

$^{16}\text{O}(^{82}\text{Se},3\text{n}\gamma)$ **2009Zh11**

Type	Author	History	Citation	Literature Cutoff Date
Full Evaluation	S. K. Basu, G. Mukherjee, A. A. Sonzogni		NDS 111, 2555 (2010)	30-Jun-2009

2009Zh11: $^{16}\text{O}(^{82}\text{Se},3\text{n}\gamma)$, E=460 MeV; measured $E\gamma$, $I\gamma$, $\gamma\gamma$; deduced angular distribution asymmetry ratios; GASP array consisting of 40 Compton-suppressed Ge detectors and a multiplicity filter of 80 BGO elements. Comparisons with shell-model calculations.

 ^{95}Mo Levels

E(level) [†]	J ^π	E(level) [†]	J ^π	E(level) [†]	J ^π	E(level) [†]	J ^π
0.0 [‡]	5/2 ⁺	2232.2 [‡] 6	(15/2 ⁺)	4048.1 [#] 8	(23/2 ⁻)	7451.2 [#] 12	(37/2 ⁻)
765.9 [‡] 4	7/2 ⁺	2580.3 [‡] 6	(17/2 ⁺)	4140.0 [‡] 10	(29/2 ⁺)	7985.5 [#] 12	(39/2 ⁻)
947.6 [‡] 4	9/2 ⁺	2611.4 [#] 6	(15/2 ⁻)	5117.6 [#] 9	(27/2 ⁻)	8424.7 [‡] 13	(37/2 ⁺)
1540.6 [‡] 5	11/2 ⁺	2618.4 [‡] 7	(19/2 ⁺)	5361.9 [‡] 11	(31/2 ⁺)	9654.8 [#] 13	(41/2 ⁻)
1551.7 5	(9/2 ⁺)	2770.1 [‡] 8	(21/2 ⁺)	5760.6 [#] 10	(31/2 ⁻)	10509.0 [#] 14	(45/2 ⁻)
1937.6 [#] 6	11/2 ⁻	3277.4 [#] 8	(19/2 ⁻)	6327.6 [‡] 12	(35/2 ⁺)		
2058.3 6	(13/2 ⁺)	3672.5 [‡] 9	(25/2 ⁺)	6708.8 [#] 11	(35/2 ⁻)		

[†] From least-squares fit to $E\gamma$'s.

[‡] Band(A): γ cascade based on 5/2⁺.

[#] Band(B): γ cascade based on (11/2⁻).

 $\gamma(^{95}\text{Mo})$

Angular distribution asymmetry ratios R_{ADO} are $I\gamma(34^\circ)/I\gamma(90^\circ)$; where $I\gamma$ were obtained from $\gamma\gamma$ coincidence spectra gated with any multipolarity. Typical values are significantly larger than 1 for $\Delta J=2$, quadrupole and <1 for $\Delta J=1$, dipole transitions. It should be noted that R_{ADO} value can be the same for $\Delta J=0$, dipole and for $\Delta J=1$, mixed transitions.

