## <sup>180</sup>**Pb** *α* decay **1999To11,1996To08**

		History	
Туре	Author	Citation	Literature Cutoff Date
Full Evaluation	M. S. Basunia	NDS 107, 791 (2006)	15-Sep-2005

Parent: <sup>180</sup>Pb: E=0.0; J<sup> $\pi$ </sup>=0<sup>+</sup>; T<sub>1/2</sub>=4.4 ms *11*; Q( $\alpha$ )=7415 *15*; % $\alpha$  decay=100.0 T<sub>1/2</sub>(<sup>180</sup>Pb) from 1999To11.

1999To11:  ${}^{90}$ Zr( ${}^{92}$ Mo,2n), E=410 MeV; Detector: double sided Si strip detector; measured E $\alpha$ , T<sub>1/2</sub>; deduced  $\alpha$ -branching ratio. 1996To08:  ${}^{144}$ Sm( ${}^{40}$ Ca,4n),E=230 MeV; detector: an array of six Si detector; measured E $\alpha$ , I $\alpha$ ; deduced evidence for  ${}^{180}$ Pb.

 $\alpha/\beta$  branchings have not been experimentally determined. The gross- $\beta$  calculations of 1973Ta30 yield  $T_{1/2}(\beta^+)=1-4$  s from which

 $\beta$  branching can be calculated as 0.1-0.4%. Any  $\beta$  branching is taken as negligible here and  $\%\alpha$ =100 is adopted. Additional information 1.

Added-in-Proof: 1997Mo25 calculated the  $\beta$  partial half-life as T<sub>1/2</sub>( $\beta$ )=618.8 ms which corresponds to a  $\beta$  branching of 0.65%.

## <sup>176</sup>Hg Levels

E(level)	$J^{\pi}$
0.0	$0^{+}$

 $\alpha$  radiations

Eα	E(level)	HF <sup>†</sup>	Comments
7250 15	0.0	1.0	<ul> <li>Eα: From 1999To11. Other value: 7230 40 (1996To08).</li> <li>Only one α group could be observed; due to the small production cross section and short half-life. Even the main α group was very weak in the spectrum. Therefore, any α less than half of the main α could not have been seen; besides, any α's to excited states would be obscured in the spectrum of 1996To08 by strong α's from other nuclei.</li> <li>Iα: the first 2<sup>+</sup> state in <sup>176</sup>Hg is at 613 keV. By requiring the hindrance factor to be &gt;1 for an α transition to this level, its intensity should be Iα&lt;1%. The calculated r<sub>0</sub> parameter retains the same value for Iα(7250)&gt;96%. The computed r<sub>0</sub> parameter of 1.53 4 is consistent with the expected value of 1.52 2 from the r<sub>0</sub> systematics.</li> </ul>

<sup>†</sup>  $r_0(^{176}\text{Hg})=1.53 \ 4 \ (1998\text{Ak04}).$