

First results on assimilation of major actinides

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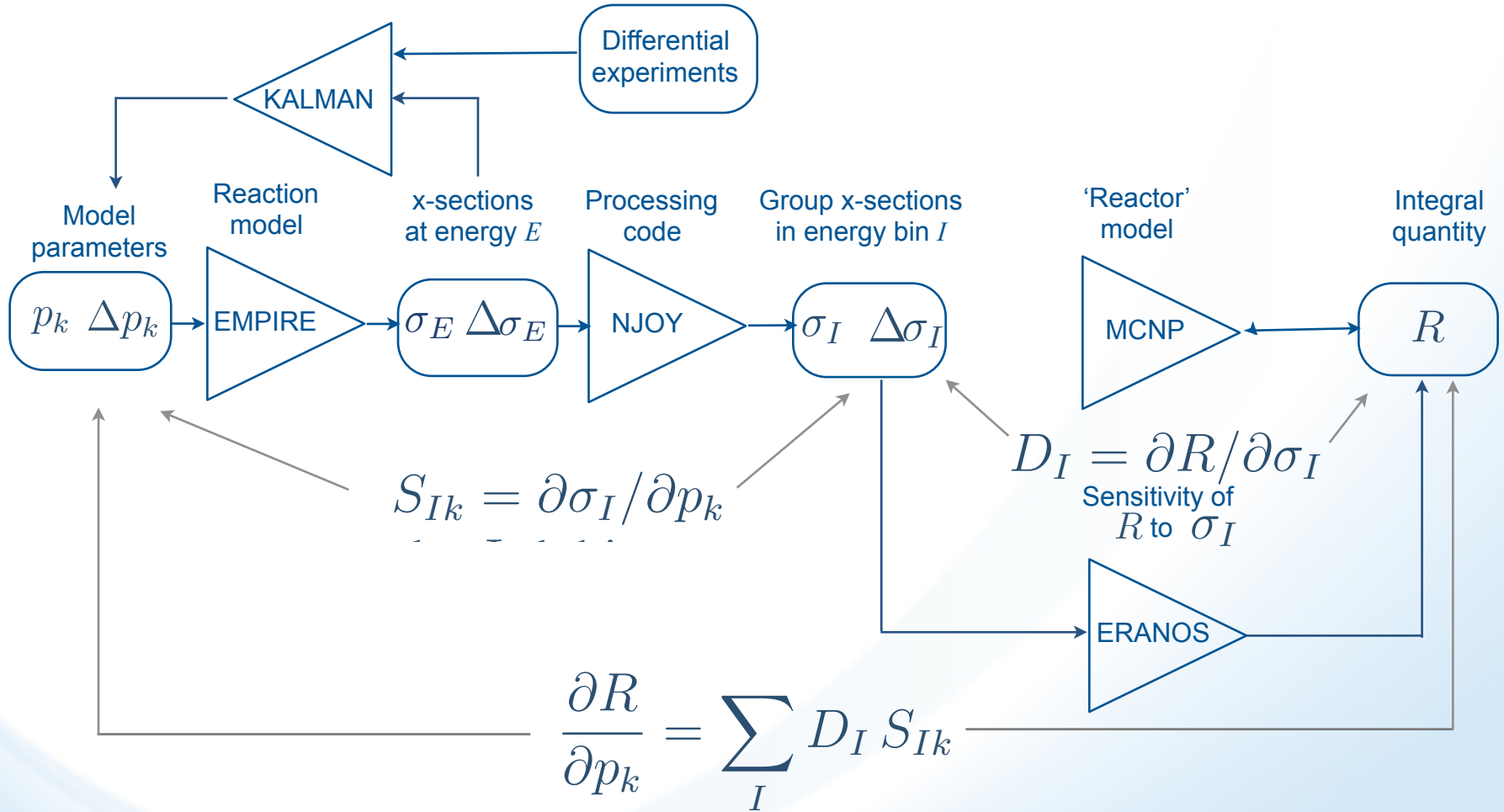