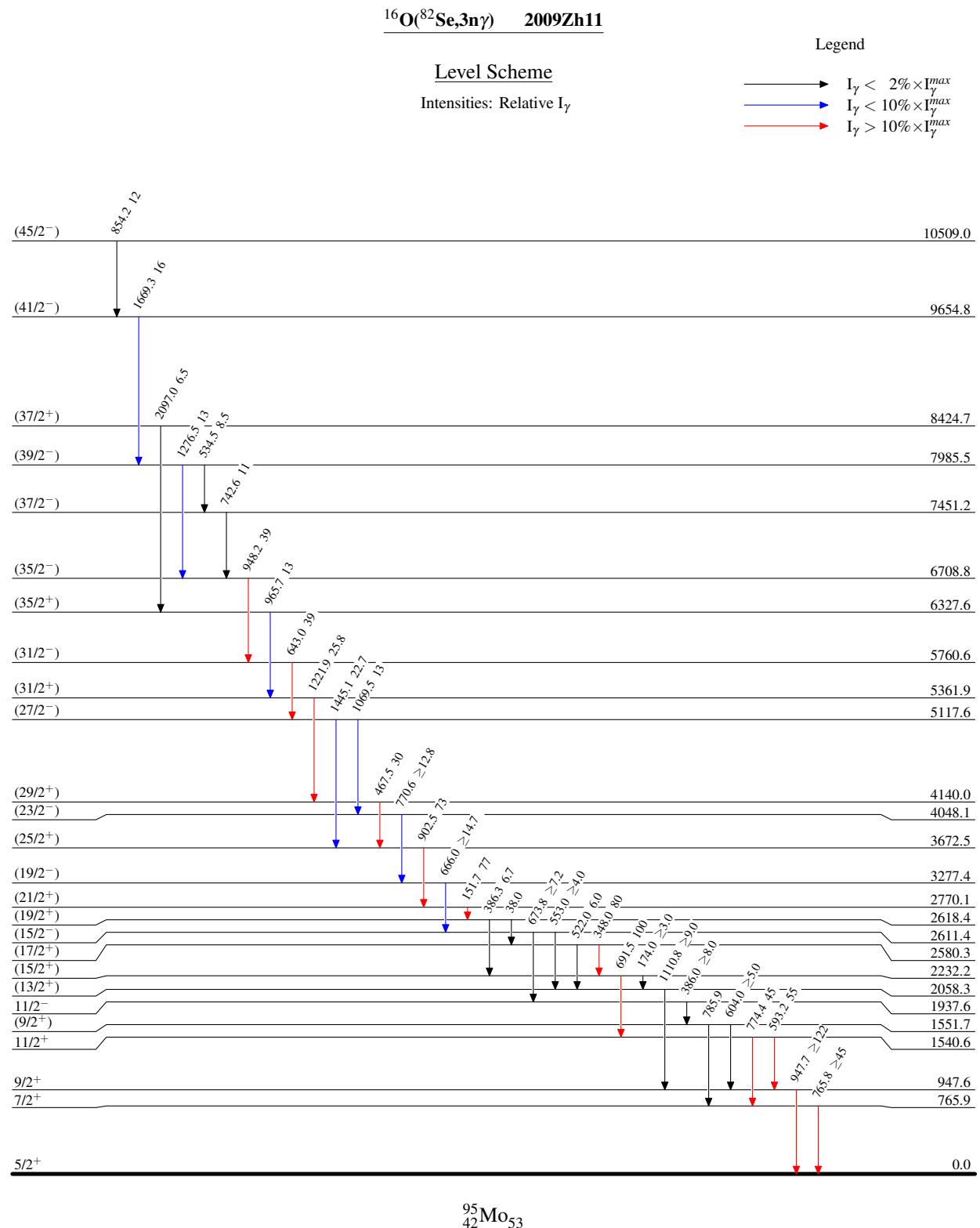
E_γ [†]	I_γ [†]	E_i (level)	J_i^π	E_f	J_f^π	Comments
38.0 5		2618.4	(19/2 ⁺)	2580.3	(17/2 ⁺)	
151.7 5	77 8	2770.1	(21/2 ⁺)	2618.4	(19/2 ⁺)	$R_{ADO}=0.86$ 21.
174.0 5	\geq 3.0	2232.2	(15/2 ⁺)	2058.3	(13/2 ⁺)	
348.0 5	80 8	2580.3	(17/2 ⁺)	2232.2	(15/2 ⁺)	$R_{ADO}=0.86$ 5.
386.0 5	\geq 8.0	1937.6	11/2 ⁻	1551.7	(9/2 ⁺)	$R_{ADO}=0.85$ 12.
386.3 5	6.7 20	2618.4	(19/2 ⁺)	2232.2	(15/2 ⁺)	$R_{ADO}=1.63$ 24.
467.5 5	30 3	4140.0	(29/2 ⁺)	3672.5	(25/2 ⁺)	$R_{ADO}=1.65$ 18.
522.0 5	6.0 18	2580.3	(17/2 ⁺)	2058.3	(13/2 ⁺)	
534.5 5	8.5 26	7985.5	(39/2 ⁻)	7451.2	(37/2 ⁻)	$R_{ADO}=0.54$ 10.
553.0 5	\geq 4.0	2611.4	(15/2 ⁻)	2058.3	(13/2 ⁺)	
593.2 5	55 6	1540.6	11/2 ⁺	947.6	9/2 ⁺	$R_{ADO}=0.80$ 11.
604.0 5	\geq 5.0	1551.7	(9/2 ⁺)	947.6	9/2 ⁺	
643.0 5	39 4	5760.6	(31/2 ⁻)	5117.6	(27/2 ⁻)	$R_{ADO}=1.92$ 16.
666.0 5	\geq 14.7	3277.4	(19/2 ⁻)	2611.4	(15/2 ⁻)	$R_{ADO}=1.80$ 20.
673.8 5	\geq 7.2	2611.4	(15/2 ⁻)	1937.6	11/2 ⁻	$R_{ADO}=2.07$ 25.
691.5 5	100 10	2232.2	(15/2 ⁺)	1540.6	11/2 ⁺	$R_{ADO}=1.65$ 14.
742.6 5	11 3	7451.2	(37/2 ⁻)	6708.8	(35/2 ⁻)	$R_{ADO}=1.12$ 13.
765.8 5	\geq 45	765.9	7/2 ⁺	0.0	5/2 ⁺	$R_{ADO}=0.90$ 8.
770.6 5	\geq 12.8	4048.1	(23/2 ⁻)	3277.4	(19/2 ⁻)	$R_{ADO}=1.90$ 35.
774.4 5	45 5	1540.6	11/2 ⁺	765.9	7/2 ⁺	$R_{ADO}=1.60$ 13.
785.9 5		1551.7	(9/2 ⁺)	765.9	7/2 ⁺	

