### <sup>190</sup>W IT decay (111 ns) **2010La16,2011St21**

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Parent:  $^{190}$ W: E=1743.6 10;  $J^{\pi}$ =(8<sup>+</sup>);  $T_{1/2}$ =111 ns 17; %IT decay=100.0

2010La16: <sup>190</sup>W ions were produced by E=840 MeV <sup>136</sup>Xe beam from the ATLAS accelerator at ANL, bombarding a 20 mg/cm<sup>2</sup> <sup>192</sup>Os target. Measured Εγ, Ιγ, γγ, delayed γ spectra using Gammasphere array at ATLAS-ANL facility. In reference 24, authors mentioned that full details of this study are to be published elsewhere. But in evaluators' search of literature, no further publication from this group seems to have appeared.

2011St21 (also 2012Al05, 2010FaZX, 2009Al30, 2009Fa06, 2009Al16, 2008StZY, 2005Ca02, 2003Po14, 2001Ca13, 2000Po26, 2000Sc32, 2000Wo13):  $^{190}$ W nuclide formed by in-flight fragmentation of  $^{208}$ Pb beam at 1 GeV/nucleon from the GSI UNILAC and SIS-18 accelerator complex. Target thickness of 2.526 g/cm², backed by  $^{93}$ Nb foil of thickness 0.223 g/cm². Fragments separated and identified in flight by the Fragment Separator (FRS) operated in achromatic mode, based on time of flight, B $\rho$  and energy loss. Transmitted ions slowed in Al degraders and stopped in a plastic catcher. The stopper was surrounded by the RISING  $\gamma$ -ray spectrometer. Measured E $\gamma$ , I $\gamma$ , delayed  $\gamma$  rays, isomer lifetime.  $^{190}$ W beam was fully-stripped or mixture of H- or He-like nuclei.

#### Additional information 1.

The level scheme is essentially from 2010La16. 2011St21 provide energies of four gamma rays from the decay of  $108-\mu s$  isomer.

### <sup>190</sup>W Levels

E(level) <sup>†</sup>	$J^{\pi \ddagger}$	$T_{1/2}$	Comments				
0.0#	$0^{+}$						
206.8 <sup>#</sup> 5	$(2^{+})$						
565.1 <sup>#</sup> 7	$(4^{+})$						
1049.4 <sup>#</sup> 9	$(6^{+})$						
1642.2 <sup>#</sup> <i>12</i>	(8+)		Transition intensity balance at this level is 25% 28, in contrast to expected zero balance. Note that the uncertainty is large for the 593.6 $\gamma$ . Or it is possible that the photon intensity of the $102\gamma$ is underestimated.				
			E(level): level from 2010La16 only, and proposed as member of the g.s. band, as in 2005Ca02 and 2000Po26. This level and feeding $\gamma$ of 102 keV are not reported by 2011St21.				
1743.6 <i>10</i>	(8+)	111 ns <i>17</i>	%IT=100 $T_{1/2}$ : from $\gamma(t)$ with gates on the prompt 324 $\gamma$ and 356 $\gamma$ feeding this level (2010La16). Configuration= $\nu$ 9/2[505] $\otimes \nu$ 7/2[503] (2010La16,2009Fa06).				

<sup>&</sup>lt;sup>†</sup> From E $\gamma$  data, assuming 1 keV for 102 $\gamma$ .

$$\gamma$$
(190W)

<sup>&</sup>lt;sup>‡</sup> From the Adopted Levels, essentially, based on data and interpretation from 2010La16.

<sup>#</sup> Band(A): g.s. band.