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Assimilation

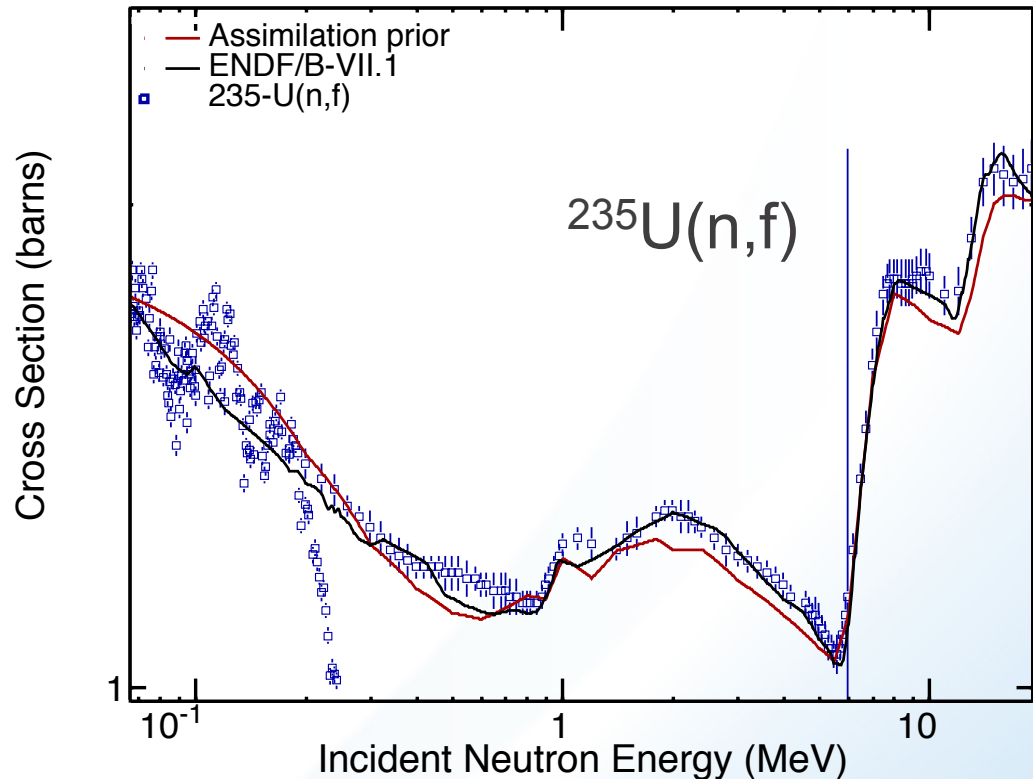
Linking integral experiments with reaction model parameters



Using S_{Ik} differential exp. data, and Kalman filter $\implies \langle \Delta p_k \Delta p_l \rangle$ covariance matrix, which contains constraints imposed by microscopic exp. data.

^{235}U assimilation starting point (prior)

- EMPIRE calculations
 - EGSM lev. den.
 - CC with RIPL #2408
4 coll. lev. & 76 in continuum
 - Exciton model
 - E1 strength MLO1
 - OM for fission
 - Fission barriers RIPL1
- Roughly adjusted parameters but **no** energy-dependent tuning!
- PFNS and nu-bars from VII.1
 - **Input to INL**
 - central values in 33 energy groups
 - group-wise sensitivity matrices for 50 parameters (no PFNS & nu-bars!)
 - parameter covariance matrix



Performance of the prior file

k_{eff} results (experimental $k_{\text{eff}}=1.0$ ($\pm 100\sim 300\text{pcm}$))

Experiment	EMPIRE ($\pm\text{pcm}$)	ENDF/B-VII.0 ($\pm\text{pcm}$)
JEZEBEL-239	0.98567 (± 8)	0.99986 (± 9)
GODIVA	0.99072 (± 9)	0.99983 (± 9)
FLATTOP-Pu	0.98838 (± 18)	1.00097 (± 18)
FLATTOP-25	1.00182 (± 17)	1.00217(± 17)

C/E ratio of spectral indices at the center of JEZEBEL-239 and GODIVA

	JEZEBEL-239		GODIVA	
	EMPIRE	ENDF/B-VII.0	EMPIRE	ENDF/B-VII.0
$\sigma_f(^{238}\text{U})/\sigma_f(^{235}\text{U})$	0.956 \pm 0.009	0.974 \pm 0.009	1.053 \pm 0.013	0.954 \pm 0.012
$\sigma_f(^{233}\text{U})/\sigma_f(^{235}\text{U})$	1.000 \pm 0.017	0.986 \pm 0.017	0.996 \pm 0.019	0.987 \pm 0.019
$\sigma_f(^{237}\text{Np})/\sigma_f(^{235}\text{U})$	0.999 \pm 0.017	1.009 \pm 0.017	1.070 \pm 0.017	0.990 \pm 0.016
$\sigma_f(^{239}\text{Pu})/\sigma_f(^{235}\text{U})$	0.971 \pm 0.020	0.984 \pm 0.020	0.992 \pm 0.018	0.986 \pm 0.018

