

$^{60}\text{Ni}(\alpha, \alpha)$  1978Zi01, 1990Se07

Type	Author	History	Citation	Literature Cutoff Date
Full Evaluation	M. Shamsuzzoha Basunia		NDS 151, 1 (2018)	1-Apr-2018

Others: 1965Br29, 1966Fo01, 1966Le12, 1968Ru02, 1969Ar20, 2013Sc06, 2013ScZZ. Also (p,d),  $^{58}\text{Ni}(\alpha, ^3\text{He}), (d,p)$  were studied by 2013Sc06, 2013ScZZ.

1990Se07:  $E(^3\text{He})=33$  MeV; FWHM $\approx$ 60 keV;  $\theta(\text{lab})=12.5^\circ-115^\circ$  in steps of  $2.5^\circ$ ;  $\Delta E$ -E telescopes and particle identification; 98.5%  $^{60}\text{Ni}$  target. DWBA analysis of  $\sigma(\theta)$ .

1978Zi01:  $E(^3\text{He})=38$  MeV; FWHM $\approx$ 40 keV,  $\theta=9^\circ-21^\circ$ , magnetic spectrograph, particle identification. Measured  $\sigma(\theta)$ .

1968Ru02:  $E(^3\text{He})\approx 25$  MeV; 99.8%  $^{60}\text{Ni}$  target, FWHM=100-140 keV,  $\theta(\text{c.m.})\approx 15^\circ-93^\circ$ . DWBA analysis of  $\sigma(\theta)$ . Same authors as 1965Br29.

1966Fo01:  $E(^3\text{He})=15$  MeV; 99.2%  $^{60}\text{Ni}$  target,  $\theta(\text{C.M.})\approx 12^\circ-93^\circ$ , evaluator estimated FWHM $\approx$ 40 keV (not stated by authors), semi detectors. DWBA analysis of  $\sigma(\theta)$ .

1965Br29:  $E(^3\text{He})=24.5$  MeV; 99.8%  $^{60}\text{Ni}$  target, FWHM=100 keV, semi detectors,  $\theta(\text{lab})=15^\circ-90^\circ$  ( $5^\circ$  steps).

 $^{59}\text{Ni}$  Levels

E(level) <sup>†</sup>	$J^\pi$ <sup>‡</sup>	L	$C^2S$ <sup>#</sup>	Comments
0.0	$3/2^-$	1	1.26	
338	$5/2^-$	3	1.66	
469	$1/2^-$	1	0.27	
878	$3/2^-$	1	0.19	
1192	$5/2^-$	3	0.12	
1313	$1/2^-$	1	0.26	
1681	$5/2^-$	3	0.24	
1951	$7/2^-$	3	0.91	
2414				From fig. 1 of 1990Se07 only.
2627	$7/2^-$	3	3.12	E(level): level assumed to be same as 2670 35 level from 1965Br29.
3049	$7/2^-$	3	1.38	E(level): level assumed to be same as 3100 35 level from 1965Br29.
3542				From fig. 1 of 1990Se07 only.
3730 10	$7/2^-$	3	0.34	
4218 10	$7/2^-$	3	1.12	E(level): 4181 in 1990Se07. Apparent doublet exists near 4200 50 in 1966Fo01 (based on shape for 4200-keV peak).
4286 10	$7/2^-$	3@	0.39@	
4386 10	$7/2^-$	3@	0.39@	
4441 10	$7/2^-$	3@	0.38@	
4541 10	$7/2^-$	3@	0.17@	
4639	$7/2^-$	3	0.66	
4709 10	$7/2^-$	3@	0.38@	
5148				E(level): level assumed to be same as 4760 35 level from 1965Br29.
5264	$7/2^-$	3	0.13	From fig. 1 of 1990Se07 only.
5586	$1/2^+, 3/2^+$	0,2	0.46,0.72	Level assumed to be same as 5310 35 level from 1965Br29.
5740 35				E(level): reported only by 1965Br29; peak (in spectrum of fig. 2) is too broad to be attributed to a single level.
5867	$1/2^+, 3/2^+$	0,2	0.34,0.54	
6600 35				E(level): reported only by 1965Br29.
7068	$7/2^-$	3	0.13	
7330 & 50	$7/2^-$	3	1.41	$^{59}\text{Co}$ $7/2^-$ g.s. analogue.
8482	$3/2^-$	1	0.11	Possibly the $^{59}\text{Co}$ 1099-keV, $3/2^-$ analogue (1990Se07).
9081	$7/2^-$	3	0.19	Possibly the $^{59}\text{Co}$ 1745-keV, $7/2^-$ analogue (1990Se07).
10085	$1/2^+, 3/2^+$	0,2	0.26,0.60	Possibly the $^{59}\text{Co}$ 2713-keV, $1/2^+$ analogue (1990Se07).
10527	$1/2^+, 3/2^+$	0,2	0.21,0.58	Possibly the $^{59}\text{Co}$ 3162-keV, $3/2^+$ analogue (1990Se07).

<sup>†</sup> From 1978Zi01 if  $\Delta E=10$  keV, from 1965Br29 if  $\Delta E=35$  keV, from 1966Fo01 if  $\Delta E=50$  keV and from 1990Se07 if  $\Delta E$

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 $^{60}\text{Ni}(^3\text{He},\alpha)$  [1978Zi01,1990Se07](#) (continued) $^{59}\text{Ni}$  Levels (continued)

unstated. E from [1990Se07](#) deviates from adopted values by  $\leq 4$  keV for  $E < 3000$ .

‡ Assumed by authors for calculation of  $C^2S$ .

# From comparison of  $\sigma(\theta)$  with finite range DWBA calculations ([1990Se07](#)), except as noted.  $C^2S$  normalized so sum for  $2p_{1/2}$ ,  $2p_{3/2}$  and  $1f_{5/2}$  transfer to  $E(\text{level}) \leq 1700$  keV equals 4.

@ From [1978Zi01](#). Note that  $C^2S$  from [1978Zi01](#) and [1990Se07](#), in general, differ significantly.

&  $E=7332$  in [1990Se07](#). The  $E=7520$  35 level of [1965Br29](#) may well be the same as this level since the authors' stated E deviates significantly from an otherwise linear energy calibration for the spectrum in their fig. 2.