

$^{12}\text{C}(^{10}\text{B}, ^9\text{Be})$

Type	Author	History	Citation	Literature Cutoff Date
Full Evaluation	J. H. Kelley, C. G. Sheu and J. E. Purcell		NDS 198,1 (2024)	1-Aug-2024

[1965Sa07](#): $^{12}\text{C}(^{10}\text{B}, ^9\text{Be})$ E=105 MeV; populated ground state and unresolved states at $E_x=3.5$ MeV.

[1974Na20](#): $^{12}\text{C}(^{10}\text{B}, ^9\text{Be})$ E=100 MeV. Measured $\sigma(\theta)$ for $\theta=10^\circ$ to 40° , deduced level energies. Considered model dependencies for S.

[2022Ar03](#): $^{12}\text{C}(^{10}\text{B}, ^9\text{Be})$ E=41.3 MeV; measured $\sigma(\theta)$ for $\theta=7^\circ$ to 51° . Analyzed data using a modified DWBA; deduced p+ $^{12}\text{C} \rightarrow ^{13}\text{N}$ ANC and analyzed astrophysical reaction rates. The study included a reanalysis of the collection of low-energy (p, γ) data resulting in $\Gamma_p=33.5$ keV 10, 46.0 keV 34 and 280 keV 50 and $\Gamma_\gamma=0.63$ eV 7, 0.35 eV 8 and 2200 eV 500 for $^{13}\text{N}^*$ (2.36, 3.50, 10.26), respectively. Note: the tail of the 10.26 MeV state is fit using data below the $E_x=2$ MeV region.

[1975Ra13](#): DWBA analysis of ([1974Na20](#)) data.

[1976Ku06](#): DWBA analysis of $^{13}\text{N}^*$ (2.37 MeV) data from ([1974Na20](#)).

[2000Fe08](#): $^{12}\text{C}(^{10}\text{B}, ^9\text{Be})$ E=100 MeV; analyzed ANC method.

 ^{13}N Levels

$E(\text{level})^\dagger$
0
2.37×10^3
$3.51 \times 10^3 \ddagger$
$3.55 \times 10^3 \ddagger$
$6.9 \times 10^3 \ddagger$
$7.17 \times 10^3 \ddagger$
$7.39 \times 10^3 \ddagger$
8.82×10^3

† From ([1974Na20](#)).

‡ Unresolved.