Assimilation of the ^{235}U data

C/E for GODIVA

Experiment	old C/E $\pm \sigma$	new C/E $\pm \sigma$
K_{eff}	0.9907 ± 0.002	1.0010 ± 0.002
Fis. ^{238}U /Fis. ^{235}U	1.0527 ± 0.013	1.0357 ± 0.004
Fis. ^{239}Pu /Fis. ^{235}U	0.9917 ± 0.018	0.9771 ± 0.003
Fis. ^{237}Np /Fis. ^{235}U	1.0703 ± 0.017	1.0536 ± 0.003
Fis. ^{233}U /Fis. ^{235}U	0.9964 ± 0.019	0.9820 ± 0.004

- a) Factor multiplying the reaction (fusion, absorption, compound nucleus formation) cross section,
 b) Factor multiplying total cross section,
 c) Asymptotic level density parameter in Compound Nucleus, d) Pairing energy in the level densities at the saddle point in Compound Nucleus (first chance fission),
 e) Height of the second hump in the fission barrier in Compound Nucleus, f) Real depth of the Optical Model potential for n + target, g) Surface imaginary Optical Model potential radius for n + target, h) Surface imaginary Optical Model potential depth for n + target,
 i) Factor on the gamma emission width in Compound Nucleus (scales capture).

Parameter variations and standard deviations obtained by data assimilation

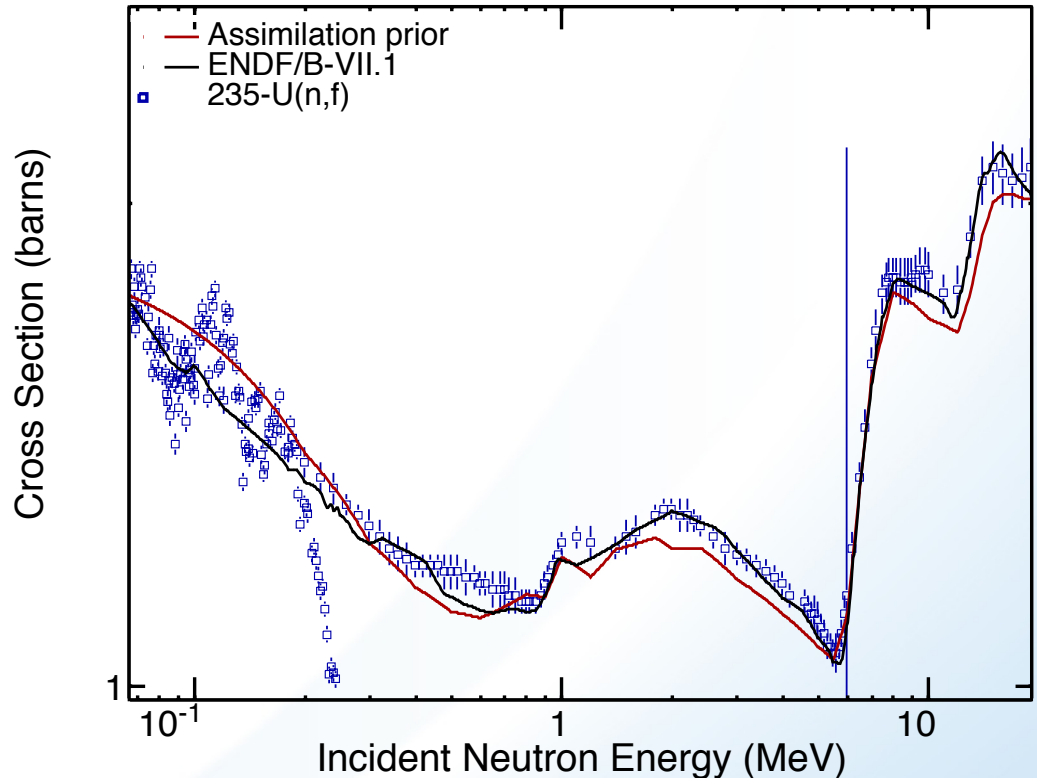
Parameter	Variation (%)	Init. Stand. Dev. (%)	Final Stand. Dev. (%)
FUSRED a)	1.402	1.257	0.878
TOTRED b)	0.461	0.966	0.917
ATILNO c)	-0.236	0.950	0.946
DELTAf d)	-0.025	0.649	0.621
VB0 e)	-0.006	0.133	0.118
UOMPVV101 f)	0.033	0.116	0.116
UOMPRS101 g)	0.072	0.834	0.834
UOMPWS101 h)	-0.110	2.023	2.022
TUNE000000 i)	-0.099	1.908	1.908

