

$^{13}\text{C}(\alpha,\text{n})$ 

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For analyses and measurements of astrophysical S-factors and ANC<sup>2</sup> values see ([1968Da05](#),[2006Jo11](#),[2008Pe09](#),[2014LaZU](#)).

- 1954Tr09:**  $^{13}\text{C}(\alpha,\text{n})$ , E=1.0-3.5 MeV; measured reaction products,  $E_n$ ,  $I_n$ ; deduced neutron yields; calculated energy levels.
- 1956Bo61:**  $^{13}\text{C}(\alpha,\text{n})$ , E=1.8-5.3 MeV; cross sections and widths of the resonances were determined.
- 1957Wa46:**  $^{13}\text{C}(\alpha,\text{n})$ ; measured products; deduced  $\sigma$ ,  $\sigma(E)$ , resonance parameters.
- 1963Sp02:**  $^{13}\text{C}(\alpha,\text{ny})$ , E=5-10 MeV; the excitation function for 6 and 7 MeV gamma radiation from the reaction  $^{13}\text{C}(\alpha,\text{ny})$  has been studied at intervals of approximately 10 keV for  $E_\alpha=5$  to 10 MeV.
- 1965Ba32:** Cross sections for the reaction  $^{13}\text{C}(\alpha,\alpha)$  at  $\theta_{cm}=54.7^\circ$ ,  $107.9^\circ$ ,  $142.6^\circ$ ,  $169.6^\circ$  and for the reaction  $^{13}\text{C}(\alpha,\text{n})^{16}\text{O}$  at  $\theta_{cm}=0^\circ$  were measured. A beam of  $E(\alpha)=2$ -3.5 MeV from the 5.5-MeV Van de Graaff accelerator bombarded a self-supporting foils made either from 41.6%  $^{13}\text{C}$ -enriched methyl iodide, or from 56.7%  $^{13}\text{C}$ -enriched methane with thickness  $\approx 15 \mu\text{g}/\text{cm}^2$ . Using dispersion-theory analysis, a consistent set of  $J^\pi$  and partial-width values for 11 excitation energies  $E_x=8$ -9 MeV were obtained. See also ([1965BaZY](#)).
- 1967Se07:**  $^{13}\text{C}(\alpha,\text{n})$ , E=1.95-5.57 MeV; measured total cross section. Deduced  $\Gamma_a$ ,  $\Gamma_n$  along with the reduced widths  $\gamma_a^2$  and  $\gamma_n^2$  for the levels corresponding to the 2.68, 2.81, 3.72, and 4.62-MeV resonances.
- 1968Ke02:** Cross sections of reactions  $^{13}\text{C}(\alpha,\alpha_0)$  and  $^{13}\text{C}(\alpha,\text{n})$  were measured by bombardment of an  $E_\alpha=12$  MeV beam on to self-supporting, 20-30  $\mu\text{g}/\text{cm}^2$  thick, enriched  $^{13}\text{C}$  targets at the Van de Graaff facility/Australian National University. Two surface-barrier detectors (for  $(\alpha,\alpha_0)$ ) and two 2.5 cm×5 cm plastic scintillators (for  $(\alpha,\text{n})$ ) were used to detect particles. Using a dispersion-theory analysis, the  $J^\pi$  and partial width values were obtained for 11 states of  $^{17}\text{O}$  with  $E_x=9$ -10 MeV.
- 1969Sc04:**  $^{13}\text{C}(\alpha,\text{n})$ , E=1.38-2.26 MeV; measured angular distribution of n-polarization.  $^{17}\text{O}$  deduced resonances,  $J$ ,  $\pi$ , level-width.
- 1970Ro08:** Thin foil targets of 50.5% enriched  $^{13}\text{C}$ , 30.3  $\mu\text{g}/\text{cm}^2$  17 thick, were bombarded with  $E(\alpha)=4.5$ -10.5 MeV ion beams produced from the University of Virginia 5.5 MeV Van de Graaff accelerator. Neutrons were detected using a 2.5 cm×2.5 cm long stilbene crystal and a 2.5 cm×5.0 cm long Ne 213 liquid scintillator with detection efficiencies of  $\pm 6\%$ . The energy scale for the excitation function was calibrated with an uncertainty of  $\pm 10$  keV. The excitation function was measured at  $\theta=0^\circ$  and angular distributions were measured at  $\theta=0^\circ$ - $170^\circ$ . Energy levels of  $^{17}\text{O}$  at  $E_x=10$ -13 MeV with  $J^\pi$  values up to 9/2 were deduced.
