

Adopted Levels, Gammas

| Type            | Author     | History Citation   | Literature Cutoff Date |
|-----------------|------------|--------------------|------------------------|
| Full Evaluation | M. R. Bhat | NDS 85, 415 (1998) | 24-Sep-1998            |

$Q(\beta^-) = -1.476 \times 10^4$  SY;  $S(n) = 1.714 \times 10^4$  SY;  $S(p) = 690.3$  4;  $Q(\alpha) = -7074.5$  18    [2012Wa38](#)

Note: Current evaluation has used the following Q record.

$\Delta Q(\beta^-)$ : = 140 keV from systematics.

$\Delta S(n)$ : = 140 keV from systematics.

$Q(\beta^-) = -14.62 \times 10^3$  SY;  $S(n) = 16.78 \times 10^3$  SY;  $S(p) = 695$  19;  $Q(\alpha) = -7091$  24    [1995Au04](#)

[1998Re01](#): calculated reaction rates for  $^{57}\text{Ni}(p,\gamma)$  in stellar interiors from the spectroscopic factors of the low-lying states in  $^{57}\text{Ni}$  using charge symmetry.

 $^{57}\text{Cu}$  Levels

$J^\pi, T$ : see  $^{57}\text{Zn} \beta^+$  decay? and ( $^7\text{Li}, ^8\text{He}$ ) for other proposed assignments.

Cross Reference (XREF) Flags

- A  $^{58}\text{Ni} (^7\text{Li}, ^8\text{He}), (^{14}\text{N}, ^{15}\text{C})$
- B  $^{57}\text{Zn} \beta^+$  decay:?
- C  $^1\text{H} (^{58}\text{Ni}, ^{57}\text{Cu} \gamma)$

| E(level) <sup>†</sup> | $J^\pi$          | $T_{1/2}$  | XREF | Comments  |
|-----------------------|------------------|------------|------|---|
| 0.0                   | $3/2^-$          | 196.3 ms 7 | ABC  | $\% \epsilon + \% \beta^+ = 100$<br>$T_z = -1/2$<br>J, ISPINZ super-allowed decay to $J^\pi = 3/2^-$ , $T_z = +1/2$ , g.s. of $^{57}\text{Ni}$ ( <a href="#">1984Sh28</a> ).<br>$T_{1/2}$ : from <a href="#">1996Se01</a> . Others: 199.4 ms 32 ( <a href="#">1987HaZN</a> ), 223 ms 16 ( <a href="#">1984Sh28</a> ). |
| 1028 4                | $5/2^- \ddagger$ |            | A C  |   |
| 1106 4                | $1/2^- \ddagger$ |            | C    |   |
| 2398 10               | $5/2^- \ddagger$ |            | C    |   |
| 2520 25               |                  |            | A    |   |
| $3.28 \times 10^3$ 5  |                  |            | B    | $\% p = 100$  |
| 3510 25               |                  |            | A    |   |
| $5.35 \times 10^3$ 5  |                  |            | B    | $\% p = 100$  |
| 5710 25               |                  |            | A    |   |

<sup>†</sup> Levels below 2500 keV are from ( $^{58}\text{Ni}, ^{57}\text{Cu} \gamma$ ); the rest are from ( $^7\text{Li}, ^8\text{He}$ ) or  $^{57}\text{Zn} \beta^+$  decay.

<sup>‡</sup> By comparison with  $J^\pi$  assignments in the mirror nuclide  $^{57}\text{Ni}$  ([1996Zh02](#)).

 $\gamma(^{57}\text{Cu})$ 

| $E_i(\text{level})$ | $J_i^\pi$ | $E_\gamma$ | $I_\gamma$ | $E_f$ | $J_f^\pi$ |
|---------------------|-----------|------------|------------|-------|-----------|
| 1028                | $5/2^-$   | 1028 4     | 100        | 0.0   | $3/2^-$   |
| 1106                | $1/2^-$   | 1106 4     | 100        | 0.0   | $3/2^-$   |
| 2398                | $5/2^-$   | 2398 10    | 100        | 0.0   | $3/2^-$   |

**Adopted Levels, Gammas****Level Scheme**

Intensities: Relative photon branching from each level

