$^{134}_{52}$ Te $_{82}$ -1

¹³⁵Sb β⁻n decay 1981Ho07,1979Kr03

History					
Туре	Author	Citation	Literature Cutoff Date		
Full Evaluation	A. A. Sonzogni	NDS 103, 1 (2004)	31-Jul-2004		

Parent: ¹³⁵Sb: E=0.0; $J^{\pi}=(7/2^+)$; $T_{1/2}=1.68 \text{ s } 2$; $Q(\beta^-n)=4.78\times10^3 \ 10$; $\%\beta^-n \text{ decay}=17.6 \ 22^{-135}\text{Sb-}\%\beta^-n \text{ decay}$: $\%\beta^-n=22 \ 3 \ (2002\text{Sh08})$.

1989H008, 1981H007: measured Ey, Iy, $\gamma\gamma$, β -strength functions. Deduced β^{-} delayed neutron feedings.

1979Kr03 (also 1976Kr18): measured delayed neutron energies and intensities, γ -ray spectra.

Others: 1966To02, 1968To19, 1968To18, 1974Sh18, 1974Fr09, 1976Lu02, 1977Ru04, 1978Cr03, 1979Kr03, 1980Lu04, 1993Ru01. See 1975Iz03, 1977Ru10, 1982Ru01, 1984Ma39, 1989BrZI for analysis and compilations.

¹³⁴Te Levels

E(level) [†]	J ^π ‡	Comments
0.0	0+	Neutron feeding=62% 6 (1989Ho08), 47% 12 (1979Kr03).
1279.07 10	2+	Neutron feeding=21% 3 (1989Ho08), 29% 6 (1981Ho07), 26% 10 (1979Kr03).
1576.15 15	4+	Neutron feeding=11% 2 (1989Ho08), 15% 4 (1981Ho07), 3% 3 (1979Kr03).
1691.45 18	6+	Neutron feeding=6% 1 (1989Ho08), 9% 2 (1981Ho07), 21% 5 (1979Kr03).
2397.4	$(6)^{+}$	Neutron feeding=4% 3 (1979Kr03). No feeding shown by 1981Ho07 and 1989Ho08.

[†] From least-squares fit to $E\gamma$.

[‡] From Adopted Levels.

 $\gamma(^{134}\text{Te})$

I γ normalization: Ti(1279 γ)=38 6.

E_{γ}^{\dagger}	$I_{\gamma}^{\ddagger@}$	E_i (level)	\mathbf{J}_i^{π}	$E_f J_f^{\pi}$	Mult. [#]	α ^{&}	Comments
115.30 10	15.5 12	1691.45	6+	1576.15 4+	E2	1.04	$\alpha(K)=0.755\ 23;\ \alpha(L)=0.224\ 7;\ \alpha(M)=0.0463\ 14;$ $\alpha(N+)=0.0104\ 4$
297.08 10	43 2	1576.15	4+	1279.07 2+	E2	0.0399	α (K)=0.0331 <i>10</i> ; α (L)=0.00540 <i>17</i> ; α (M)=0.00109 4; α (N+)=0.00025 <i>1</i>
1279.06 10	94 <i>3</i>	1279.07	2^{+}	$0.0 0^+$	E2	0.00086	α =0.00086; α (K)=0.00074 2

[†] From 1989Ho08.

[±] From 1989H008, relative to $I\gamma(1127.0\gamma \text{ in } {}^{135}\text{Te})=100.$

[#] From adopted gammas.

[@] For absolute intensity per 100 decays, multiply by 0.070 14.

 $^{\&}$ Total theoretical internal conversion coefficients, calculated using the BrIcc code (2008Ki07) with Frozen orbital approximation based on γ -ray energies, assigned multipolarities, and mixing ratios, unless otherwise specified.

Delayed Neutrons (¹³⁴Te)

E(n) [†]	E(¹³⁴ Te)	I(n) ^{‡#}	E(¹³⁵ Te)
	0.0	62 6	
162 <i>3</i>	1576.15	0.96	5240
458 2	1279.07	1.4	5240
499 <i>3</i>	2397.4	1.0	6400
549 <i>3</i>	1576.15	1.1	5630
623 <i>3</i>	1691.45	0.96	5810

135 Sb β^- n decay 1981Ho07,1979Kr03 (continued)

Delayed Neutrons (continued)

E(n) [†]	E(¹³⁴ Te)	I(n) ^{‡#}	E(¹³⁵ Te)	Comments
783 <i>3</i>	1576.15	0.96	5810	
848 2	1279.07	1.7	5630	
977 2	1691.45	2.2	6170	
1042 2	1691.45	7.7	6240	This transition could also feed 1279 level from 5810 level.
1201 2	1691.45	4.7	6400	
1251 4	1576.15	0.87	6330	
1322 4	1576.15	2.1	6400	
1384 4	1279.07	1.5	6170	
1458 2	1279.07	8.7	6240	
1549 <i>3</i>	1279.07	2.3	6330	
1618 5	1279.07	0.87	6400	

[†] Recoil-corrected energies from 1979Kr03.

 \ddagger From 1979Kr03, normalized so that total reported neutron feeding to first 2⁺, 4⁺ and 6⁺ states is 38%, based on 62% 6 (1989Ho08) as the neutron feeding of g.s..
For absolute intensity per 100 decays, multiply by 0.176 22.

¹³⁵Sb β⁻n decay 1981Ho07,1979Kr03



 γ Intensities: $I_{(\gamma+ce)}$ per 100 parent decays I(n) Intensities: Relative I(n)

