### <sup>71</sup>Ga(<sup>3</sup>He,t),(<sup>3</sup>He,tγ) 2015Fr02,1998Ej03,1988Ch25

History									
Туре	Author	Citation	Literature Cutoff Date						
Full Evaluation	Balraj Singh and Jun Chen	NDS 188,1 (2023)	17-Jan-2023						

Target  $J^{\pi}=3/2^{-}$ .

The <sup>71</sup>Ga(<sup>3</sup>He,t) studies are relevant to determination of neutrino capture rates in <sup>71</sup>Ga, a detector of solar neutrino flux.

2015Fr02 (also 2011Fr15): (<sup>3</sup>He,t),E=420 MeV (450 MeV in 2011Fr15). Measured triton spectra, and angular distributions using Grand Raiden spectrometer with FWHM=45 keV at RCNP, Osaka cyclotron facility. DWBA analysis of  $\sigma(\theta)$  data. Deduced B(GT) and solar-neutrino capture rates.

1998Ej03 (also 1996Fu06,1994Fu11): (<sup>3</sup>He,t),(<sup>3</sup>He,t $\gamma$ ),E=450 MeV from RCNP ring cyclotron. Measured triton energies at 0-2° in the forward direction in singles and coin with  $\gamma$  rays. Tritons were analyzed by GRAND RAIDEN magnetic spectrometer with FWHM=300-400 keV. The  $\gamma$  rays were detected using six NaI(Tl) detectors with FWHM=100-300 keV. The target was enriched and 2 mg/cm<sup>2</sup> thick. The CASCADE and GEANT codes were used to evaluate and simulate  $\gamma$ -spectrum. Deduced Gamow-Teller strengths and solar neutrino capture rates.

1988Ch25: (<sup>3</sup>He,t $\gamma$ ),(<sup>3</sup>He,t): E=29.9 MeV. Measured E(t) in singles (at 0°) and coin with  $\gamma$  rays using Q3D magnetic spectrometer for tritons and NaI(Tl) detector for  $\gamma$  rays. 99.8% enriched target. Detailed study of the isobaric analog state at 8.9 MeV.

1984Ko10 (from the same group as 1988Ch25): (<sup>3</sup>He,t) E=20.42 MeV. Measured triton spectra using Q3D spectrometer, deduced Q value and excited states in <sup>71</sup>Ge.

Not all the peaks shown in triton spectra of 1998Ej03 and 1984Ko10 are identified.

#### <sup>71</sup>Ge Levels

B(GT)=2.88 9 (1998Ej03) for 1580-7420 energy region.

E(level) <sup>†</sup>	$J^{\pi \ddagger}$	B(GT) <sup>#</sup>	Comments
0	1/2-	0.0852 11	$d\sigma/d\Omega$ (q=0)=0.786 mb/sr 9 (2015Fr02). Solar neutrino capture rate=109.8 SNU 13 (2015Fr02). B(GT)=0.0891 13 (1908Fi03)
175 <i>1</i>	5/2-	0.0034 26	$d\sigma/d\Omega(q=0)=0.071 \text{ mb/sr } 4 (2015\text{Fr02}).$ Solar neutrino capture rate=1.2 SNU 7 (2015Fr02). B(GT)=0.0049 <i>l</i> 8 (1998Fi03)
500 1	3/2-	0.0176 14	$d\sigma/d\Omega(q=0)=0.171 \text{ mb/sr } 4 (2015\text{Fr02}).$ Solar neutrino capture rate=2.7 SNU 2 (2015Fr02). B(GT)=0.0208 2 <i>I</i> (1998F103)
708 1	3/2-	0.0011 5	$d\sigma/d\Omega(q=0)=0.018$ mb/sr 1 (2015Fr02). Solar neutrino capture rate=0.03 SNU 1 (2015Fr02).
808 1	1/2-	0.0229 10	E(level): 830 in 1998Ej03. $d\sigma/d\Omega(q=0)=0.210$ mb/sr 4 (2015Fr02). Solar neutrino capture rate=0.61 SNU 7 (2015Fr02). B(GT)=0.0237.23 (1908E103)
1096 <i>1</i>	3/2-	0.0183 17	$d\sigma/d\Omega(q=0)=0.184 \text{ mb/sr } 4$ (2015Fr02). Solar neutrino capture rate=0.33 SNU 3 (2015Fr02). B(GT)=0.0233 24 (1998Ej03). E(level): 1096 (1984Ko10) 1160 (1998Ej03)
1299 <i>1</i>	3/2-	0.0133 8	$d\sigma/d\Omega(q=0)=0.126 \text{ mb/sr } 2 (2015\text{Fr02}).$ Solar neutrino capture rate=0.17 SNU <i>I</i> (2015Fr02). B(GT)=0.0201 <i>23</i> (1998Ej03). E(level): 1299 (1984Ko10), 1360 in 1998Ei03 could be a composite of 1299 and 1378 keV peaks.
1378 <i>1</i>	5/2-	0.0033 4	$d\sigma/d\Omega(q=0)=0.035$ mb/sr 3 (2015Fr02). Solar neutrino capture rate=0.041 SNU 5 (2015Fr02)
1598 <i>1</i>	3/2-	0.0011 5	$d\sigma/d\Omega(q=0)=0.018$ mb/sr 2 (2015Fr02).
1744 <i>1</i>	3/2-	0.0068 2	$d\sigma/d\Omega(q=0)=0.061 \text{ mb/sr } I (2015Fr02).$
1964 <i>1</i>	3/2-	0.0012 6	Solar neutrino capture rate=0.07/ SNU 3 (2015Fr02). $d\sigma/d\Omega(q=0)=0.020$ mb/sr 2 (2015Fr02).

