

$^{12}\text{C}(\text{n},\gamma):\text{res}$

Type	Author	History	Citation	Literature Cutoff Date
Full Evaluation	J. H. Kelley, C. G. Sheu and J. E. Purcell		NDS 198,1 (2024)	1-Aug-2024

- 1950Ki68: $^{12}\text{C}(\text{n},\gamma)$ E=slow; measured E_γ , I_γ .
- 1963Ma60: $^{12}\text{C}(\text{n},\gamma)$ E \approx 30,65 keV; measured neutron radiative capture cross sections for $Z\leq 92$.
- 1975Ar19: $^{12}\text{C}(\text{n},\gamma_0)$ E=14 MeV; measured γ -yields, $I_\gamma(\theta)$. There is an indication γ_0 peaks at backward angles.
- 1984Wo05: $^{12}\text{C}(\text{n},\gamma_0)$, (pol. n,γ_0) E=12-18.8 MeV; measured $\sigma(\theta)$, analyzing power for $\theta\approx 50^\circ-140^\circ$. Deduced states at $E_x=16-21.75$ MeV. Combined analysis with other data indicates a secondary doorway state. Deduce a_1 , a_2 , b_1 , b_2 from Legendre fit.
- 1985AuZZ: $^{12}\text{C}(\text{n},\gamma)$ E=6.5-18.5 MeV; measured $\sigma(\theta)$. $^{12}\text{C}(\text{pol. n},\gamma)$ E=12-18.8 MeV; measured analyzing, vector analyzing power vs θ . ^{13}C deduced resonances, doorway characteristics.
- 1986Be17: $^{12}\text{C}(\text{n},\gamma)$ E=7-19.5 MeV; measured $\sigma(E,\theta=90^\circ)$. Deduced $T_{<}=1/2$ giant resonance at $E_x=20.8$ MeV with $\Gamma\approx 2.5$ MeV along with lower unresolved strength near $E_n\approx 10$ MeV identified as the *pygmy* resonance. DSD model.
- 1987Au02: $^{12}\text{C}(\text{n},\gamma)$ E=6.5-18.5 MeV; measured $\sigma(E,\theta)$, $\theta=90^\circ$. Pygmy and giant resonance region. Direct-semidirect calculations.
- 1988McZT: $^{12}\text{C}(\text{n},\gamma)$; analyzed $\sigma(E_n)$.
- 1989Hu15: $^{12}\text{C}(\text{n},\gamma)$ E=14.2 MeV; measured $\gamma(\theta)$; deduced γ -multipolarity. Direct-semidirect model.
- 1990Ha19: $^{12}\text{C}(\text{n},\gamma)$ E=8-11 MeV; measured $\sigma(\theta=90^\circ)$ in the Pygmy and giant resonance region.
- 1990Ma52: $^{12}\text{C}(\text{n},\gamma)$ E<46 keV; measured effective capture $\sigma(E)$; deduced Maxwellian averaged σ .
- 1991Hu05: $^{12}\text{C}(\text{n},\gamma)$ E=7-14 MeV; measured $\sigma(\theta)$ vs E, E_γ , I_γ , $\gamma(\theta)$. Analyzed pygmy resonance.
- 1991Na06,1991Na19: $^{12}\text{C}(\text{n},\gamma)$ E=30 keV, stellar energy; measured capture σ , E_γ , I_γ ; deduced nucleosynthesis implications.
- 1992Wi08: $^{12}\text{C}(\text{pol. n},\gamma)$ E=20-35 MeV; measured $\gamma(\theta)$, A_γ vs E, $\gamma(\text{recoil})$ -coin; deduced E2, E1 capture interference. Direct semidirect model. Deduce a_1 , a_2 , b_1 , b_2 from Legendre fit.
- 1994Oh02, 1996Na27: $^{12}\text{C}(\text{n},\gamma)$ E=10-250 keV; measured E_γ , I_γ , $\sigma(E)$; deduced Maxwellian averaged σ and astrophysical implications.
- 1998Ki09: $^{12}\text{C}(\text{n},\gamma)$ E=550 keV; measured E_γ , I_γ ; deduced partial capture σ . Deduced spectroscopic factor of $^{13}\text{C}^*$ (3.09 MeV).
- 1999Oh04: $^{12}\text{C}(\text{n},\gamma)$ E \approx 42 keV; measured capture $\sigma(E)$, E_γ , I_γ .
- 2008Oh05: XUNDL dataset compiled by McMaster, 2008.
- E(n)=10-80 keV neutrons produced in the $^7\text{Li}(\text{p},\text{n})$ reaction using the 3.2 MV Pelletron accelerator at the Tokyo Institute of Technology. Measured E_γ , I_γ , $\gamma\gamma$ coin using anti-Compton NaI(Tl) spectrometer, time-of-flight method. Non-resonant study.

