

$^9\text{Be}(p,\gamma)$ res 1988Aj01

Type	Author	History	Citation	Literature Cutoff Date
Full Evaluation	J. H. Kelley, C. G. Sheu and J. L. Godwin, et al.		NP A745 155 (2004)	31-Mar-2004
<p>1964Bi18: $^9\text{Be}(P,\gamma)$ E=336 keV, measured $\sigma(E_\gamma, \theta_\gamma, \theta(\gamma-\gamma))$. ^{10}B level deduced J, π. 1968Si07: $^9\text{Be}(P,\gamma)$ E=2-2.1 MeV. 1970BeYS: $^9\text{Be}(P,\gamma)$ E not given, measured $\gamma-\gamma(\theta, H)$. ^{10}B level deduced g. 1972Av01: $^9\text{Be}(P,\gamma-\gamma)$ E=1.084 MeV, H=17.6 Kg, measured $\gamma-\gamma(\theta, H)$. ^{10}B deduced g-factor 1^+ first excited state. 1972Ha63: $^9\text{Be}(P,\gamma)$ E=1.083 MeV, measured E_γ. ^{10}B deduced γ-branching, isospin impurity fo 6.88-MeV state. 1972Re07: $^9\text{Be}(P,\gamma)$ E=0.20-0.85 MeV, measured $\sigma(E)$. ^{10}B deduced γ-branching, isospin impurity of 6.88-MeV state. 1973Sz07: $^9\text{Be}(P,\gamma)$ E=200-750 keV, measured $\sigma(E, E_\gamma)$. 1975Au02: $^9\text{Be}(P,\gamma)$ E=0.22-1.2 MeV, measured $\sigma(E, E_\gamma, \theta=0$ degree). ^{10}B deduced levels Γ-level, J, π, T. 1977Ki17: $^9\text{Be}(P,\gamma)$ E=1.29 MeV, measured $\sigma(E_\gamma)$, Doppler broadening. ^{10}B level deduced $T_{1/2}$. 1982Ri04: $^9\text{Be}(P,\gamma)$ E=320 keV, measured $\sigma(E_\gamma)$, I_γ. ^{10}B deduced isoscalar E2 transition strength upper limit, γ-branching ratio. 1985Ki07: $^9\text{Be}(P,\gamma)$ E=2.4-4.2 MeV, measured thick target relative γ yields, E_γ, I_γ. 1986CaZV: $^9\text{Be}(P,\gamma)$ E=193-407 keV, measured γ yield vs. $E, \gamma(\theta)$. ^{10}B deduced isospin mixed state characteristics. 1987Ra23: $^9\text{Be}(P,\gamma)$ E=7-9 MeV, measured absolute thick target γ yield, relative neutron yield. 1992Ce02: $^9\text{Be}(P,\gamma)$ E=40-180 keV, measured capture E_γ, I_γ, $\gamma(\theta)$. Deduced astrophysical S-factor. ^{10}B levels deduced γ-ray to charged particle branching ratio. 1995Za04: $^9\text{Be}(P,\gamma)$ E=75-1800 keV, measured E_γ, I_γ. Deduced $\sigma(E)$, thermonuclear reaction rates, astrophysical S-factor. 1998Wu05: $^9\text{Be}(\text{pol. } p,\gamma)$ E=100 keV, measured E_γ, I_γ, $A_Y(\text{THETA})$. Deduced astrophysical S-factor. 1999Ga21: $^9\text{Be}(\text{pol. } p,\gamma)$ E=280 keV, measured $A_Y(\text{THETA})$. Deduced ground-state capture S-factor. ^{10}B level deduced J, π. 2002Ba09: $^9\text{Be}(P,\gamma)$ E=75-1800 keV, analyzed $\sigma(E)$, $A_Y(\text{THETA})$. Deduced R-matrix parameters. 2002Ga11: $^9\text{Be}(P,\gamma)$ $E_{C.M.}=0-1.6$ MeV, analyzed S-factor.</p>				

