# $^{58}$ Ni( $^{40}$ Ca,2p2n $\gamma$ ) 2003Ma24

History

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E=135 MeV, measured E $\gamma$ , I $\gamma$ ,  $\gamma\gamma$ , using the GASP detector array comprised of 40 Compton-suppressed Ge detectors in conjuction with an 80 element BGO inner ball where the six elements of the most forward ring were replaced by the n-ring detector consisting of six liquid scintillator detectors. The ISIS Si ball, a 40-element  $\Delta$ E-E telescope array, was also used.

### 94Pd Levels

E(level) <sup>†</sup>	$J^{\pi \#}$	$T_{1/2}^{\bigcirc}$	E(level) <sup>†</sup>	$J^{\pi \#}$	E(level) <sup>†</sup>	$J^{\pi \#}$	$T_{1/2}^{@}$
$0_{\ddagger}$	$0_{+}$	9.0 s 5	2538.0 16 2699.0 17 2704.0 <sup>‡</sup> 19 3051.0 21	(5 <sup>-</sup> )	3796.0 <sup>‡</sup> 21	10 <sup>+</sup>	
814.0 <sup>‡</sup> <i>10</i>	2+		2699.0 <i>17</i>	$(6^{+})$	4663.0 23		
1720.0 <sup>‡</sup> <i>14</i>	4+		2704.0 <sup>‡</sup> 19	8+	4790.0 <sup>‡</sup> 23	12+	
2380.0 <sup>‡</sup> <i>16</i>	6+		3051.0 <i>21</i>	$(8^{+})$	4886 <sup>‡</sup> <i>3</i>	14+	$0.530~\mu s~10$

<sup>&</sup>lt;sup>†</sup> From least-squares fit to Ey assuming  $\Delta$ Ey=1 keV.

#### $\gamma$ (94Pd)

$E_{\gamma}$	$E_i(level)$	$\mathbf{J}_i^{\pi}$	$\mathbf{E}_f$	$\mathbf{J}_f^{\pi}$	Mult.	$\alpha^{\dagger}$	Comments
96	4886	14+	4790.0	12+	E2	1.63	$\alpha(\exp)=1.9 \ 4$ $\alpha=1.63; \ \alpha(K)=1.25 \ 4; \ \alpha(L)=0.311 \ 10; \ \alpha(M)=0.0597 \ 18; \ \alpha(N+)=0.0104$
							u=1.03, u(K)=1.23 7, $u(L)=0.311$ 10, $u(M)=0.037$ 10, $u(M+)=0.010$ 4
							Mult.: from $\alpha(\exp)=1.9$ 4, which was determined from intensity balance.
158	2538.0	$(5^{-})$	2380.0	6+			•
324	2704.0	8+	2380.0	6+			
347	3051.0	$(8^{+})$	2704.0	8+			
660	2380.0	6+	1720.0	4+			
745	3796.0	10+	3051.0	$(8^{+})$			
814	814.0	2+	0	0+	[E2]		
818	2538.0	$(5^{-})$	1720.0	4+			
867	4663.0		3796.0	$10^{+}$			
906	1720.0	4+	814.0	2+			
979	2699.0	$(6^{+})$	1720.0	4+			
994	4790.0	12+	3796.0	$10^{+}$			
1092	3796.0	10 <sup>+</sup>	2704.0	8+			

<sup>&</sup>lt;sup>†</sup> Total theoretical internal conversion coefficients, calculated using the BrIcc code (2008Ki07) with Frozen orbital approximation based on  $\gamma$ -ray energies, assigned multipolarities, and mixing ratios, unless otherwise specified.

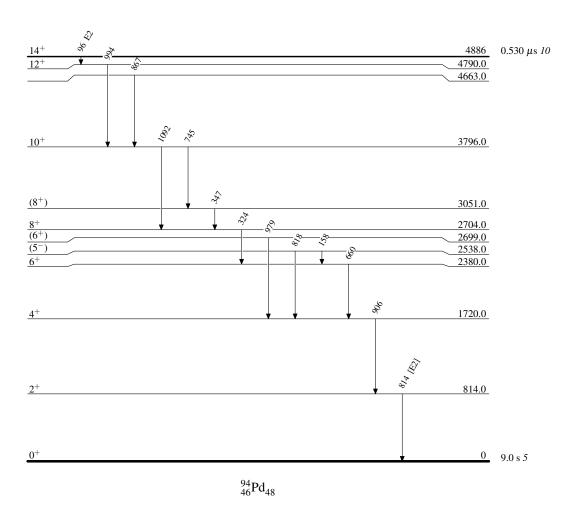
<sup>‡</sup> Band(A): yrast band.

<sup>&</sup>lt;sup>#</sup> As given by 2003Ma24, based on  $\gamma(\theta)$ , systematics of N=48 nuclei, and supported by shell model calculations.

<sup>&</sup>lt;sup>®</sup> From Adopted Levels.

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## Level Scheme



# $\frac{58}{\text{Ni}}(^{40}\text{Ca},2\text{p2n}\gamma)$ 2003Ma24

 $Band(A)\hbox{:}\ Yrast\ band$ 

