

Adopted Levels, Gammas

Type	Author	Citation	Literature Cutoff Date
Full Evaluation	J. H. Kelley, G. C. Sheu	ENSDF	16-Jan-2018

Q(β⁻)=-18.7×10³ 10; S(n)=15558 20; S(p)=1469 8; Q(α)=-9040 10 2017Wa10

Theoretical analyses of ¹⁷Ne.

Nuclear Mass and Level Properties:

General analyses of the ¹⁷Ne nuclear properties are given in (1973Re17,1978Gu10,2004Ge02,2018Fo04). Comparisons of the ¹⁷Ne level properties with those of levels in the ¹⁷N mirror nucleus are found in (1996Ti02, 2004Ga07,2006Fo08,2006Li18,2018Fo02).

Halo Character.

Because of the unbound nature of ¹⁶F, the nucleus ¹⁷Ne is considered a Borromean nuclear system similar to ^{6,8}He and ¹¹Li. The low binding energy suggests the ground or first excited states may exhibit properties that can be described as a ¹⁵O core surrounded by a diffuse 2-proton halo. Detailed studies on ¹⁷Ne are found in (1990Ha29, 1995Zh35, 1997Li22, 1998Na01, 2001Fo05, 2003Zh29, 2004Ta40, 2005Gr11, 2008Ne13, 2013Zh27, 2016Hw01). More general analyses of one- and two- nucleon halo nuclides, including ¹⁷Ne, are given in (1994Fe01, 1995Ta06, 1999Kn04, 2001Oz04, 2002Gu10, 2007Be58, 2011Al11, 2014Ch39, 2017Ah08).

Possible soft dipole excitations and pygme dipole resonances are discussed in (2011OI01, 2012Ma12, 2017Lv02).

3-Body nature of ¹⁷Ne.

Studies of the 2-proton correlations associated with the 3-body nature of ¹⁷Ne are given in (2001Gr29, 2003Gr01, 2004Ga10, 2005Ga04, 2005Ga49, 2005Pf01, 2005Pf02, 2007Gr12, 2007Gr13, 2010Oi01).

The ¹⁵O+2p resonances of ¹⁷Ne are relevant for the 2p capture rate on ¹⁵O, see (2006Gr20,2016Ca38). The Γ_{2p} width of the E_x=1288 keV: J^π=3/2⁻ state is of particular interest (2000Gr16) since it is bound to one proton emission, unbound to 2p emission, but γ decays 100%.

3-body nature of ¹⁹Mg.

Some states in ¹⁷Ne are important for the 2p decay of ¹⁹Mg and the 2p+¹⁷Ne system, see (2001Mu23, 2003Gr04, 2003Gr24). See other results in (2000Gr18, 2004Pf02, 2007Fo07, 2010Gr06, 2013OI02, 2017Go17, 2017Kw01).

β-decay studies.

The β decay of ¹⁷Ne is rather complex. The first forbidden decay to ¹⁷F is stronger than expected (1972To03, 1994Jo04, 1997Mi08, 1999Ba21, 2000Ni14, 2003Sm02); the observation may be connected with the halo nature of ¹⁷Ne. See other general analyses and comparison with ¹⁷N decay in (1965Ha31, 1970Wi02). The decay is dominated by β-p reactions to levels in ¹⁶O (1963Ba63, 1964Da13, 1964Fl03, 1964Mc16, 1965Ha20, 1967Es02, 1988Bo39, 2002Ch61, 2002Mo19). Additional β-α branches to ¹³C are also present. See theoretical discussion in (1973Ha77, 1977Ce05).

¹⁷Ne Levels

Cross Reference (XREF) Flags

A	¹ H(¹⁷ Ne,P)	H	⁹ Be(¹⁸ Ne, ¹⁷ Ne)	O	²⁰ Ne(³ He, ⁶ He)
B	¹ H(¹⁷ Ne, ¹⁵ O2p)	I	⁹ Be(²⁰ Ne, ¹⁷ Ne)	P	Mg(o,p, ¹⁷ Ne)
C	¹ H(¹⁸ Ne,D)	J	¹⁶ O(³ He,2n)	Q	Mg(p, ¹⁷ Ne)
D	⁹ Be(¹² C,X), ¹² C(¹² C,X)	K	¹⁶ O(¹⁴ N, ¹³ B)	R	Si(¹⁷ Ne,X)
E	⁹ Be(¹⁷ Ne, ¹⁷ Ne')	L	¹⁸ Na p decay	S	Ni(²⁰ Ne, ¹⁷ Ne)
F	⁹ Be(¹⁷ Ne,X)	M	¹⁹ F(p,3n)	T	Au(¹⁷ Ne, ¹⁷ Ne'γ)
G	⁹ Be, ¹² C, ²⁷ Al(¹⁷ Ne,X)	N	¹⁹ Mg 2p decay:4.0 ps	U	Pb(¹⁷ Ne, ¹⁷ Ne)

E(level)	J ^π	T _{1/2}	XREF	Comments
0.0	1/2 ⁻	109.2 ms 6	A CD FG IJ LMNOPQRSTU	%ε+%β ⁺ =100; %εp=95.2 46 (2002Mo19); %εα=2.77 19 (2002Mo19) T=3/2 μ=+0.7873 14 (2005Ge06) T _{1/2} : weighted average of: 109.3 ms 6 (1988Bo39), 109 ms 1 (1971Ha05). Other results can be found in (1967Es02,1967Fi10,1964Mc16) and Hardy, et al., Nucl. Isospin, Proc. 1969 Conf. (Academic Press, 1969) 725.