### <sup>190</sup>W IT decay (111 ns) **2010La16,2011St21** (continued)

# $\gamma$ <sup>(190</sup>W) (continued)

$E_{\gamma}^{\dagger}$	$I_{\gamma}^{\#}$	$E_i(level)$	$\mathbf{J}_i^{\pi}$	$\mathrm{E}_f \qquad \mathrm{J}_f^\pi$	Mult.	$\alpha^{@}$	$I_{(\gamma+ce)}^{\#}$	Comments
206.8 5	78.4 <sup>‡</sup> 3	206.8	(2+)	0.0 0+	[E2]	0.275	100	be more likely for (8 <sup>+</sup> ) to (8 <sup>+</sup> ) transition. Reduced hindrance factor $f_v$ =5.4 2, $v$ =7, from B(M1)(W.u.)= $7 \times 10^{-6}$ 2, as given in 2010La16. $\alpha(K)$ =0.1540 23; $\alpha(L)$ =0.0921 14; $\alpha(M)$ =0.0229 4 $\alpha(N)$ =0.00541 9; $\alpha(O)$ =0.000765 12; $\alpha(P)$ =1.228×10 <sup>-5</sup> 18 E <sub><math>\gamma</math></sub> : others: 207.03 (2005Ca02), 206
358.3 5	95.15 <sup>‡</sup> 7	565.1	$(4^+)$	206.8 (2+)	[E2]	0.0510	100	(2010La16), 207 (2000Po26). I <sub>γ</sub> : others: 81 10 (2011St21), 129 33 (renormalized from 27 7 in 2005Ca02), 105 24 (renormalized from 65 15 2000Po26). α(K)=0.0362 6; α(L)=0.01134 17;
								$\alpha(M)=0.00274 \ 4$ $\alpha(N)=0.000652 \ 10; \ \alpha(O)=9.63\times10^{-5} \ 14;$ $\alpha(P)=3.19\times10^{-6} \ 5$ $E_{\gamma}$ : others: 357.4 3 (2005Ca02), 358 (2010La16), 357 (2000Po26).
								I <sub>γ</sub> : others: 92 <i>10</i> (2011St21), 114 <i>33</i> (renormalized from 24 <i>7</i> in 2005Ca02), 107 <i>24</i> (renormalized from 67 <i>15</i> 2000Po26).
484.3 5	97.77 <sup>‡</sup> 3	1049.4	(6 <sup>+</sup> )	565.1 (4+)	[E2]	0.0228	100	$\alpha(K)$ =0.01729 25; $\alpha(L)$ =0.00419 6; $\alpha(M)$ =0.000996 15 $\alpha(N)$ =0.000238 4; $\alpha(O)$ =3.61×10 <sup>-5</sup> 6; $\alpha(P)$ =1.575×10 <sup>-6</sup> 23 $E_{\gamma}$ : others: 484.0 4 (2005Ca02), 484 (2010La16), 485 (2000Po26). $I_{\gamma}$ : others: 100 11 (2011St21), (100 33 (renormalized from 21 7 in 2005Ca02), 100 32 (renormalized from 62 20
593.6 11	20.6 38	1642.2	(8+)	1049.4 (6 <sup>+</sup> )	[E2]	0.01387		2000Po26). $\alpha(K)$ =0.01088 $16$ ; $\alpha(L)$ =0.00230 $4$ ; $\alpha(M)$ =0.000541 $9$ $\alpha(N)$ =0.0001293 $20$ ; $\alpha(O)$ =2.00×10 <sup>-5</sup> $3$ ; $\alpha(P)$ =1.003×10 <sup>-6</sup> $15$ $E_{\gamma}$ : from 2005Ca02. Others: E $\gamma$ =592 (2010La16), 591 (2000Po26). This $\gamma$ was not seen by 2009Al30, and not reported by 2011St21, perhaps due to its weak population. $I_{\gamma}$ : deduced by evaluators from intensity balance at 1642 level. Others: 43 $27$ (2000Po26), renormalized from 27 $17$ , 100 $38$ , renormalized from 21 $8$ in 2005Ca02. This $\gamma$ is as strong as the 484.1 $\gamma$ with $I_{\gamma}$ =21 $7$ in Table 1 of 2005Ca02, but seen in the $\gamma$ spectrum in Fig. 4 of 2005Ca02, much weaker
694.0 5	78.5 37	1743.6	(8+)	1049.4 (6+)	[E2]	0.00972		than the 484.1 $\gamma$ peak. $\alpha(K)$ =0.00776 11; $\alpha(L)$ =0.001509 22; $\alpha(M)$ =0.000351 5

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#### $\gamma$ (190W) (continued)

 $E_{\gamma}^{\dagger}$   $E_{i}(level)$ 

Comments

 $\alpha({\rm N}){=}8.41{\times}10^{-5}~12;~\alpha({\rm O}){=}1.313{\times}10^{-5}~19;~\alpha({\rm P}){=}7.19{\times}10^{-7}~11$ 

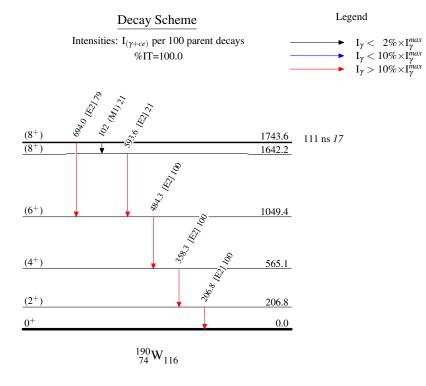
 $E_{\gamma}$ : others: 694 (2010La16), 694.0 7 (2005Ca02), 695 (2000Po26); in 2005Ca02 and 2000Po26, this  $\gamma$  was placed from a 2335,  $10^+$  level, as member of g.s. band). The decay scheme was modified by 2009A130 and 2009Fa06 to place 694 $\gamma$  from this level, and omitting the 594 $\gamma$  from 1640 level as the 594 $\gamma$  was absent in the delayed  $\gamma$  spectrum obtained by 2009A130.

delayed  $\gamma$  spectrum obtained by 2009Al30.  $I_{\gamma}$ : from <sup>190</sup>W IT decay (166  $\mu$ s), based on data in 2010La16. Others: 71 *10* (2011St21), 76 *33* (renormalized from 16 7 in 2005Ca02), 64 *24* (renormalized from 62 *20* 2000Po26).

Reduced hindrance factor  $f_v=3.7$  1, v=6, from B(E2)(W.u.)= $3.9\times10^{-4}$  6, as given in 2010La16.

- <sup>‡</sup> From  $I(\gamma+ce)=100$  and  $\alpha$ (theory) for E2.
- # Absolute intensity per 100 decays.
- <sup>@</sup> Total theoretical internal conversion coefficients, calculated using the BrIcc code (2008Ki07) with Frozen orbital approximation based on  $\gamma$ -ray energies, assigned multipolarities, and mixing ratios, unless otherwise specified.

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<sup>&</sup>lt;sup>†</sup> From 2011St21, unless otherwise stated. Uncertainty of 0.5 keV for Eγ values in 2011St21 are assigned in consultation with Zs. Podolyak (co-author of 2011St21) in January 2012.

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Band(A): g.s. band