- 1971Ba06:**  $^{13}\text{C}(\alpha,\text{n})$ , E=3.36-4.80 MeV; measured  $\sigma(E; \theta)$ ,  $P(n)(E; \theta)$ .  $^{17}\text{O}$  resonances deduced  $J$ ,  $\pi$ .
- 1973Ba10:**  $^{13}\text{C}(\alpha,\text{n})$ , E=1-5 MeV; measured  $\sigma(E)$ ;  $^{17}\text{O}$  deduced resonances, level-width.
- 1973Bu14:**  $^{13}\text{C}(\alpha,\text{n})$ , E=2.075,2.25,2.43 MeV; measured n-polarization( $\theta$ ),  $\sigma(\theta)$ .  $^{17}\text{O}$  levels deduced  $J$ ,  $\pi$ .
- 1973Lo16:**  $^{13}\text{C}(\alpha,\text{n})$ , measured  $E_n$ ,  $I_n$ .
- 1975Be44:**  $^{13}\text{C}(\alpha,\text{ny})$ , measured  $\sigma(E,E_\gamma)$ .  $^{17}\text{O}$  deduced resonances,  $\Gamma$ .
- 1976Mc11:**  $^{13}\text{C}(\alpha,\text{n})$ , E=4.2-8.7 MeV; measured  $\sigma(E,E_n,\theta)$ .  $^{17}\text{O}$  deduced T=3/2 levels,  $\Gamma$ . Enriched target.
- 1976Ra36:**  $^{13}\text{C}(\alpha,\text{n})$ , E=0.60-1.15 MeV; measured  $\sigma(E)$ ; deduced astrophysical  $\sigma$  factors.  $^{17}\text{O}$  1.056-MeV resonance deduced parameters. Enriched target.
- 1981HaZV:**  $^{13}\text{C}(\alpha,\text{n})$ , E≈35 MeV; measured  $\sigma(En,\theta)$ . Cluster transfer, DWBA analysis. Tof, solid state counters, magnetic spectrograph.
- 1982CrZY:**  $^{13}\text{C}(\alpha,\text{n})$ , E=35 MeV; measured  $\sigma(E_n)$ ,  $\sigma(\theta)$ . DWBA, cluster transfer, knockout mechanisms.
- 1983HaZX:**  $^{13}\text{C}(\alpha,\text{n})$ , E=35 MeV; measured  $\sigma(\theta)$ . DWBA, three nucleon stripping, semimicroscopic, cluster models.
- 1990We10:**  $^{13}\text{C}(\alpha,\text{n})$ , E=2.406-3.308 MeV; measured  $\sigma(\theta)$ , polarization. New design high pressure  $^4\text{He}$ -polarimeter.
- 1993Br17:**  $^{13}\text{C}(\alpha,\text{n})$ , E=650-1600 KeV; measured yield vs E.  $^{17}\text{O}$  deduced resonances,  $\Gamma_n$ ,  $\Gamma_\alpha$ , resonance strengths.
- 1993Dr08:**  $^{13}\text{C}(\alpha,\text{n})$ , E(cm)≈275-1075 keV; measured  $\sigma(E)$ ; deduced astrophysical S-factor, reaction rates at helium burning temperatures.
- 1993DrZZ:** E=0.35-1.4 MeV; measured  $E_n$ ,  $\sigma$ , resonance energy, width, strength, S-factor, reaction rate. Comparison with other measurements.
- 2001He22:**  $^{13}\text{C}(\alpha,\alpha)$ , E=2.6-6.2 MeV; measured  $\sigma(\theta)$ .  $^{13}\text{C}(\alpha,\text{n})$ , E=0-2 MeV; calculated  $\sigma$ , S-factor following r-matrix analysis of elastic scattering data. Comparison with earlier results for s-process conditions.
- 2003Ka51:**  $^{13}\text{C}(\alpha,\text{n})$ , E(cm)≈200-800 keV; deduced astrophysical S-factors, reaction rate.
- 2003Ku36:**  $^{13}\text{C}(\alpha,\text{n})$ , E(cm)=0-800 keV; deduced astrophysical S-factors, reaction rates.
- 2005Ha69:**  $^{13}\text{C}(\alpha,\text{n})$ , E=0.8-8.0 MeV; measured  $\sigma$ , neutron yields.
- 2007PeZZ:**  $^{13}\text{C}(\alpha,\text{n})$ , deduced S $_\alpha$  factor.
- 2008He11:**  $^{13}\text{C}(\alpha,\text{n})$ , E(cm)=320-700 keV;  $^{13}\text{C}(\alpha,\alpha)$ , E=2.6-6.2 MeV; measured radii,  $\sigma$ ,  $\sigma(\theta)$ , S-factor.  $^{17}\text{O}$  deduced levels,  $J$ ,  $\pi$ ,

