

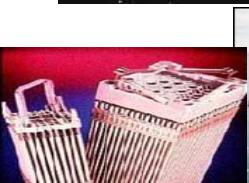
New Neutron Cross-Section Measurements from ORELA

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Nuclear Astrophysics

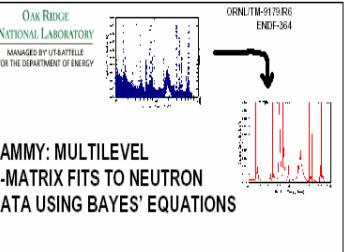
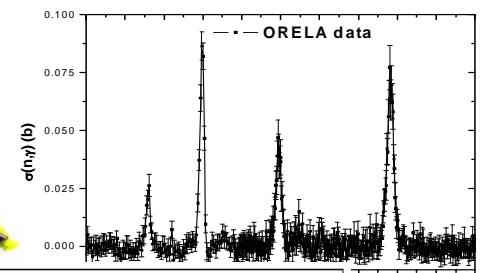


Applications

ORELA



Basic Science



SAMMY

Cross-Section Evaluations

ORNL Data Support for Nuclear Applications

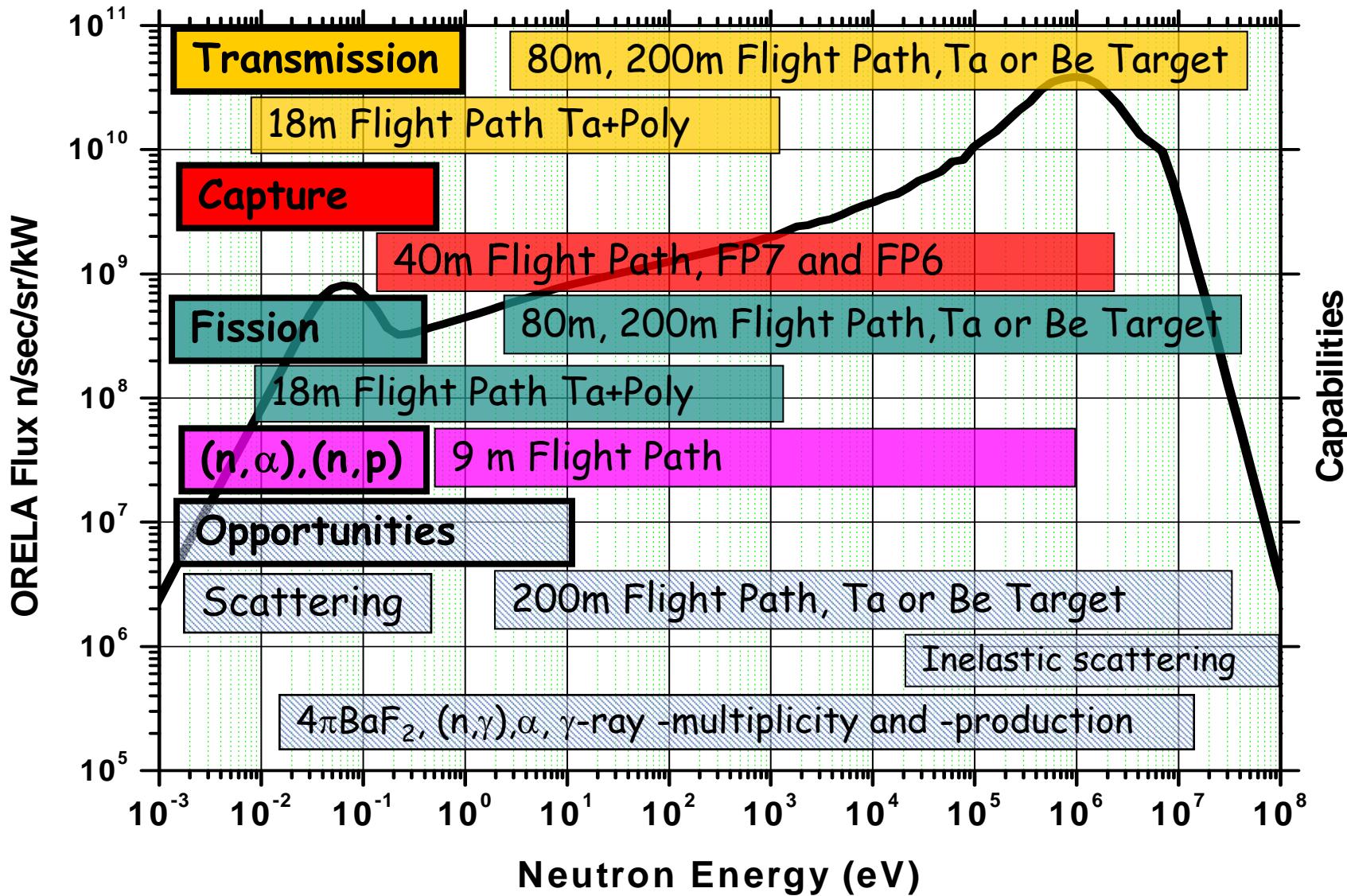
Computational modeling

SCALE

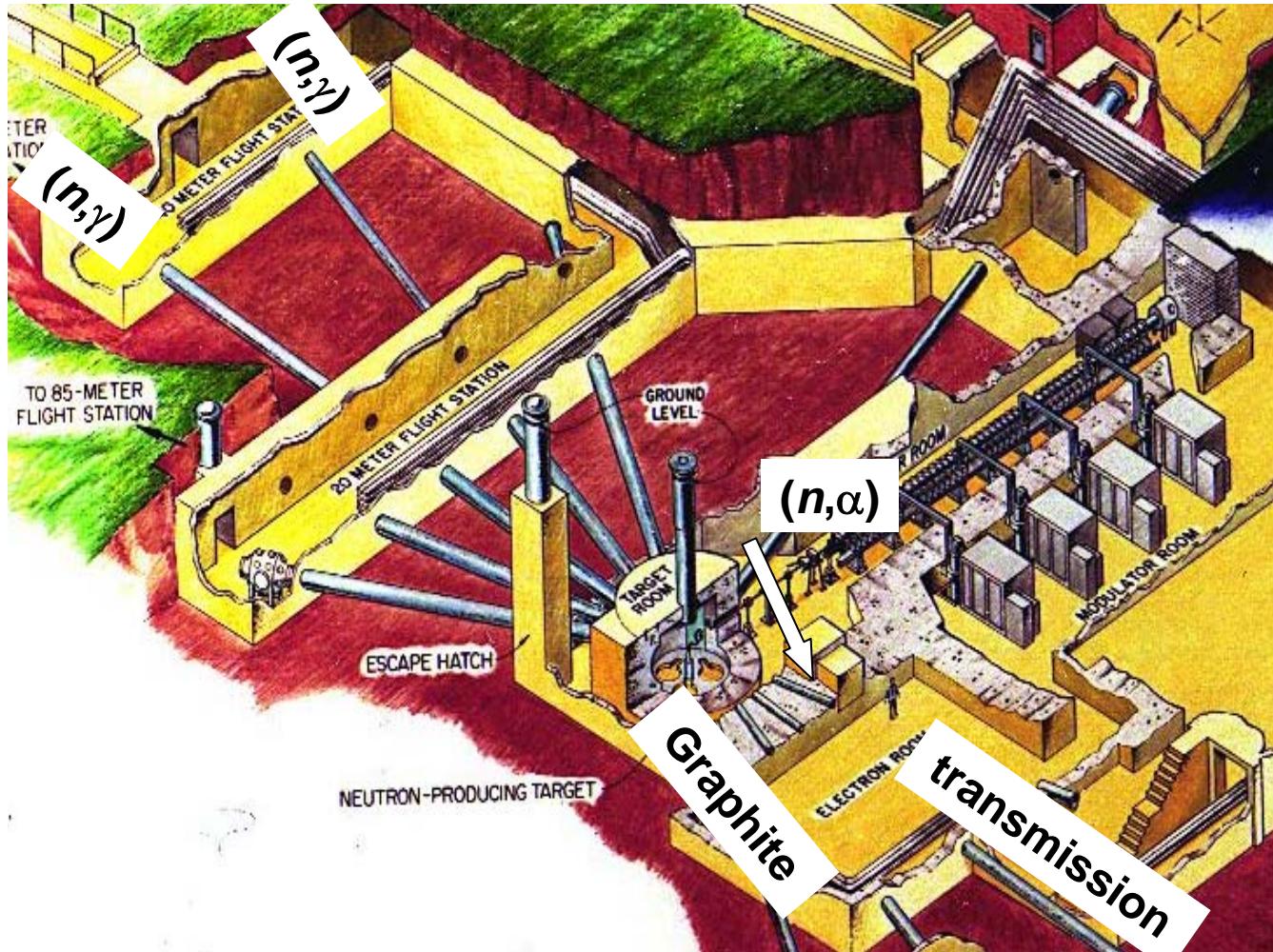
OAK RIDGE NATIONAL LABORATORY
U. S. DEPARTMENT OF ENERGY

