

$^2\text{H}(^{12}\text{B},\text{p})$ **2010Ba06**

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$J^\pi(^{12}\text{B g.s.})=1^+$.

2010Ba06: XUNDL dataset compiled by TUNL, 2010.

A 75 MeV/nucleon beam of ^{12}B ions, produced by bombarding a cryogenic deuterium gas cell with ^{11}B ions at the ANL/ATLAS facility, impinged on a $73\text{ }\mu\text{g}/\text{cm}^2$ CD_2 target located at the HELIOS (HELical Orbit Spectrometer) target position. Reaction protons were emitted in the backwards direction and followed a single helical orbit in the 1.05 T axial magnetic field before reaching a barrel shaped array of position sensitive Si detectors that surrounded the incident beam axis. The forward moving ^{13}C ions were stopped in a $\Delta\text{E-E}$ telescope that covered $\theta_{\text{lab}}=0.5^\circ\text{--}2.8^\circ$.

The momentum of the emitted proton was determined and excited states were resolved with $\Delta\text{E}\approx 100\text{ keV}$ FWHM. The angular distribution for population of $^{13}\text{B}(3.48, 3.68\text{ MeV})$ was determined over $\theta_{\text{c.m.}}=8^\circ\text{--}30^\circ$ by analyzing the $\text{p}+^{13}\text{B}$ coincidences. The angular distributions were analyzed via DWBA analysis.

2010Le02: XUNDL dataset compiled by TUNL, 2010.

A 75 MeV/nucleon beam of ^{12}B ions, from the ANL/ATLAS facility, impinged on a $150\text{ }\mu\text{g}/\text{cm}^2$ CD_2 target. A set of three position-sensitive annular Si detectors measured protons at $\theta_{\text{lab.}}=110^\circ\text{--}161^\circ$ while forward moving boron isotopes were identified in a $\Delta\text{E-E}$ telescope that covered $\theta_{\text{lab}}=1.3^\circ\text{--}7.2^\circ$. Neutron bound and unbound states of ^{13}B were identified at $E_x=0, 3.48, 3.68, 5.105, 5.388\text{ MeV}$; only the ground state was resolved.

The angular distribution was determined for the ground state over $\theta_{\text{c.m.}}=7.5^\circ\text{--}30^\circ$, and it was analyzed via DWBA analysis to obtain spectroscopic data useful for determining the astrophysical $^{12}\text{B}(\text{n},\gamma)$ reaction rates. Also see [2008WuZY](#), [2011BaZX](#) for other ANL reports.

See ([2021Du10](#)) for a calculation of the cross section at astrophysically relevant energies.

 ^{13}B Levels

| E(level) [†] | J^π [#] | L [#] | S [#] | Comments |
|--------------------------------|----------------------|----------------|----------------|--|
| 0 | $3/2^-$ | 1 | 1.1 3 | |
| 3.48×10^3 [‡] | $(1/2^+)$ | 0,2 | | $S_{L=2} \leq 0.05 S_{L=0}$ component (2010Ba06). |
| 3.68×10^3 [‡] | $(5/2^+, 3/2^+)$ | 2,0 | | The L=0 component is less than $\approx 2\%$ of the L=0 component for the 3.48 MeV state. The authors suggest that $J^\pi=5/2^+$ is favored based on absence of L=0 component in the angular distribution and a better fit to the ratios of spectroscopic factor (2010Ba06). |
| 5105 [‡] | | | | |
| 5388 [‡] | | | | |

[†] Nominal values given in ([2010Ba06](#), [2010Le02](#)).

[‡] Unresolved in ([2010Le02](#)).

[#] From DWBA analysis of spectroscopic factors in ([2010Le02](#)).