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## <sup>71</sup>Ga(<sup>3</sup>He,t),(<sup>3</sup>He,tγ) 2015Fr02,1998Ej03,1988Ch25 (continued)

# <sup>71</sup>Ge Levels (continued)

E(level) <sup>†</sup>	$J^{\pi \ddagger}$	T <sub>1/2</sub>	B(GT) <sup>#</sup>	Comments
2041 1	(2 2-5 2-)		0.0191.2	Solar neutrino capture rate=0.013 SNU 6 (2015Fr02). d = (40)(a = 0) = 0.187 mb/sm 2 (2015Fr02)
2041 1	(3/2 ,3/2 )		0.0181 2	$d\sigma/ds_2(q=0)=0.187$ III0/SI 5 (2015F102). Solar neutrino canture rate=0.187 SNU 6 (2015Fr02)
2145 <i>I</i>	$(3/2^{-}, 5/2^{-})$		0.0031 6	$d\sigma/d\Omega(q=0)=0.036$ mb/sr <i>1</i> (2015Fr02).
	., ,, ,			Solar neutrino capture rate=0.031 SNU 6 (2015Fr02).
2276 1	$(3/2^-, 5/2^-)$		0.0036 9	$d\sigma/d\Omega(q=0)=0.046 \text{ mb/sr } 1 \text{ (2015Fr02)}.$
2252 1	(2 0-5 0-)		0.0120.28	Solar neutrino capture rate= $0.035 \text{ SNU } 9 (2015\text{Fr}02)$ .
2552 1	(3/2 ,3/2 )		0.0130 28	$d\sigma/d\Omega_2(q=0)=0.137$ mb/sr 2 (2015F102). Solar neutrino capture rate=0.12 SNU 3 (2015Fr02)
2435 1	$(3/2^{-}, 5/2^{-})$		0.0123 17	$d\sigma/d\Omega(q=0)=0.133$ mb/sr 2 (2015Fr02).
	., ,, ,			Solar neutrino capture rate=0.11 SNU 2 (2015Fr02).
2642 1	$(5/2^{-})$		0.0054 10	$d\sigma/d\Omega(q=0)=0.062 \text{ mb/sr } 1 \text{ (2015Fr02)}.$
1 9770	$(5/2^{-})$		0.0058 12	Solar neutrino capture rate= $0.046$ SNU 9 (2015Fr02).
2770 1	(3/2)		0.0038 12	d0/d22(q=0)=0.070 mb/sr 7 (2015F102). Solar neutrino canture rate=0.048 SNU 10 (2015Fr02)
2806 1	$(5/2^{-})$		0.0172 12	$d\sigma/d\Omega(q=0)=0.165 \text{ mb/sr } 3 \text{ (2015Fr02)}.$
	., ,			Solar neutrino capture rate=0.140 SNU 11 (2015Fr02).
2888 1	$(5/2^{-})$		0.0019 7	$d\sigma/d\Omega(q=0)=0.028$ mb/sr 4 (2015Fr02).
2024 1	(5)		0.0000.14	Solar neutrino capture rate= $0.015$ SNU 6 (2015Fr02).
2924 1	(5/2)		0.0033 14	$d\sigma/d\Omega(q=0)=0.052$ mb/sr I (2015Fr02).
$325 \times 10^{1}$ 25			0 1 10 70	F(level): 3.0-3.5  MeV energy range
525710 25			0.110 10	Solar neutrino capture rate=0.77 SNU 7 (2015Fr02).
375×10 <sup>1</sup> 25			0.165 16	E(level): 3.5-4.0 MeV energy range.
				Solar neutrino capture rate=0.96 SNU 10 (2015Fr02).
$425 \times 10^{1} 25$			0.191 17	E(level): 4.0-4.5 MeV energy range.
175 101 25			0.000.16	Solar neutrino capture rate=0.92 SNU 9 (2015Fr02).
475×10 <sup>1</sup> 25			0.209 16	E(level): 4.5-5.0 MeV energy range.
