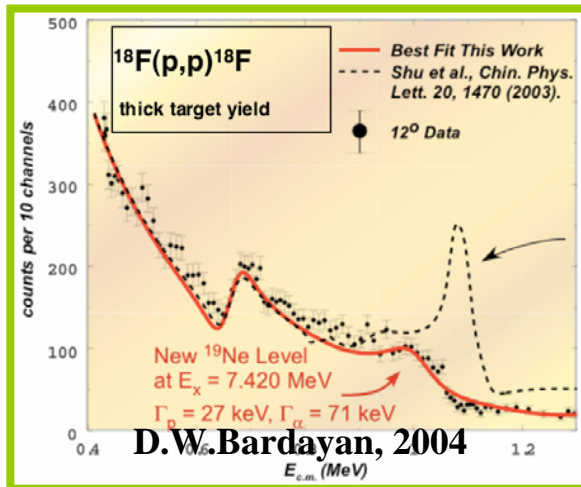
$^{21}\text{Na}(p,\gamma)^{22}\text{Mg}$  reaction rate based on recent ISAC results (D'Auria et al. PRC 2004)

Resonance energies in keV

# ORNL : Evaluation of $^{18}\text{F}(p,\alpha)^{15}\text{O}$ & $^{18}\text{F}(p,\gamma)^{19}\text{Ne}$

C. Nesaraja

- Important for novae & X-ray bursts
- Many experiments performed with radioactive beams require new evaluations
- New evaluation of  $^{19}\text{F}$  levels from the legacy of  $^{15}\text{N}(\alpha,\alpha)$  data



**Detailed nuclear structure information for  $^{19}\text{Ne}$**   
 ~ 30 levels between  $E_x = 6.411 - 8.100$  MeV

**Excitation energies for missing levels**

(scaled using Thomas-Ehrman shift calculations )

**spin and parity assignments**

(angular momentum transfer or  $J^\pi$  of mirror nuclei)

**Partial gamma widths  $\Gamma_\gamma$**

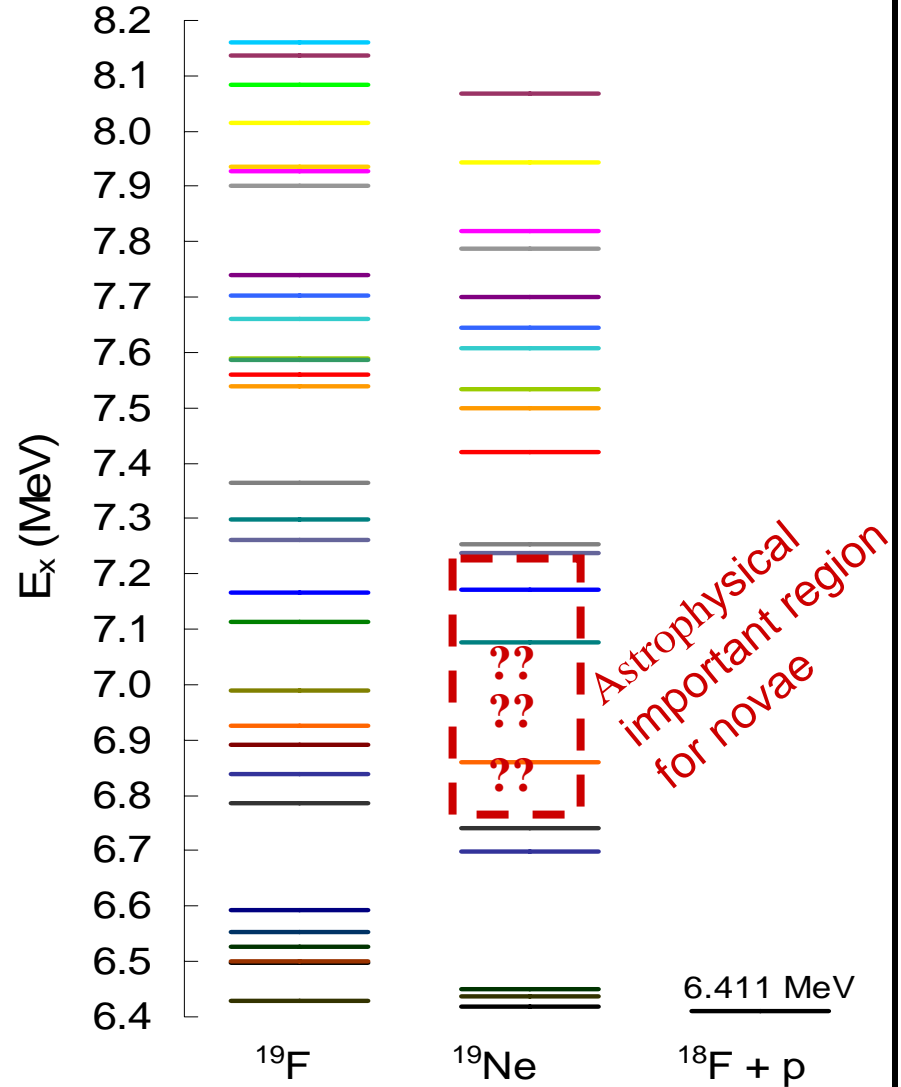
(corrected for phase space & reduced transition probability for  $\Gamma_\gamma$  of mirror nuclei)

