

^{178}Pt α decay 1998Ak04

Type	Author	History	Citation	Literature Cutoff Date
Full Evaluation	E. Browne, Huo Junde		NDS 87, 15 (1999)	1-Nov-1998

Parent: ^{178}Pt : $E=0.0$; $J^\pi=0^+$; $T_{1/2}=21.1$ s 6; $Q(\alpha)=5573$ 3; $\% \alpha$ decay=4.6 31

The measured half-lives are 22 s 2 (1993Me13), 21 s 1 (1982Bo04), 19 s 2 (1970Ha18), 21.2 s 8 (1968De01) and 21.3 s 15 (1966Si08). $T_{1/2}(^{178}\text{Pt})=21.1$ s 6, the weighted average of the measured values is adopted in 1994Br18, and used here.

The intensities of $I_\alpha(5446\alpha)=1.3\%$ and $I_\alpha(5284\alpha)=0.07\%$ were obtained by 1966Si08 from the relative α intensities and the relative ($^{16}\text{O},\text{xn}$) production cross-sections of platinum isotopes for ^{168}Yb - ^{174}Yb enriched targets. It was assumed that the ratios of maximum cross sections for $x=6,7$, and 8 are the same for all ytterbium targets.

The α intensities of $I_\alpha(5446\alpha)=7.2\%$ 8 and $I_\alpha(5284\alpha)=0.20\%$ 7, reported by 1970Ha18, give α -decay branching=7.4% 9. The isotope was produced in the ^{182}Hg α decay, and α branching was obtained by comparing parent and daughter α activities. The 5284-keV α peak was obscured in the singles spectrum by ^{210}Po contamination; its assignment was based on α - γ coincidences. The reported $I_\alpha(5446\alpha)/I_\alpha(5284\alpha)$ ratio was obtained by assuming 10% of the 5284-keV α peak was from ^{182}Au , produced in the β decay of ^{182}Hg .

The α branching was reported by 1980Sc09 as 7.5% 3, presumably for the main α group. In this case, the 5284 α intensity should be added. The isotope ^{178}Pt was produced by 1980Sc09 in ^{186}Pb α decay. The α branching of ^{178}Pt was determined from comparison of intensities of parent and daughter activities in the same spectrum. The considered activities of ^{178}Ir and ^{174}Os (β and α daughters of ^{178}Pt) and their intensities are not given in 1980Sc09.

$\% \alpha=7.7$ 3 is adopted in 1994Br18 from $I_\alpha(5440\alpha)=7.5\%$ 3 (1980Sc09) and $I_\alpha(5284\alpha)/I_\alpha(5440\alpha)=0.028$ 10 (1970Ha18). Adoption of the ratio $I_\alpha(5284\alpha)/I_\alpha(5440\alpha)=0.07/1.3=0.054$ (1966Si08) would give $\% \alpha=7.9$ 4.

The calculated radius parameters are $r_0=1.570$ 3 and $r_0=1.484$, for $\% \alpha=7.7$ 3 and $\% \alpha=1.4$, respectively. The r_0 systematics suggests $r_0=1.540$ from which the partial α half-life of the 5446 α is calculated as 4.6×10^2 s, yielding an absolute value of $I_\alpha(5446\alpha)=4.6\%$ and therefore, the α branching of $\% \alpha=4.8$.

Based on the uncertainties of the experimental branchings due to the assumptions had been made and based on the r_0 systematics, $\% \alpha=4.6$ 31, an average value with a large uncertainty in order to cover all values is suggested here.

$Q(\alpha)(^{178}\text{Pt})=5573.4$ 26 is from 1995Au04. $E_\alpha=5446$ 3 gives $Q(\alpha)=5571$ 3.

 ^{174}Os Levels

E(level)	J^π
0.0	0^+
158.7 3	2^+

 α radiations

E_α	E(level)	$I_\alpha^{\dagger\#}$	HF^{\ddagger}	Comments
5291 4	158.7	5.1 24	3.3 16	E_α : calculated from $E(\text{level})=158.7$ 3 and $E_\alpha(\text{g.s.})=5446$ 3.
5446 3	0.0	94.9 24	1.0	E_α : the measured α energies (adjusted, as recommended by 1991Ry01, for the changes in the calibration energies) are 5440 10 (1966Si08), 5443 10 (1978Fa16), 5458 5 (1979Ha10), 5445 3 (1982Bo04). $E_\alpha=5446$ 3 is recommended by 1991Ry01.

† α intensity per 100 α decays. Relative α intensities of $I_\alpha(\text{to } 2^+)/I_\alpha(\text{to } 0^+)=7/130=0.054$, measured by 1966Si08, is recommended here, since the α peaks in the spectrum of 1966Si08 were not obscured. $I_\alpha(\text{to } 2^+)/I_\alpha(\text{to } 0^+)=0.028$ 10 was measured by 1970Ha18. This ratio gives $I_\alpha(5446\alpha)=97.3$, $I_\alpha(5284\alpha)=2.7$ per 100 α decays. Uncertainties on I_α (no ΔI_α 's are reported in 1966Si08) are assigned here such that I_α 's of 1970Ha18 will be covered.

‡ $r_0(^{174}\text{Os})=1.54$ 3 is calculated by requiring $\text{Hf}(5446\alpha)=1.0$.

$^{\#}$ For absolute intensity per 100 decays, multiply by 0.046 31.