Continued on next page (footnotes at end of table)

$^{16}\text{O}(^{82}\text{Se},3n\gamma)$ **2009Zh11 (continued)** $\gamma(^{95}\text{Mo})$ (continued)

E_γ^\dagger	I_γ^\dagger	$E_i(\text{level})$	J_i^π	E_f	J_f^π	Comments
854.2 5	12 4	10509.0	(45/2 ⁻)	9654.8	(41/2 ⁻)	$R_{ADO}=1.45~15.$
902.5 5	73 7	3672.5	(25/2 ⁺)	2770.1	(21/2 ⁺)	$R_{ADO}=1.70~14.$
947.7 5	≥ 122	947.6	9/2 ⁺	0.0	5/2 ⁺	$R_{ADO}=2.2~4.$
948.2 5	39 4	6708.8	(35/2 ⁻)	5760.6	(31/2 ⁻)	$R_{ADO}=1.6~4.$
965.7 5	13 4	6327.6	(35/2 ⁺)	5361.9	(31/2 ⁺)	$R_{ADO}=1.7~3.$
1069.5 5	13 4	5117.6	(27/2 ⁻)	4048.1	(23/2 ⁻)	$R_{ADO}=1.60~15.$
1110.8 5	≥ 9.0	2058.3	(13/2 ⁺)	947.6	9/2 ⁺	$R_{ADO}=1.7~3.$
1221.9 5	25.8 26	5361.9	(31/2 ⁺)	4140.0	(29/2 ⁺)	$R_{ADO}=0.65~10.$
1276.5 5	13 4	7985.5	(39/2 ⁻)	6708.8	(35/2 ⁻)	$R_{ADO}=1.78~20.$
1445.1 5	22.7 23	5117.6	(27/2 ⁻)	3672.5	(25/2 ⁺)	$R_{ADO}=0.85~8.$
1669.3 5	16 5	9654.8	(41/2 ⁻)	7985.5	(39/2 ⁻)	$R_{ADO}=1.15~18.$
2097.0 5	6.5 20	8424.7	(37/2 ⁺)	6327.6	(35/2 ⁺)	$R_{ADO}=1.4~3.$

[†] [2009Zh11](#) state that energy uncertainties are within 0.5 keV and the intensity uncertainties within 30%. A general energy uncertainty of 0.5 keV for all γ rays, 10% intensity uncertainty for $I_\gamma > 20$ and 30% for others have been assumed.



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