Theory:

- 1971AI33,1971AIYV: $^{12}\text{C}(\text{n},\gamma)$ E \approx 30 keV; compiled experimental Maxwellian averaged σ ; deduced empirical correlation between σ and nucleosynthesis abundances.
- 1974Ma10: $^{12}\text{C}(\text{n}_0,\gamma)$; analyzed isospin splitting in the giant dipole resonance.
- 1986Li16: $^{12}\text{C}(\text{pol. n},\gamma)$ E \leq 9 MeV; calculated polarization effects.
- 1987LyZY: $^{12}\text{C}(\text{n},\gamma)$ E=slow; analyzed data; deduced model parameters, capture mechanism.
- 1990Wa22: $^{12}\text{C}(\text{n},\gamma)$; analyzed data; deduced calibration γ -energies. Proposed $E_\gamma=3683.915$ keV 15 for transition 3.684 \rightarrow 0.
- 1991Ho18: $^{12}\text{C}(\text{n},\gamma)$ E=threshold-30 keV; calculated σ ; deduced reaction mechanism.
- 1993Ho06: $^{12}\text{C}(\text{n},\gamma)$ E \approx 8-20 MeV; analyzed $\sigma(\theta)$ vs E; deduced GDR, resonance parameters. Unified formalism.
- 1994Ot04: $^{12}\text{C}(\text{n},\gamma)$ E<0.5 MeV; calculated $\sigma(E)$; deduced S-factor. Kinematically complete approach.
- 1995Li31: $^{12}\text{C}(\text{n},\gamma)$ E \approx 6-20 MeV; calculated capture $\sigma(\theta)$ vs E. Direct-semidirect model.
- 1995Me14: $^{12}\text{C}(\text{n},\gamma)$ E \leq 500 keV; calculated capture $\sigma(E)$. Direct capture model.
- 1996Re16: $^{12}\text{C}(\text{n},\gamma)$; analyzed inverse Coulomb dissociation reaction and relevance for astrophysical input.
- 1997Du09: $^{12}\text{C}(\text{n},\gamma)$ $E_{\text{c.m.}}\leq 0.5$ MeV; calculated capture $\sigma(E)$. Calculated levels, $B(\lambda)$, rms radius vs $R(\text{c})$. Multicenter approach.
- 1997Li10: $^{12}\text{C}(\text{n},\gamma)$ E<600 keV; calculated $\sigma(\text{En})$; deduced influence of scattering potential depth. Consistent direct-semidirect model.
- 1997Ti03: Analyzed vertex constants for capture reactions.
- 1998Ki01: $^{12}\text{C}(\text{n},\gamma)$ E<1 MeV; calculated σ ; deduced optical potential features.
- 1999MeZW: $^{12}\text{C}(\text{n},\gamma)$ E<0.8 MeV; analyzed capture σ ; deduced parameters.
- 2003Wu01: $^{12}\text{C}(\text{n},\gamma)$ E=0-1 MeV; calculated σ . Asymptotic normalization coefficient method, comparison with data.
- 2004Ba62: $^{12}\text{C}(\text{n},\gamma)$ E=0-0.6 MeV; calculated S-factors, $\sigma(E)$ for radiative capture. Taylor expansion.
- 2004Hu10: $^{12}\text{C}(\text{n},\gamma)$ E=low; calculated astrophysical reaction rate, resonance effects.
- 2009Wa17: $^{12}\text{C}(\text{n},\gamma)$ $E_{\text{c.m.}}< 1$ MeV; analyzed σ , spectroscopic factors and other parameters for nonresonant neutron capture using

$^{12}\text{C}(\text{n},\gamma)$:res (continued)

simple polynomials obtained from Taylor expansions. Comparison with experimental data.

[2010Hu11](#): $^{12}\text{C}(\text{n},\gamma)$ $E_{\text{c.m.}} < 2$ MeV; calculated binding energies, σ , S-factors, spectroscopic factors. Single-particle potential model.

[2012Pr13](#), [2020Pr08](#): $^{12}\text{C}(\text{n},\gamma)$ $E < 20$ MeV; calculated Maxwellian-averaged σ , astrophysical reaction rates, neutron thermal σ , Westcott factors, resonance integrals.

[2013Di12](#): $^{12}\text{C}(\text{n},\gamma)$ $E < 20$ MeV; analyzed available data; deduced recommended σ , k_{eff} .

[2013Du08](#): $^{12}\text{C}(\text{n},\gamma)$ $E < 1$ MeV; calculated σ , low-energy phase shifts. Potential cluster model, comparison with available data.

[2013Du16](#): $^{12}\text{C}(\text{n},\gamma)$ $E < 1$ MeV; calculated σ , phase shifts. Young diagrams, potential cluster model.

[2017HaZY](#): $^{12}\text{C}(\text{n},\gamma)$ $E = 0-0.2$ MeV; calculated σ .

[2018Br05](#): $^{12}\text{C}(\text{n},\gamma)$ $E = 30$ keV; calculated Maxwellian-averaged σ .

[2024Sa13](#): $^{12}\text{C}(\text{n},\gamma)$; R-matrix analysis of the astrophysically relevant region of the reaction.

 ^{13}C Levels

<u>E(level)[†]</u>	<u>J^π</u>	<u>Γ</u>	<u>Comments</u>
0	1/2 ⁻		
20520 70		510 keV 70	E(level),Γ: From the analyzing power data in (1984Wo05): secondary doorway state.
21050 60		4.2 MeV 4	E(level),Γ: From the analyzing power data in(1984Wo05): primary doorway state. See also $E_x \approx 20.81$ MeV and $\Gamma \approx 2.5$ MeV in (1986Be17).

[†] ([1986Be17](#)) also observed an unresolved broad peak at $E_n(\text{res}) \approx 10$ MeV ($E_x \approx 14180$ keV).

 $\gamma(^{13}\text{C})$

<u>E_γ[†]</u>	<u>$E_i(\text{level})$</u>	<u>E_f</u>	<u>J_f^π</u>
20.520×10^3	20520	0	1/2 ⁻
21.050×10^3	21050	0	1/2 ⁻

[†] From level energy difference.

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