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### $^{13}\text{C}(\alpha,\text{n})$ (continued)

resonance parameters.

**2009Ma70:**  $^{13}\text{C}(\alpha,\text{n}), (\alpha,\gamma)$ , E=2.000,2.270 MeV; measured  $E\gamma$ ,  $I\gamma$ ,  $\gamma(\theta)$ ,  $E_n$ ,  $\sigma$ , and  $\sigma(\theta)$ ; deduced astrophysical S factors.

Comparison with previous experimental data.

**2018Sm01:**  $^{13}\text{C}(\alpha,\text{n})$ , E not given; measured reaction products,  $E\gamma$ ,  $I\gamma$ ,  $E_n$ ,  $I_n$ ; deduced yields.

**2020Fe06:**  $^{13}\text{C}(\alpha,\text{n}_{0,1,2})$ , E=5.2-6.4 MeV; analyzed excitations functions and impact on backgrounds for large volume  $\nu$  detectors.

**Theory:**

**1977Li19:**  $^{13}\text{C}(\alpha,\text{n})$ , E<7 MeV; analyzed  $\sigma(E)$ .

**1987De38:**  $^{13}\text{C}(\alpha,\text{n}), (\alpha,\alpha)$ , E=low; calculated transfer, elastic  $\sigma(\theta)$ ,  $\sigma(E)$ . Generator coordinate method.

**1996Le06:**  $^{17}\text{O}$ ; calculated levels using parameters for  $^{13}\text{C}+\alpha$  cluster system. Semi-microscopic algebraic cluster model.

**2003Ku03:**  $^{13}\text{C}(\alpha,\text{n})$ , E(cm)=0-800 keV; calculated astrophysical S-factors, reaction rates.

**2008St11:**  $^{13}\text{C}(\alpha,\text{n})$ , analyzed reaction rates.

**1997Ha37:**  $^{13}\text{C}(\alpha,\text{n})$ , E not given; analyzed reaction  $\sigma$  data; deduced astrophysical S-factor vs E, extrapolation methods accuracy. Astrophysical implications.

**1999An35:**  $^{13}\text{C}(\alpha,\text{n})$ , E<10 MeV; compiled, analyzed  $\sigma$ , S-factors; calculated astrophysical reaction rates vs  $T_9$ . Analytical approximations.

**2001Du11:**  $^{13}\text{C}(\alpha,\text{n})$ , E(cm)=0.1-1 MeV; calculated  $\sigma$ , S-factor. Comparison with data, Astrophysics interest.

**2001Du12:**  $^{13}\text{C}(\alpha,\text{n})$ , E(cm)≈0.2-2 MeV; calculated S-factors. Comparisons with data.

**2005Ad03:**  $^{13}\text{C}(\alpha,\text{n})$ , E=0-5 MeV; calculated phase shifts, transition amplitudes. Comparison of DWBA and microscopic cluster model.