AMPX

UT-BATTELLE



Existing Experiments at ORELA



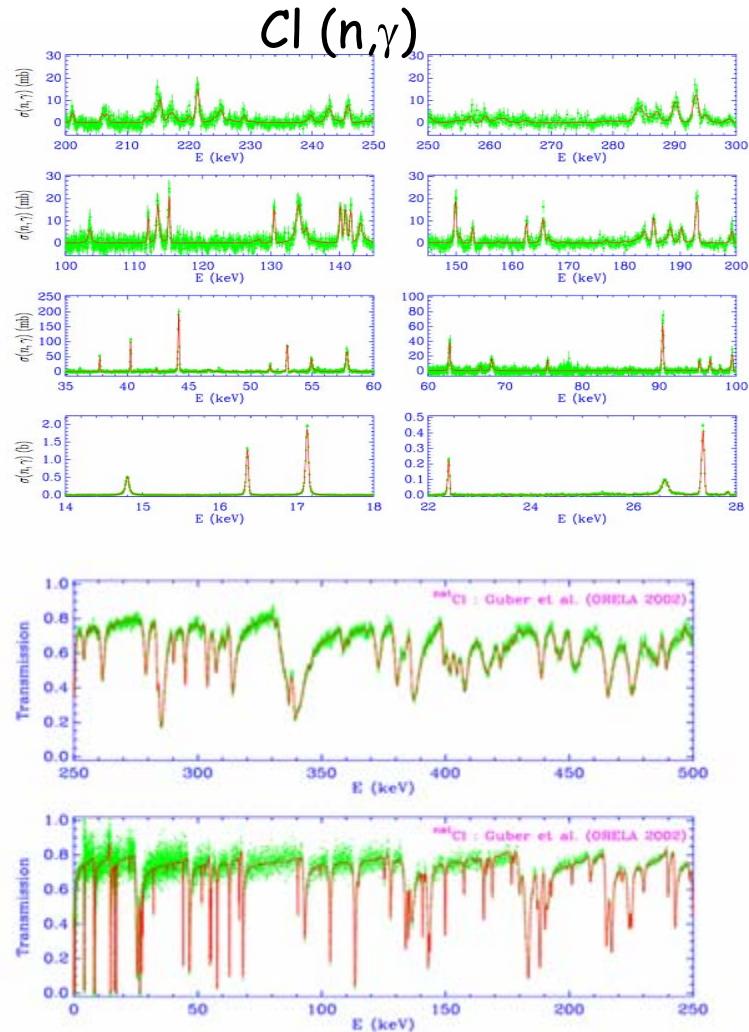
- 11 Flight paths
- Flight Stations:
 - 8-18, 20, 35, 40, 85, 150, and 200 m

New NCSP Measurements

- Completed ^{41}KCl capture and transmission.
- Measured Mn capture using thick sample (0.018at/b).
- Started natural Cr transmission (energy range 100eV to 50 keV), good high energy data available. Completed capture on natural Cr.
- Started capture on ^{58}Ni , good transmission data available.