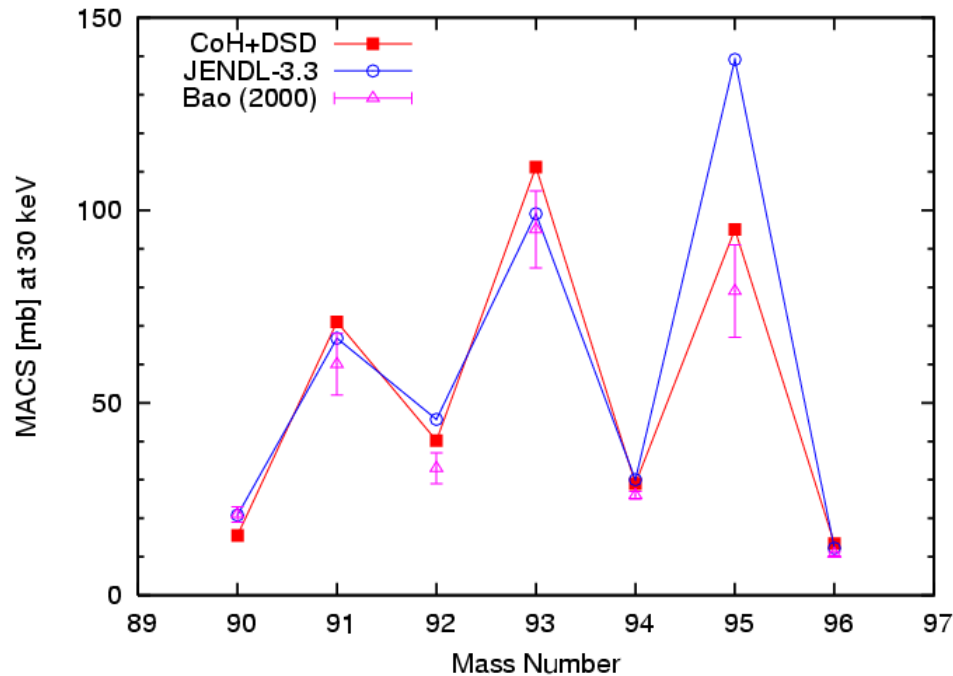
**Partial proton widths  $\Gamma_p$**

(spectroscopic factor and using a diffuse surface potential or by direct measurements )

**Partial alpha widths  $\Gamma_\alpha$**

(assuming analog states have same reduced widths & correcting for the Coulomb barrier penetration)





Calculation of  $^{95}\text{Zr}(n,\gamma)^{96}\text{Zr}$  cross section based on new experimental data for  $^{90,91,92,94}\text{Zr}$  and systematics of nuclear model parameters (with R. Reifarth & F. Herwig of LANL)

Important for heavy element synthesis in the slow neutron capture process in red giant stars



# LANL: Nuclear Theory

T. Kawano

Parameterization of phenomenological Nuclear Level Density using shell & pairing energies of KTUY05 mass formula  
(LANL - JAERI collaboration)

Important in calculations of reaction cross sections on unstable nuclei for use in r-process element synthesis

Develop computer code for Direct Capture cross section calculations

Using DSD theory with Hartree-Fock BCS calculations for bound state wave functions