**2005Du20:**  $^{13}\text{C}(\alpha,\text{n})$ , E(cm)≈0.1-1.2 MeV; calculated S-factors.  $^{17}\text{O}$ ; calculated levels, J,  $\pi$ . Generator coordinate method, comparison with data.

**2005Pi19:**  $^{13}\text{C}(\alpha,\text{n})$ , E=low; analyzed astrophysical reaction rates.

**2007Mu10:**  $^{13}\text{C}(\alpha,\text{n})$ , E=0-0.9 MeV; calculated astrophysical S-factor. Asymptotic normalization coefficient method. Comparison with data. **2008KaZX:**  $^{13}\text{C}(\alpha,\text{n})$ , E≈0.1-10 MeV; analyzed S-factors.

**2008Lu01:**  $^{13}\text{C}(\alpha,\text{n})$ , E not given; analyzed reaction rates as neutron sources for s process in AGB stars.

**2012Mi24:**  $^{13}\text{C}(\alpha,\text{n})$ , E<1 MeV; analyzed available data; calculated reaction rates, isotope abundances. Comparison with available data.

**2014Ku13:**  $^{13}\text{C}(\alpha,\text{n})$ , E=0.7-4.7 MeV; calculated  $\sigma$  using multi-channel R-matrix with care for covariances; deduced resonances. Compared to ENDF/B-VII.1 and Harisopoulos data.

**2015LaZW:**  $^{13}\text{C}(\alpha,\text{n})$ , E(cm)=0-1.2 MeV; calculated S-factor using R-matrix. Compared to data.

**2015Vi01:**  $^{13}\text{C}(\alpha,\text{n})$ , E=4-9 MeV; analyzed available data; deduced thick target yields and their uncertainties.

**2016La06:**  $^{13}\text{C}(\alpha,\text{n})$ , E(cm)<1 MeV; calculated S-factors. R-matrix, Trojan horse method (THM) resonance parameters.

**2016Sp03:**  $^{13}\text{C}(\alpha,\text{n})$ , E(cm)=0-1.2 MeV; compiled S-factor data, fitting THM (Trojan Horse Method) and calculations.

**2017HaZY:**  $^{13}\text{C}(\alpha,\text{n})$ , E=0-5.4 MeV ; calculated  $\sigma$ .  $^{13}\text{C}(\alpha,\alpha)$ , E=2-5.7 MeV; calculated  $\sigma(\theta)$ .

**2017Mu14:**  $^{13}\text{C}(\alpha,\text{n})$ , E(cm)<1.1 MeV; calculated astrophysical S factor as a function of the  $\alpha$ - $^{13}\text{C}$  relative kinetic energy using the R-matrix approach for resonances and Trojan horse method (THM) for S factor. Comparison with experimental data. Relevance to neutron generation in low-mass AGB stars.

**2017Pa45:**  $^{13}\text{C}(\alpha,\text{n})$ , E not given; analyzed available data; deduced isotopic abundance ratios of s-elements in presolar SiC grains.

**2017Pe13:**  $^{13}\text{C}(\alpha,\text{n})$ , E=0.8-8.0 MeV; analyzed  $\sigma(E)$  measured in the work of **2005Ha69**; deduced uncertainty in the cross section of about 50% above 5 MeV, due to changes in neutron detector efficiency due to different neutron energies that are possible above the  $^{16}\text{O}$  first excited-state, and which were not adequately accounted for in **2005Ha69** who used a moderated neutron detector. Relevance to s-process nucleosynthesis.

**2017Tr03:**  $^{13}\text{C}(\alpha,\text{n})$ , E(cm)<1 MeV; analyzed available data; deduced  $\sigma$ , S-factors, astrophysical reaction rates, asymptotic normalization coefficients.

**2018Cr02:**  $^{13}\text{C}(\alpha,\text{n})$ , E not given; analyzed available data; deduced impact of reaction rate variations on variations of the element surface distributions.

**2018Mo15:**  $^{13}\text{C}(\alpha,\text{n})$ , E=0.8-8 MeV; analyzed previous experiments for  $\sigma(E)$  data and deduced corrected  $\sigma(E)$  based on improved determination of neutron detection efficiency; calculated branching ratios for the  $(\alpha,\text{n}0)$ ,  $(\alpha,\text{n}1)$ ,  $(\alpha,\text{n}2)$ ,  $(\alpha,\text{n}3)$  and  $(\alpha,\text{n}4)$  channels using TALYS code, and compared with experimental values.

**2018Ze01:**  $^{13}\text{C}(\alpha,\text{n})$ , E<9 MeV; calculated  $\sigma$  of inverse reaction using the reciprocity theorem and Web calculator.

**$^{13}\text{C}(\alpha,\text{n})$  (continued)** **$^{17}\text{O}$  Levels***Notes:*

Many authors viewed the uncertainty in excitation energy as equal to the uncertainty in resonance energy; it was not possible to resolve this issue.