New ORNL Evaluation for Cl helps to extract ^{41}K parameters

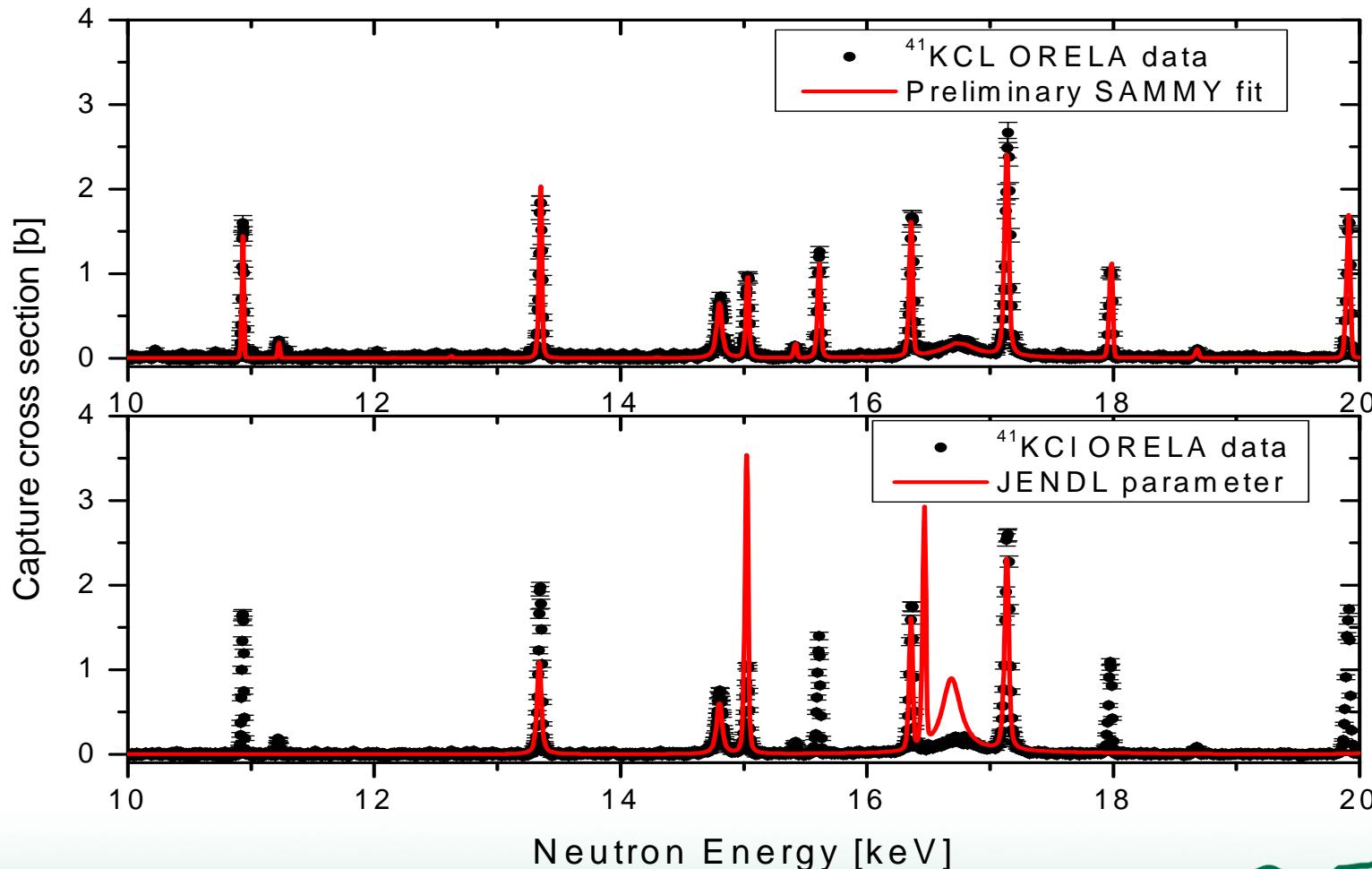
- ^{41}KCl sample used with 99.17% enrichment. (0.00797 at/b)
- Including the new resonance parameter set from the ORNL Cl evaluation it will be possible to extract reliable parameter for ^{41}K .