 ^{10}B Levels

E(level)	J^π	$T_{1/2}$	$\omega-\gamma$ (eV)	Comments
0				
720				
1.74×10^3				
2.15×10^3				
3.59×10^3		133 fs 21		Γ : from $T_{\text{mean}}=192$ fs +35-29 (1977Ki17).
5.11×10^3				
5.16×10^3				
5.18×10^3				
5.92×10^3				
6873 5	1^-	120 keV 5	0.14 4	$\Gamma_p/\Gamma=0.30$; $\Gamma_\gamma=4.8$ eV E(level): Γ : from $E_p=319$ keV 5 (1975Au02).
7427 5	$2^{(-)}$	140 keV 30	1.25 18	$\Gamma_p/\Gamma=0.7$; $\Gamma_\gamma=2.4$ eV E(level): from $E_{\text{res}}=934$ keV 6, the weighted average of $E_p=932$ keV 8 (1962El06) and 938 keV 10 (1964Ho02). Γ : from (1964Ho02).
7472? 5	$(1^+, 2^+)$	≈ 80 keV		E(level): Γ : from $E_p=985$ keV 5 (1962El06).
7480 2	2^-	72 keV 4		$T=1$; $\Gamma_p/\Gamma \approx 0.65$; $\Gamma_\gamma=25.6$ eV E(level): from $E_p=993$ keV 2, the weighted average of $E_p=992$ keV 2 (1964Ho02) and $E_{\text{res}}=998$ keV 5 (1962El06). Γ : from (1964Ho02). Also see (1962El06) $\Gamma \approx 80$ keV.
7559.9 4	0^+	2.65 keV 18		$T=1$; $\Gamma_p/\Gamma=1.0$; $\Gamma_\gamma=8.5$ eV E(level): from $E_p=1083.2$ keV 4 (1964Bo13). Γ : from $\Gamma_{\text{lab}}=2.94$ keV 20 (1972Ha63). Other

Continued on next page (footnotes at end of table)

$^9\text{Be}(p,\gamma)$ res 1988Aj01 (continued) ^{10}B Levels (continued)

<u>E(level)</u>	<u>J$^\pi$</u>	<u>T$_{1/2}$</u>	<u>Comments</u>
			discrepant values are $\Gamma_{\text{C.M.}}=3.25$ keV 33 (1964Bo13) and $\Gamma_{\text{C.M.}}=3.4$ keV 5 (1964Ho02).
7.75×10^3 3	2^-	210 keV 60	$\Gamma_p/\Gamma \approx 0.65$; $\Gamma_\gamma = 8.5$ eV E(level): Γ : from $E_p = 1290$ keV 30 (1964Ho02). T=(1).
8894 2	2^+	36 keV 2	T=1 E(level): Γ : from $E_p = 2567$ keV 2 (1956Ma55) and (Mackin Thesis, Cal Tech, 1953 see (1974Aj01)).
10.83×10^3	$2^+, 3^+, 4^+$	≈ 500 keV	E(level): from $E_p = 4.72 \times 10^3$ keV (1952Ha10) and (Fisher Thesis, Stanford 1970, see (1974Aj01)). Γ : from (1952Ha10).
12.6×10^3	$0^+, 1^+, 2^+$	< 200 keV	E(level): Γ : from $E_p = 6.7 \times 10^3$ keV (Fisher Thesis, Stanford 1970, see (1974Aj01)).
$12.9 \times 10^3?$	$+$	≈ 100 keV	E(level): Γ : from $E_p = 7.0 \times 10^3$ keV (Fisher Thesis, Stanford 1970, see (1974Aj01)).
13.3×10^3	$0^+, 1^+, 2^+$	≈ 300 keV	E(level): Γ : from $E_p = 7.5 \times 10^3$ keV (Fisher Thesis, Stanford 1970, see (1974Aj01)).
14.1×10^3	$0^+, 1^+, 2^+$	≈ 250 keV	E(level): Γ : from $E_p = 8.4 \times 10^3$ keV (Fisher Thesis, Stanford 1970, see (1974Aj01)).
$14.6 \times 10^3?$	$2^+, 3^+, 4^+$	≈ 150 keV	E(level): Γ : from $E_p = 8.9 \times 10^3$ keV (Fisher Thesis, Stanford 1970, see (1974Aj01)).
$15.6 \times 10^3?$	$2^+, 3^+, 4^+$	≈ 400 keV	E(level): Γ : from $E_p = 10.0 \times 10^3$ keV (Fisher Thesis, Stanford 1970, see (1974Aj01)).
$19.7 \times 10^3?$	$2^-, 3^-, 4^-$	≈ 50 keV	E(level): Γ : from $E_p = 14.6 \times 10^3$ keV (Fisher Thesis, Stanford 1970, see (1974Aj01)).