E(level)	J <sup>π</sup>	Γ	L	E <sub>α</sub> (res) (keV)	Comments
6860.3 <sup>†</sup> 7				656.0 7	E(level): from E <sub>α</sub> =656.0 keV 7 ( <a href="#">1993Br17</a> ). $(\omega\gamma)_n=1.85\times10^{-4}$ eV 20, $(\omega\gamma)_\gamma<5$ μeV ( <a href="#">1993Br17</a> ).
6972.1 <sup>†</sup> 8				802.2 8	E(level): from E <sub>α</sub> =802.2 keV 8 ( <a href="#">1993Br17</a> ). $(\omega\gamma)_n=4.54\times10^{-4}$ eV 35, $(\omega\gamma)_\gamma<8$ μeV ( <a href="#">1993Br17</a> ).
7166.5 <sup>‡</sup> 15	5/2 <sup>#</sup>	1.5 <sup>‡</sup> keV 2		1056.3 15	E(level): from E <sub>α</sub> =1056.3 keV 15 ( <a href="#">1973Ba10</a> ). See also E <sub>α</sub> (keV)=1056 ( <a href="#">1976Ra36</a> ), 1054 ( <a href="#">2005Ha69</a> ). Γ: See also Γ=5 keV ( <a href="#">1957Wa46</a> ), $\Gamma_n/\Gamma_\alpha=(1300)$ ( <a href="#">1957Wa46</a> ). $\omega\gamma=12.1$ eV 6 ( <a href="#">2005Ha69</a> ), 11.9 eV 6 (Deduced from the resonance yield, $4434\pm135$ n/μC ( <a href="#">1992Br05</a> , <a href="#">1973Ba10</a> ), and the stopping power in $^{13}\text{C}$ (Ziegler, The Stopping and Range of Ions in Matter, Vol. 3 (1977)); see ( <a href="#">1993Br17</a> )). E(level): authors ( <a href="#">2008Pe09</a> ) listed the value from ( <a href="#">1993Ti07</a> ). Involvement in $^{13}\text{C}(\alpha,\text{n})$ is suggested from its population in $^{13}\text{C}(^7\text{Li},\text{t})$ . Γ: $\Gamma_n=400$ keV, $\Gamma_\alpha=0.09$ keV ( <a href="#">2008Pe09</a> ) which is consistent with ( <a href="#">1966Li03</a> : $\Gamma_n/\Gamma>0.99$ in $^{16}\text{O}+\text{n}$ measurement).
7202?					E(level): from E <sub>α</sub> =1336.7 keV 15 ( <a href="#">1973Ba10</a> ). See also E <sub>α</sub> =1336 keV ( <a href="#">2005Ha69</a> ). Γ: See also Γ(kEV)=3 ( <a href="#">1969Sc04</a> ), ≤4 ( <a href="#">1957Wa46</a> ), $\Gamma_n/\Gamma_\alpha=(450)$ ( <a href="#">1957Wa46</a> ). $\omega\gamma=33.3$ eV 18 ( <a href="#">2005Ha69</a> ).
7380.9 <sup>‡</sup> 15	5/2 <sup>+</sup> @	0.6 <sup>‡</sup> keV +2-1		1336.7 15	E(level): from E <sub>α</sub> =1336.7 keV 15 ( <a href="#">1973Ba10</a> ). See also E <sub>α</sub> =1336 keV ( <a href="#">2005Ha69</a> ). Γ: See also Γ(kEV)=3 ( <a href="#">1969Sc04</a> ), ≤4 ( <a href="#">1957Wa46</a> ), $\Gamma_n/\Gamma_\alpha=(450)$ ( <a href="#">1957Wa46</a> ). $\omega\gamma=33.3$ eV 18 ( <a href="#">2005Ha69</a> ).
7383.9 <sup>‡</sup> 15		0.8 <sup>‡</sup> keV +3-2	3,2&	1340.6 15	E(level): from E <sub>α</sub> =1340.6 keV 15 ( <a href="#">1973Ba10</a> ).
7572.9 21	7/2 <sup>-</sup> @	≤1 <sup>‡</sup> keV	4,3&	1587.9 25	E(level): from E <sub>α</sub> =1587.9 keV 21, which is the average of ( <a href="#">1973Ba10</a> : 1590 keV 2) and ( <a href="#">1993Br17</a> : 1585.7 keV 15). See also E <sub>α</sub> =1590 keV ( <a href="#">2005Ha69</a> ). Γ: See also Γ(kEV)=3 ( <a href="#">1969Sc04</a> ), ≤4 ( <a href="#">1957Wa46</a> ). $\omega\gamma=10.8$ eV 5 ( <a href="#">1993Br17</a> ), 11.5 eV 12 ( <a href="#">2005Ha69</a> ).
7693 <sup>‡</sup> 6	5/2 <sup>+</sup> @	≤15 <sup>‡</sup> keV	3,2&	1745 6	E(level): from E <sub>α</sub> =1745 keV 6 ( <a href="#">1973Ba10</a> ). Γ: See also Γ=22 keV ( <a href="#">1957Wa46,1969Sc04</a> ).
7952 <sup>‡</sup> 8	1/2 <sup>+</sup> @	79 <sup>a</sup> keV 10	1,0&	2083 8	E(level): from E <sub>α</sub> =2083 keV 8 ( <a href="#">1973Ba10</a> ). See also E <sub>α</sub> (keV)=2090 ( <a href="#">1956Bo61</a> ), 2080 ( <a href="#">1967Se07</a> ), 2075 ( <a href="#">1973Bu14</a> ). E(level): triplet ( <a href="#">1957Wa46</a> ). Γ: See also Γ(kEV)=75 ( <a href="#">1973Ba10</a> ), 110 ( <a href="#">1957Wa46,1969Sc04</a> ), $\Gamma(\text{lab})=100$ keV ( <a href="#">1956Bo61</a> ), $\Gamma_n/\Gamma_\alpha=10$ ( <a href="#">1957Wa46</a> ). J <sup>π</sup> : See also ( <a href="#">1973Bu14</a> ).
8079 <sup>‡</sup> 8	3/2 <sup>-</sup> @	71 <sup>a</sup> keV 8	2,1&	2250 8	E(level): from E <sub>α</sub> =2250 keV 8 ( <a href="#">1973Ba10</a> ). See also E <sub>α</sub> (keV)=2250 ( <a href="#">1956Bo61,1967Se07</a> ), 2240 ( <a href="#">1963Bu14</a> ). Γ: See also Γ(kEV)=70 ( <a href="#">1957Wa46, 1969Sc04</a> ), 110 ( <a href="#">1973Ba10</a> ), $\Gamma(\text{lab})=100$ keV ( <a href="#">1956Bo61</a> ), $\Gamma_n/\Gamma_\alpha=10$ ( <a href="#">1957Wa46</a> ). J <sup>π</sup> : See also ( <a href="#">1973Bu14</a> ) who suggest the π is ambiguous.
8199 <sup>‡</sup> 8	3/2 <sup>+</sup> @	71 <sup>a</sup> keV 5	1,2&	2407 8	E(level): from 2407 keV 8 ( <a href="#">1973Ba10</a> ). See also

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**$^{13}\text{C}(\alpha, \text{n})$  (continued)** **$^{17}\text{O}$  Levels (continued)**