$525 \times 10^{1}$ 25			0 101 18	Solar neutrino capture rate= $0.82 \text{ SiNO 7}$ (2013F102). E(level): 5.0.5.5 MeV energy range
525×10 25			0.191 10	Solar neutrino capture rate=0.61 SNU 6 (2015Fr02).
575×10 <sup>1</sup> 25			0.265 19	E(level): 5.5-6.0 MeV energy range.
				Solar neutrino capture rate=0.67 SNU 5 (2015Fr02).
$625 \times 10^1 \ 25$			0.338 18	E(level): 6.0-6.5 MeV energy range.
				Solar neutrino capture rate= $0.66$ SNU 4 (2015Fr02).
$6/5 \times 10^{1}$ 25			0.315 17	E(level): 6.5-7.0 MeV energy range.
$721 \times 10^{1} 21$			0 280 16	Solar neutrino capture rate= $0.47$ SNU 5 (2015FT02). E(level): 7.0-7.42 MeV energy range
721×10 21			0.289 10	Solar neutrino capture rate=0.33 SNU 2 (2015Fr02).
794×10 <sup>1</sup> 52			0.645 39	E(level): 7.42-8.46 MeV energy range.
				B(GT)=0.74 <i>10</i> (1998Ej03).
8913 5	3/2-	≈50 keV		%n≈100 (1988Ch25)
				B(Fermi)= $0.0900\ 22\ (2011Fr15)$ .
				$^{70}$ Ge with subsequent observation of $1040_{\text{V}}$ in $^{70}$ Ge
				The $\gamma$ decay of IAS to <sup>71</sup> Ga through single high-energy $\gamma$ ray or a
				cascade of two $\gamma$ ravs is determined to be less than 11%
				(1988Ch25).
				E(level): from 2011Fr15 and 2015Fr02, with uncertainty estimated as
				5 keV by evaluators based on FWHM=45 keV. Other: 8932 13
				(1700 - 1123). $I^{\pi}$ : IAS of $3/2^{-1}$ as in $71$ Ga
				$\Gamma$ from 2011Fr15.
				$d\sigma/d\Omega(q=0)=9.04$ mb/sr 12 (2011Fr15).
$11.75 \times 10^{3}$				Gamow-Teller giant resonance (GTGR), wide structure from 10-16

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## <sup>71</sup>Ga(<sup>3</sup>He,t),(<sup>3</sup>He,tγ) 2015Fr02,1998Ej03,1988Ch25 (continued)

### <sup>71</sup>Ge Levels (continued)

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E(lawal)	t –
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Comments

 $18.0 \times 10^{3}$ 

MeV, peaking at  $\approx 11.75$  MeV (2015Fr02,2011Fr15). Interpreted as T<sub>></sub> component of the GTGR (2015Fr02,2011Fr15).

<sup>†</sup> From 2015Fr02, unless otherwise noted.

<sup>‡</sup> States populated by Gamow-Teller excitations are expected to be  $1/2^-, 3/2^-$  or  $5/2^-$ .  $J^{\pi}$  assignments are as given in Table I of 2015Fr02, which are based on literature assignments for low-lying levels, and from  $\sigma(\theta)$  data with comparison to DWBA calculations in the higher energy region.

<sup>#</sup> From 2015Fr02.