Verification - EMPIRE/MCNP calculations using posterior parameters

k_{eff} C/E for GODIVA

	prior C/E $\pm \sigma$	posterior C/E $\pm \sigma$
INL	0.9907 \pm 0.002	1.0010 \pm 0.002
BNL	0.98418 \pm 0.00008	0.99526 \pm 0.00008

INL and BNL calculations show ~ 1000 pcm improvement in spite of the difference in the starting values.

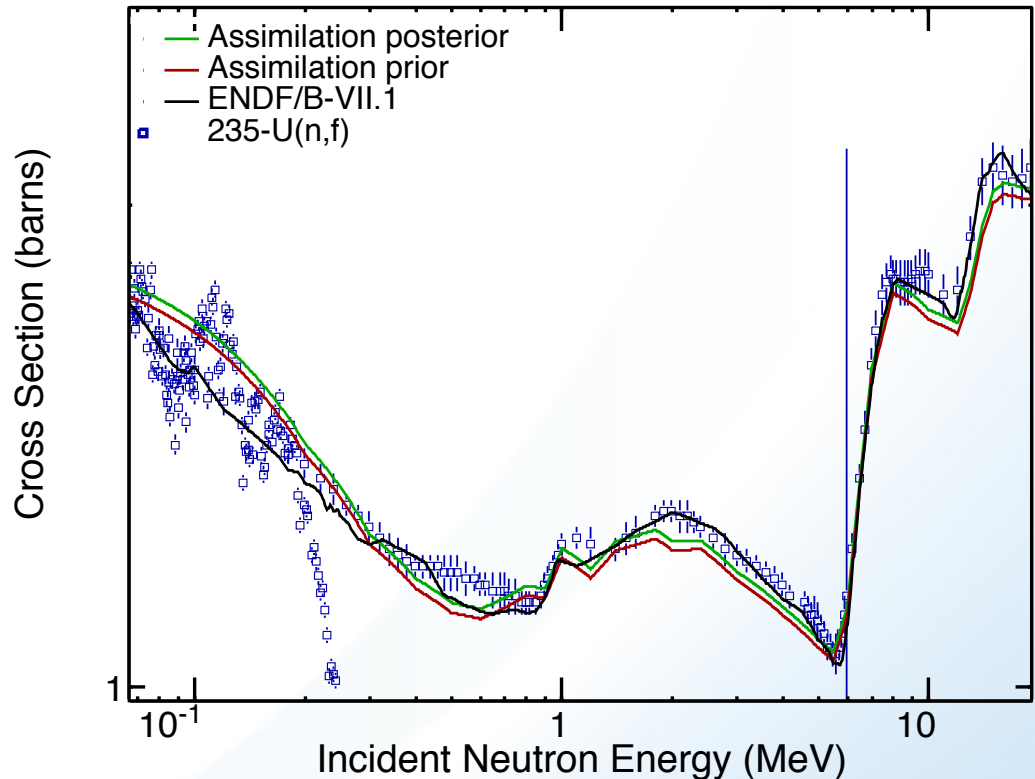


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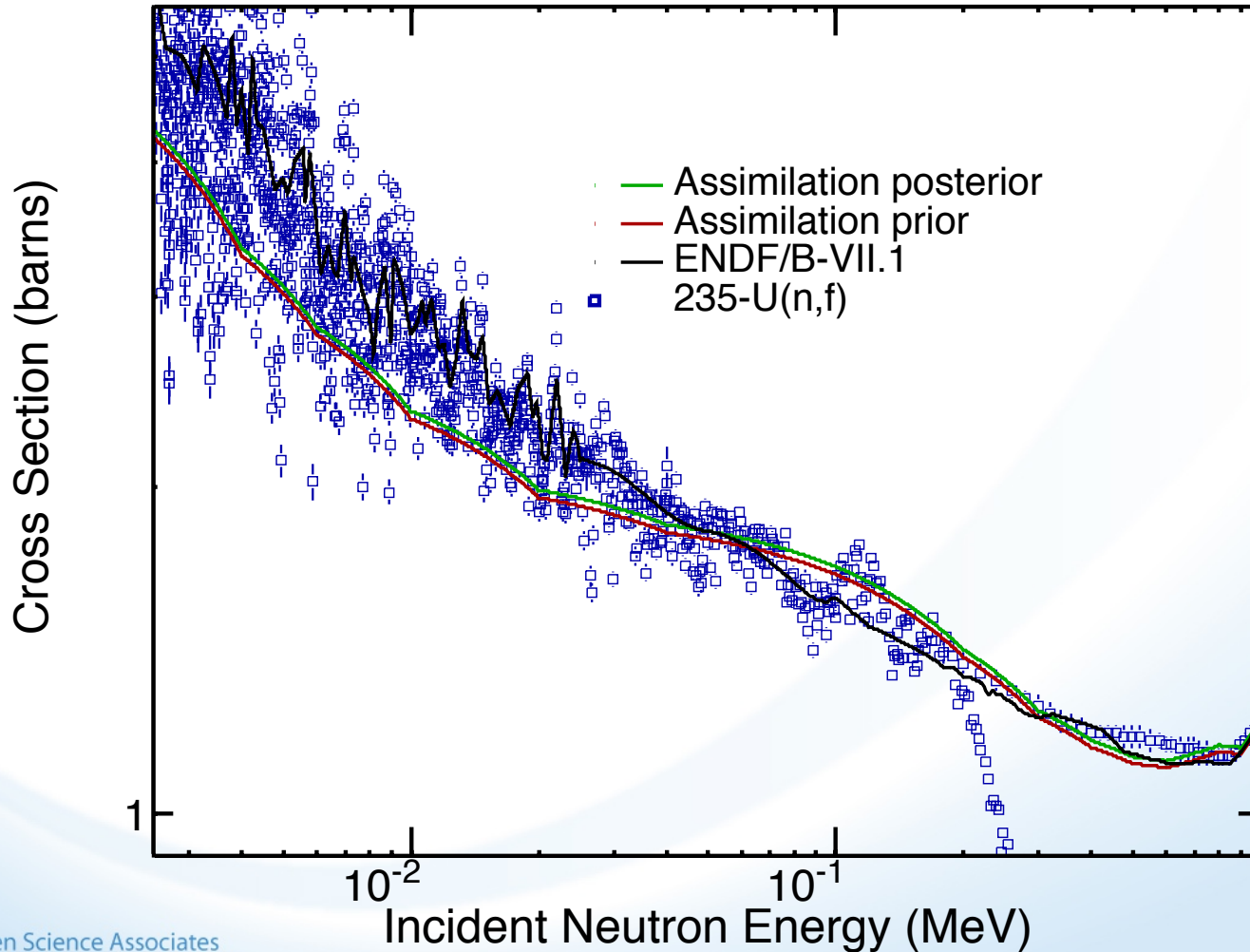
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Further developments

- Better 'a priori' calculations
- Developing PFNS capabilities in EMPIRE
- Allowing variations of PFNS and nu-bars
- Allowing variations of resonance parameters
- Including more integral experiments

Requisites for assimilation

- Adequate set of reaction models
- Entire evaluation expressed in terms of model parameters
- Reaction model and its parameterization flexible enough to reproduce differential and integral data
- Clean, well defined, integral experiments predominantly sensitive to a single material.

Conclusions

- Assimilation is feasible!
- Very small changes in cross sections can be enough to fix k_{eff}
- Pretty 'bad' cross sections can still produce reasonable k_{eff} , i.e., integral data should not overwrite differential data, lot of space for error compensation
- Non-linearity must be kept under control
- Advantages of consistent assimilation over Total MC
 - provides better insight into physics (sensitivities)
 - does not make 'unnecessary violence' to the parameters
 - calculation time - can be used for multi-material and multi-experiment adjustment
- Total MC advantages
 - not affected by non-linearity issues
 - can be used to find other local minima