E(level)	$J^\pi$	$\Gamma$	L	$E_\alpha(\text{res})$ (keV)	Comments
8350 <sup>‡</sup> 4	1/2 <sup>-</sup> @	9 <sup>a</sup> keV 3	0,1&	2604 4	$E_\alpha(\text{keV})=2410$ ( <a href="#">1967Se07</a> ), 2420 ( <a href="#">1956Bo61</a> ), 2430 ( <a href="#">1973Bu14</a> ), 2440 ( <a href="#">1954Tr09</a> ). $\Gamma$ : See also $\Gamma(\text{keV})=60$ ( <a href="#">1957Wa46</a> , <a href="#">1969Sc04</a> ), 70 ( <a href="#">1973Ba10</a> ), $\Gamma(\text{lab})=80$ keV ( <a href="#">1956Bo61</a> ), $\Gamma_n/\Gamma_\alpha=13$ ( <a href="#">1957Wa46</a> ). $J^\pi$ : See also ( <a href="#">1973Bu14</a> ) who suggest the $\pi$ is ambiguous. E(level): from $E_\alpha=2604$ keV 4 ( <a href="#">1973Ba10</a> ). See also $E_\alpha(\text{keV})=2610$ ( <a href="#">1967Se07</a> ), 2605 ( <a href="#">1956Bo61</a> ), $E_n=4440$ keV ( <a href="#">1956Be98</a> : Fig. 5 top).
8408 <sup>‡</sup> 3	5/2 <sup>+</sup> <sup>a</sup>	4 <sup>a</sup> keV 3	3,2 <sup>b</sup>	2680 3	$\Gamma$ : See also $\Gamma(\text{keV})=18$ ( <a href="#">1957Wa46</a> , <a href="#">1969Sc04</a> ), 15 ( <a href="#">1973Ba10</a> ), $\Gamma(\text{lab})\leq 6$ keV ( <a href="#">1956Bo61</a> ), $\Gamma_n/\Gamma_\alpha=6.7$ ( <a href="#">1957Wa46</a> ). $J^\pi$ : See also ( <a href="#">1973Bu14</a> ) who suggest the $\pi$ is ambiguous. E(level): from $E_\alpha=2680$ keV 3 ( <a href="#">1973Ba10</a> ). See also $E_\alpha(\text{keV})=2680$ ( <a href="#">1967Se07</a> ), 2660 ( <a href="#">1954Tr09</a> ), 2690 ( <a href="#">1956Bo61</a> ), $E_n=4530$ keV ( <a href="#">1956Be98</a> : Fig. 5 top). $\Gamma$ : See also $\Gamma(\text{keV})=8$ ( <a href="#">1973Ba10</a> ), 11 ( <a href="#">1957Wa46</a> ), $\Gamma(\text{lab})=10$ keV ( <a href="#">1956Bo61</a> ), $\Gamma_n=3.84$ keV and $\Gamma_\alpha=0.16$ keV ( <a href="#">1967Se07</a> ), $\Gamma_n/\Gamma_\alpha=19$ ( <a href="#">1957Wa46</a> ).
8473 <sup>‡</sup> 3	7/2 <sup>#</sup>	7 <sup>a</sup> keV 3		2765 3	E(level): from $E_\alpha=2765$ keV 3 ( <a href="#">1973Ba10</a> ). See also $E_\alpha(\text{keV})=2770$ ( <a href="#">1967Se07</a> ), 2775 ( <a href="#">1956Bo61</a> ), 2760 ( <a href="#">1954Tr09</a> ), $E_n=4590$ keV ( <a href="#">1956Be98</a> : Fig. 5 top). E(level): doublet ( <a href="#">1957Wa46</a> ). $\Gamma$ : See also $\Gamma(\text{keV})=8$ ( <a href="#">1973Ba10</a> ), 9 ( <a href="#">1957Wa46</a> ), $\Gamma(\text{lab})=10$ keV ( <a href="#">1956Bo61</a> ), $\Gamma_n/\Gamma_\alpha=31$ ( <a href="#">1957Wa46</a> ). E(level): from $E_\alpha=2809$ keV 3 ( <a href="#">1973Ba10</a> ). See also $E_\alpha(\text{keV})=2810$ ( <a href="#">1967Se07</a> ), 2825 ( <a href="#">1956Bo61</a> ), $E_n=4630$ keV ( <a href="#">1956Be98</a> : Fig. 5 top). $\Gamma$ : See also $\Gamma(\text{keV})=6$ ( <a href="#">1973Ba10</a> ), 11 ( <a href="#">1957Wa46</a> ), $\Gamma(\text{lab})\leq 7$ keV ( <a href="#">1956Bo61</a> ), $\Gamma_n=4.57$ keV and $\Gamma_\alpha=0.43$ keV ( <a href="#">1967Se07</a> ), $\Gamma_n/\Gamma_\alpha=2.8$ ( <a href="#">1957Wa46</a> ).
8507 <sup>‡</sup> 3	(3/2,5/2 <sup>-</sup> ) <sup>#a</sup>	5 <sup>a</sup> keV 3	2,3 <sup>b</sup>	2809 3	E(level): from $E_\alpha=3059$ keV 5 ( <a href="#">1973Ba10</a> ). See also $E_\alpha(\text{keV})=3070$ ( <a href="#">1967Se07</a> ), 3090 ( <a href="#">1956Bo61</a> ), 3100 ( <a href="#">1971Ba06</a> ), $E_n=4850$ keV ( <a href="#">1956Be98</a> : Fig. 5 top). E(level): triplet ( <a href="#">1957Wa46</a> ). $\Gamma$ : See also $\Gamma(\text{keV})=50$ ( <a href="#">1973Ba10</a> ), 85 ( <a href="#">1957Wa46</a> ), $\Gamma(\text{lab})=90$ keV ( <a href="#">1956Bo61</a> ), $\Gamma_n/\Gamma_\alpha=17$ ( <a href="#">1957Wa46</a> ). E(level): broad; from $E_\alpha\approx 3100$ keV ( <a href="#">1971Ba06</a> ). E(level): from $E_\alpha=3318$ keV 8 ( <a href="#">1973Ba10</a> ). See also $E_\alpha(\text{keV})=3320$ ( <a href="#">1967Se07</a> ), 3300 ( <a href="#">1954Tr09</a> ), 3330 ( <a href="#">1956Bo61</a> ), 3360 ( <a href="#">1971Ba06</a> ), $E_n=5080$ keV ( <a href="#">1956Be98</a> : Fig. 5 top). $\Gamma$ : See also $\Gamma(\text{keV})=115$ ( <a href="#">1973Ba10</a> ), 110 ( <a href="#">1957Wa46</a> ), $\Gamma(\text{lab})=150$ keV ( <a href="#">1956Bo61</a> ), $\Gamma_n/\Gamma_\alpha=3.5$ ( <a href="#">1957Wa46</a> ). E(level): from $E_\alpha=3415$ keV 4 ( <a href="#">1973Ba10</a> ). See also $E_\alpha(\text{keV})=3420$ ( <a href="#">1956Bo61</a> , <a href="#">1967Se07</a> , <a href="#">1971Ba06</a> ), $E_n=5130$ keV ( <a href="#">1956Be98</a> : Fig. 5 top). $\Gamma$ : See also $\Gamma(\text{keV})=14$ ( <a href="#">1973Ba10</a> ), 35 ( <a href="#">1957Wa46</a> ), $\Gamma(\text{lab})=30$ keV ( <a href="#">1956Bo61</a> ), $\Gamma_n/\Gamma_\alpha=35$ ( <a href="#">1957Wa46</a> ). E(level): from $E_\alpha=3645$ keV 4 ( <a href="#">1973Ba10</a> ). See also $E_\alpha(\text{keV})=3650$ ( <a href="#">1967Se07</a> ), 3670 ( <a href="#">1956Bo61</a> ), 3640 ( <a href="#">1971Ba06</a> ). $\Gamma$ : See also $\Gamma=9$ keV ( <a href="#">1973Ba10</a> ), $\Gamma(\text{lab})\leq 8$ keV ( <a href="#">1956Bo61</a> ). E(level): from $E_\alpha=3690$ keV ( <a href="#">1971Ba06</a> ).
8970 <sup>‡</sup> 4	7/2 <sup>-</sup> <sup>c</sup>	21 <sup>a</sup> keV 3		3415 4	
9146 <sup>‡</sup> 4		4 <sup>a</sup> keV 3		3645 4	
9180 <sup>c</sup>				3690	