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Adopted Levels, Gammas (continued) ^{17}Ne Levels (continued)

<u>E(level)</u>	<u>J^π</u>	<u>XREF</u>				<u>Comments</u>	
						$\Delta M=16500.4$ keV 4: From ISOLTRAP measurement in (2008Ge07). Note: most level energies are taken from (1998Gu10) who measured $\Delta M=16453$ keV 32. Since other measurements (2002Ch44,2017Br07) tend to support the level energy differences found in (1998Gu10), those relative level energy spacings are preserved in the present evaluation.	
1288 8	$(3/2,1/2)^-$	C	K	O	T	%IT=100 (2002Ch44) B(E2) $\uparrow=0.0090$ 18 (2016Ma42) E(level): From (1998Gu10), also see $E_x=1275$ keV 22 from (2002Ch44) Au($^{17}\text{Ne},^{17}\text{Ne}'$). J^π : $3/2^-$ is preferred based on comparison with the ^{17}N levels (1998Gu10).	
1749 8	$(5/2,7/2)^-$	E	H	O	TU	$\Gamma_{2p}/\Gamma < 1.6 \times 10^{-4}$ (2017Sh29). %p=100 (2017Br07) J^π : $5/2^-$ is preferred based on comparison with the ^{17}N levels (1998Gu10). E(level): From average of $E_x=1764$ keV 12 (1998Gu10) and $E_x=1745$ keV 7 (2018Ch25). Also see $E_x=1770$ keV 20 from (2017Br07) $^9\text{Be}(^{17}\text{Ne},^{17}\text{Ne}')$.	
1908 15	$1/2^+, (3/2,5/2)^+$	C	H	K	O	TU	%p \approx 100 (2002Ch44) B(E1) $\uparrow < 0.00007$ (2016Ma42) J^π : $1/2^+$ is preferred (1998Gu10). E(level): From (1998Gu10), also see $E_x=1900$ keV 78 from (2002Ch44) Au($^{17}\text{Ne},^{17}\text{Ne}'$).
2651 12		C	E		O	TU	%p=100 (2016Ma42,2017Br07) $J=5/2^+, 5/2^-, 3/2^+, 7/2^-$ E(level): From (1998Gu10) $^{20}\text{Ne}(^3\text{He},^6\text{He})$. While only one peak has been observed in the region, the group is suggested as a possible doublet (1998Gu10) since relatively poor energy resolution may conceal two groups. An analysis of the angular distributions for the $E_x \approx 2623$ keV region is consistent with $L=3$ ($J^\pi=(5/2^-, 7/2^-)$), while analysis of the $E_x \approx 2765$ keV region is consistent with $L=2$ ($J^\pi=(3/2^+, 5/2^+)$). A later analysis of Pb($^{17}\text{Ne},^{17}\text{Ne}'$) (2016Ma42) suggests an unresolved doublet with $^{17}\text{Ne}^*(2614$ keV 20: $5/2^+$) and $^{17}\text{Ne}^*(2692$ keV 21: $3/2^-)$.
2997 11	$(7/2,5/2)^-$			K	O	T	J^π : $7/2^-$ is preferred based on comparison with the ^{17}N levels (1998Gu10). E(level): From (1998Gu10).
3415 38	$(5/2^-)$					U	%p=100 (2016Ma42) E(level), J^π : From (2016Ma42).
3548 20	$(9/2,11/2)^-$	E		K	O	T	%p=100 (2016Ma42,2017Br07) J^π : $9/2^-$ is preferred based on comparison with the ^{17}N levels (1998Gu10). E(level): From (1998Gu10).
4010 10	$(3/2,5/2)^+$				O	T	J^π : $3/2^+$ is preferred based on comparison with the ^{17}N levels (1998Gu10). E(level): From (1998Gu10).
4487 22					O	T	E(level): From (1998Gu10).
5141 62					O	TU	E(level): From (1998Gu10), also see $E_x=5210$ keV 79 from (2016Ma42) Pb($^{17}\text{Ne},^{17}\text{Ne}'$).
$\approx 5.3 \times 10^3?$				K			E(level): From (2000O101).
5722 23					O	T	E(level): From (1998Gu10).
6132 35					O	T	E(level): From (1998Gu10).
6366 22					O	T	E(level): From (1998Gu10).
10.06×10^3 15						U	E(level): From (2016Ma42).

Adopted Levels, Gammas (continued) $\gamma({}^{17}\text{Ne})$

$E_i(\text{level})$	J_i^π	E_γ	I_γ	E_f	J_f^π
1288	$(3/2, 1/2)^-$	1288	100	0.0	$1/2^-$

Adopted Levels, Gammas**Level Scheme**

Intensities: Relative photon branching from each level