Can help calculate rates on unmeasured reactions needed in astro simulations

# LANL: Nuclear Theory

## Calculation of fission barriers for 3000 nuclei with $A > 190$

Important for determining influence of fission on the r-process  
nuclei from the proton drip line to the neutron dripline

use a 5 dimensional 3 quadratic surface parameterization of  
nuclear shapes

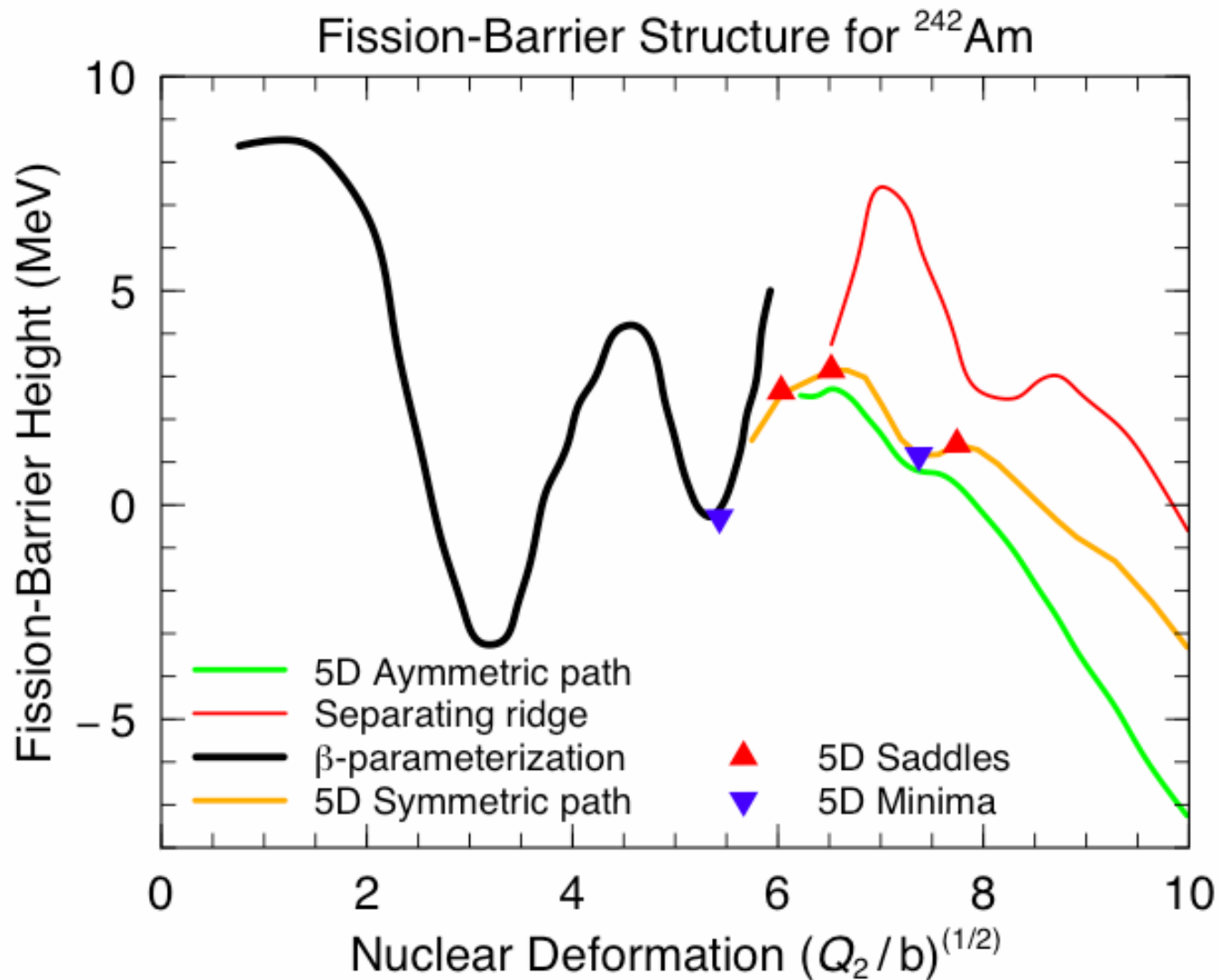
goals: determine optimum saddle point for each nucleus and,  
later, complete potential structure

developed highly automated procedures to interpret the widely-  
varying calculated surfaces.

Calculations are set up as 15 production jobs (different regions of  
the N/Z plane), each takes ~10 days with 30 500MHz computers

**Exciting result: for U and Pu there is a region of low fission  
barriers (around 4 MeV) around  $A=260$**



Calculation of fission barriers for 3000 nuclei with  $A > 190$ 

## Calculation of shape isomers (oblate/prolate, spherical/prolate...) for 7206 nuclei

Help understand structure of nuclei off stability for explosive burning in stars

Develop classification of shapes and shape-isomer types (oblate-prolate, triple...) for global N/Z plots

Will compare to experimental data and will publish in ATNDT