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$^{13}\text{C}(\alpha,\text{n})$  (continued) $^{17}\text{O}$  Levels (continued)

E(level)	J <sup>π</sup>	Γ	L	E <sub>α</sub> (res) (keV)	Comments
9199 <sup>‡</sup> 4	5/2 <sup>a</sup>	4 <sup>a</sup> keV 3	2,3 <sup>b</sup>	3714 4	E(level): from E <sub>α</sub> =3714 keV 4 ( <a href="#">1973Ba10</a> ). See also E <sub>α</sub> (keV)=3720 ( <a href="#">1967Se07</a> , <a href="#">1971Ba06</a> ), 3730 ( <a href="#">1956Bo61</a> ), E <sub>n</sub> =5380 keV ( <a href="#">1956Be98</a> : Fig. 5 top). Γ: See also Γ=8 keV ( <a href="#">1973Ba10</a> ), Γ(lab)≤5 keV ( <a href="#">1956Bo61</a> ), Γ <sub>n</sub> =3.86 keV and Γ <sub>α</sub> =0.14 keV ( <a href="#">1967Se07</a> ).
9491 <sup>‡</sup> 4		8 <sup>a</sup> keV 3		4096 4	E(level): from E <sub>α</sub> =4096 keV 4 ( <a href="#">1973Ba10</a> ). See also E <sub>α</sub> (keV)=4120 ( <a href="#">1967Se07</a> ), 4125 ( <a href="#">1956Bo61</a> ), 4110 ( <a href="#">1971Ba06</a> ). Γ: See also Γ=16 keV ( <a href="#">1973Ba10</a> ), Γ(lab)=15 keV ( <a href="#">1956Bo61</a> ).
9.6×10 <sup>3</sup> <sup>c</sup>	3/2 <sup>-c</sup>				E(level): broad; from unresolved broad level near E <sub>α</sub> ≈4.3E3 keV ( <a href="#">1971Ba06</a> ).
9719 <sup>‡</sup> 5	7/2 <sup>+c</sup>	25 <sup>‡</sup> keV		4394 5	E(level): from E <sub>α</sub> =4394 keV 5 ( <a href="#">1973Ba10</a> ). See also E <sub>α</sub> =4380 keV ( <a href="#">1971Ba06</a> ).
9739	3/2 <sup>+c</sup>	15 <sup>a</sup> keV 3		4420	E(level): from E <sub>α</sub> =4420 keV ( <a href="#">1956Bo61</a> , <a href="#">1967Se07</a> , <a href="#">1971Ba06</a> ), This level is associated with E <sub>x</sub> =9786 keV in Adopted Levels. Γ: See also Γ(lab)=25 keV ( <a href="#">1956Bo61</a> ). It appears this level is not real. No single experiment reports more than two of the three levels at 9.72, 9.74 and 9.77 MeV. The energies of these levels are better resolved in <sup>16</sup> O(n,n) ( <a href="#">1980Ce03</a> ).
9773 <sup>‡</sup> 15		≈25 <sup>‡</sup> keV		4465 15	E(level): from E <sub>α</sub> =4465 keV 15 ( <a href="#">1973Ba10</a> ). See also E <sub>α</sub> (keV)=4500 ( <a href="#">1956Bo61</a> ), (4490) ( <a href="#">1967Se07</a> ). Γ: See also Γ(lab)=70 keV ( <a href="#">1956Bo61</a> ).
9863 <sup>‡</sup> 5	(9/2 <sup>+</sup> ) <sup>c</sup>	14 <sup>‡</sup> keV		4583 5	E(level): from E <sub>α</sub> =4583 keV 5 ( <a href="#">1973Ba10</a> ). See also E <sub>α</sub> =4580 keV ( <a href="#">1971Ba06</a> ).
9876 <sup>‡</sup> 15	9/2 <sup>a</sup>	5 <sup>a</sup> keV 3	4,5 <sup>b</sup>	4600 15	E(level): from E <sub>α</sub> =4600 keV 15 ( <a href="#">1973Ba10</a> ). See also E <sub>α</sub> (keV)=4620 ( <a href="#">1967Se07</a> ), 4630 ( <a href="#">1956Bo61</a> ). Γ: See also Γ(keV)≈10 ( <a href="#">1973Ba10</a> ), Γ(lab)=15 keV ( <a href="#">1956Bo61</a> ), Γ <sub>n</sub> =4.7 keV and Γ <sub>α</sub> =0.3 keV ( <a href="#">1967Se07</a> ).
9976 <sup>‡</sup> 20	5/2 <sup>+c</sup>	≈80 <sup>‡</sup> keV		4730 20	E(level): from E <sub>α</sub> =4730 keV 20 ( <a href="#">1973Ba10</a> ). See also E <sub>α</sub> (keV)=4750 ( <a href="#">1956Bo61</a> ), 4700 ( <a href="#">1971Ba06</a> ), (4770) ( <a href="#">1967Se07</a> ). Γ: See also Γ(lab)=200 keV ( <a href="#">1956Bo61</a> ).
10045 <sup>‡</sup> 20		≈100 <sup>‡</sup> keV		4820 20	E(level): from E <sub>α</sub> =4820 keV 20 ( <a href="#">1973Ba10</a> ). See also E <sub>α</sub> =4850 keV ( <a href="#">1967Se07</a> ).
10136 <sup>c</sup>				4940	E(level): from E <sub>α</sub> =4940 keV ( <a href="#">1971Ba06</a> ).
10177 <sup>‡</sup> 5	(5/2 <sup>+</sup> , 7/2 <sup>-</sup> ) <sup>d</sup>	50 <sup>a</sup> keV 3		4993 5	E(level): from E <sub>α</sub> =4993 keV 5 ( <a href="#">1973Ba10</a> ). See also E <sub>α</sub> (keV)=4980 ( <a href="#">1971Ba06</a> ), 4995 ( <a href="#">1970Ro08</a> ), 5040 ( <a href="#">1967Se07</a> ), 5050 ( <a href="#">1956Bo61</a> ). Γ: See also 45 keV ( <a href="#">1973Ba10</a> ), Γ(lab)=65 keV ( <a href="#">1956Bo61</a> ).
10335 <sup>‡</sup> 15	(5/2 <sup>+</sup> , 7/2 <sup>-</sup> ) <sup>d</sup>	150 <sup>‡</sup> keV		5200 15	E(level): from E <sub>α</sub> =5200 keV 15 ( <a href="#">1973Ba10</a> ). See also E <sub>α</sub> =5185 keV ( <a href="#">1970Ro08</a> ).
10422.3 <sup>e</sup> 20	(5/2 <sup>-</sup> , 7/2 <sup>-</sup> ) <sup>d</sup>	14 <sup>f</sup> keV 3		5314.0 25	E(level): from E <sub>α</sub> =5314.0 keV 25 ( <a href="#">1975Be44</a> ). See also E <sub>α</sub> (keV)=5317 10 ( <a href="#">1963Sp02</a> ), 5325 10 ( <a href="#">1973Ba10</a> ), 5290 ( <a href="#">1967Se07</a> ), 5370 ( <a href="#">1970Ro08</a> ). Γ: See also Γ(keV)=11 9 ( <a href="#">1975Be44</a> ), 23 keV ( <a href="#">1973Ba10</a> ).

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**$^{13}\text{C}(\alpha,\text{n})$  (continued)** **$^{17}\text{O}$  Levels (continued)**

E(level)	J <sup>π</sup>	Γ	E <sub>α</sub> (res) (keV)	Comments
10498	(5/2 <sup>+</sup> ,7/2 <sup>-</sup> ) <sup>d</sup>		5413	E(level): from E <sub>α</sub> =5413 keV, which is the average of (1967Se07: 5410 keV) and (1970Ro08: 5415 keV).
10558.5 <sup>e</sup> 20	(7/2 <sup>-</sup> ,9/2 <sup>+</sup> ) <sup>d</sup>	51 <sup>e</sup> keV 2	5492.0 25	E(level): from E <sub>α</sub> =5492.0 keV 25 (1975Be44). See also E <sub>α</sub> (keV)=5496 10 (1963Sp02), 5540 (1970Ro08). Γ: See also Γ=50 keV 20 (1963Sp02).
10777.9 <sup>e</sup> 20	(1/2 <sup>+</sup> ,7/2 <sup>-</sup> ) <sup>d</sup>	74 <sup>e</sup> keV 3	5779.0 25	E(level): from E <sub>α</sub> =5779.0 keV 25 (1975Be44). See also E <sub>α</sub> (keV)=5771 10 (1963Sp02), 5790 (1970Ro08). Γ: See also Γ=85 keV 30 (1963Sp02).
10905 <sup>e</sup> 2	(5/2) <sup>d</sup>	46 <sup>e</sup> keV 2	5945 3	E(level): from E <sub>α</sub> =5945 keV 3 (1975Be44). See also E <sub>α</sub> (keV)=5945 10 (1963Sp02), 5995 (1970Ro08). Γ: See also Γ=55 keV 20 (1963Sp02).
11036 <sup>e</sup> 2		31 <sup>e</sup> keV 3	6117 3	E(level): from E <sub>α</sub> =6117 keV 3 (1975Be44). See also E <sub>α</sub> =6107 keV 10 (1963Sp02). Γ: See also Γ=45 keV 10 (1963Sp02).
11076 <sup>g</sup> 5	1/2 <sup>-</sup> <sup>g</sup>	5.0 <sup>g</sup> keV 11	6169	T=3/2 (1976Mc11) E(level): from E <sub>α</sub> =6169 keV which is calculated from the E <sub>x</sub> given in (1976Mc11); see also E <sub>α</sub> =6220 keV (1970Ro08). J <sup>π</sup> : See also (3/2 <sup>-</sup> ,7/2 <sup>+</sup> ) (1970Ro08). Γ: From (1976Mc11). A preliminary result, 5 keV 1, from a BAPS was picked up by (1973Ad02): $^{18}\text{O}(^3\text{He},\alpha)$ and used to derive various partial widths. See also $(\Gamma_{a0}\Gamma_{n0})^{1/2}/\Gamma_{tot}=0.23$ (1976Mc11) which corresponds to $\Gamma_{a0}/\Gamma=0.06$ 2 when combined with the value $\Gamma_{n0}/\Gamma=0.91$ 15 (1973Ad02).
11238 <sup>e</sup> 2	(3/2 <sup>-</sup> ,7/2 <sup>+</sup> ) <sup>d</sup>	80.0 <sup>e</sup> keV 25	6380 3	E(level): from E <sub>α</sub> =6380 keV 3 (1975Be44). See also E <sub>α</sub> (keV)=6367 10 (1963Sp02), 6480 (1970Ro08). Γ: See also Γ=100 keV 30 (1963Sp02).
11622 <sup>e</sup> 2		65 <sup>e</sup> keV 2	6883 3	E(level): from E <sub>α</sub> =6883 keV 3 (1975Be44). See also E <sub>α</sub> =6878 keV 10 (1963Sp02). Γ: See also Γ=120 keV 30 (1963Sp02).
11751 <sup>f</sup> 10		40 <sup>f</sup> keV 25	7051 10	E(level): from E <sub>α</sub> =7051 keV 10 (1963Sp02).
11816 <sup>f</sup> 15	(5/2,7/2) <sup>d</sup>	12 <sup>f</sup> keV 3	7136 15	E(level): from E <sub>α</sub> =7136 keV 15 (1963Sp02). See also E <sub>α</sub> =7160 keV (1970Ro08).
12005 <sup>f</sup> 15			7384 15	E(level): from E <sub>α</sub> =7348 keV 15 (1963Sp02).
12109 <sup>f</sup> 20		150 <sup>f</sup> keV 50	7520 20	E(level): from E <sub>α</sub> =7520 keV 20 (1963Sp02).
12274 <sup>f</sup> 15		100 <sup>f</sup> keV 30	7736 15	E(level): from E <sub>α</sub> =7736 keV 15 (1963Sp02).
12385 <sup>f</sup> 20			7880 20	E(level): from E <sub>α</sub> =7880 keV 20 (1963Sp02).
12421 <sup>f</sup> 15			7927 15	E(level): from E <sub>α</sub> =7927 keV 15 (1963Sp02).
12458 <sup>g</sup> 5	3/2 <sup>-</sup> <sup>g</sup>	8 <sup>g</sup> keV 2	7976	T=3/2 (1976Mc11) E(level): from E <sub>α</sub> =7976 keV which is calculated from the E <sub>x</sub> given in (1976Mc11). See also E <sub>α</sub> =8000 keV (1970Ro08). J <sup>π</sup> : See also (3/2 <sup>-</sup> ,9/2 <sup>+</sup> ) (1970Ro08).
12596 <sup>f</sup> 15		75 <sup>f</sup> keV 30	8156 15	E(level): from E <sub>α</sub> =8156 keV 15 (1963Sp02).
12670 <sup>f</sup> 15	(3/2 <sup>-</sup> ,9/2 <sup>+</sup> ) <sup>d</sup>	≈75 <sup>f</sup> keV	8253 15	Γ: Beginning in (1971Aj02) the Γ for this level was listed at ≈5 keV attributed to $^{13}\text{C}(\alpha,\text{n})$ (1969Sp02). However this was a typo. (1969Sp02) report ≈75 keV. E(level): from E <sub>α</sub> =8253 keV 15 (1963Sp02). See also E <sub>α</sub> =8328 keV (1970Ro08).
12813 <sup>f</sup> 25			8440 25	E(level): from E <sub>α</sub> =8440 keV 25 (1963Sp02).
12928 <sup>f</sup> 20	(1/2 <sup>+</sup> ,7/2 <sup>-</sup> ) <sup>d</sup>	≥150 <sup>f</sup> keV	8590 20	E(level): from E <sub>α</sub> =8590 keV 20 (1963Sp02). See also E <sub>α</sub> =8657 keV (1970Ro08).
12944 <sup>g</sup> 6	1/2 <sup>+g</sup>	6 <sup>g</sup> keV 2	8612	T=3/2 (1976Mc11)

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$^{13}\text{C}(\alpha,\text{n})$  (continued) $^{17}\text{O}$  Levels (continued)

E(level)	$J^\pi$	$\Gamma$	$E_\alpha(\text{res})$ (keV)	Comments
12993 <sup>g</sup> 6	$5/2^-$ <sup>g</sup>	$\leq 3$ <sup>g</sup> keV	8676	E(level): from $E_\alpha=8612$ keV which is calculated from the $E_x$ given in (1976Mc11). T=3/2 (1976Mc11)
13077 <sup>f</sup> 15		$16$ <sup>f</sup> keV 4	8785 15	E(level): from $E_\alpha=8785$ keV 15 (1963Sp02).
13485 <sup>f</sup> 15		$\approx 120$ <sup>f</sup> keV	9319 15	E(level): from $E_\alpha=9319$ keV 15 (1963Sp02).
13610 <sup>f</sup> 15		$250$ <sup>f</sup> keV 100	9483 15	E(level): from $E_\alpha=9483$ keV 15 (1963Sp02). $\Gamma$ : =150-350 keV.

<sup>†</sup> From (1993Br17).<sup>‡</sup> From (1973Ba10).<sup>#</sup> From (1957Wa46).<sup>@</sup> From (1969Sc04). But see (1973Bu14) who claim  $\pi$  deduced by (1969Sc04) is sometimes ambiguous.<sup>&</sup>  $L_\alpha, L_{\alpha'}$  (1969Sc04).<sup>a</sup> From (1967Se07).<sup>b</sup>  $L_\alpha, L_n$  (1967Se07).<sup>c</sup> From (1971Ba06).<sup>d</sup> From (1970Ro08).<sup>e</sup> From (1975Be44).<sup>f</sup> From (1963Sp02).<sup>g</sup> From (1976Mc11).