Cl transmission

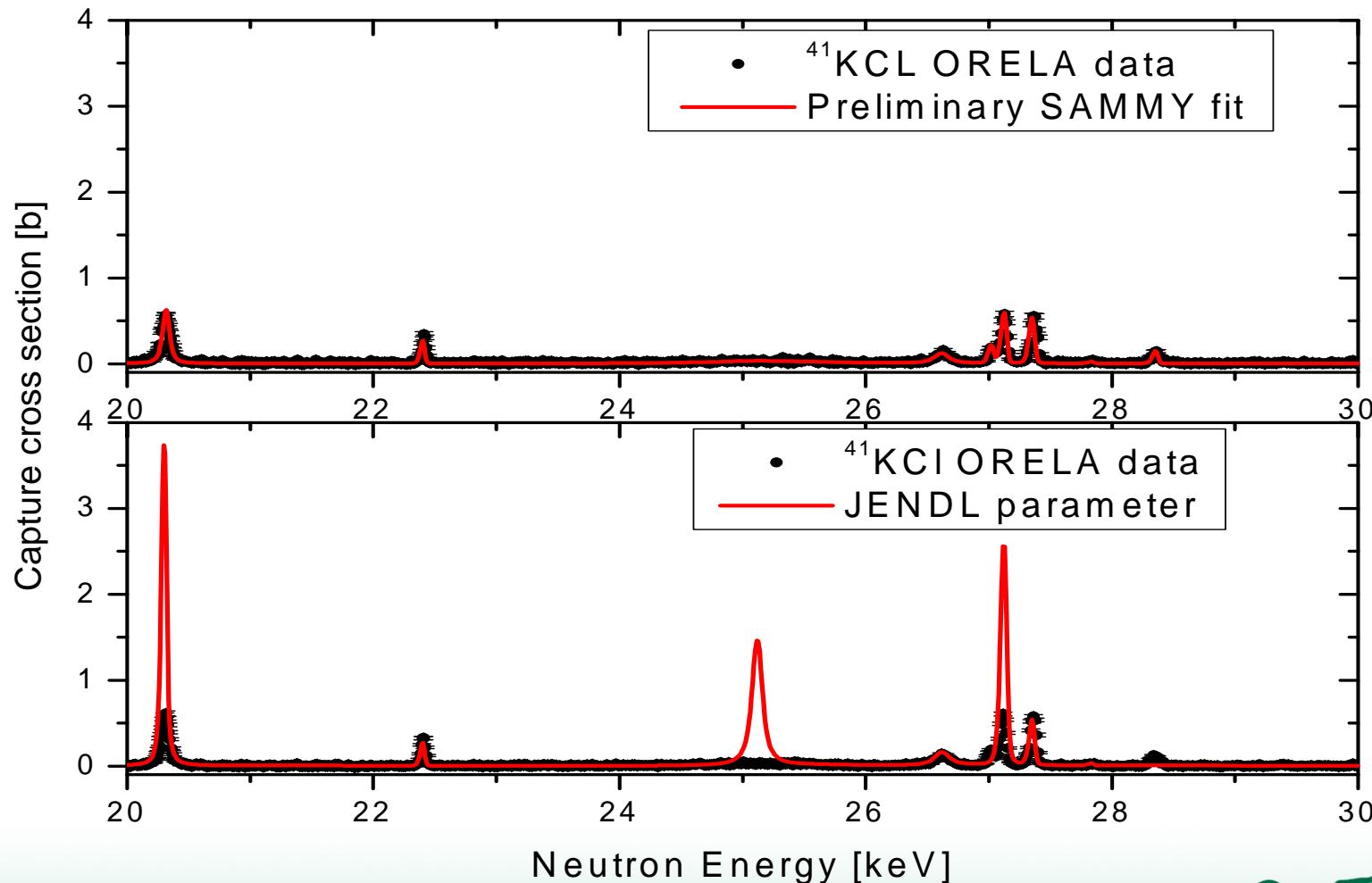
ORELA capture data for ^{41}KCl compared to JENDL3.3 and Preliminary SAMMY fits.

Several resonance areas too large (neutron sensitivity) in evaluation



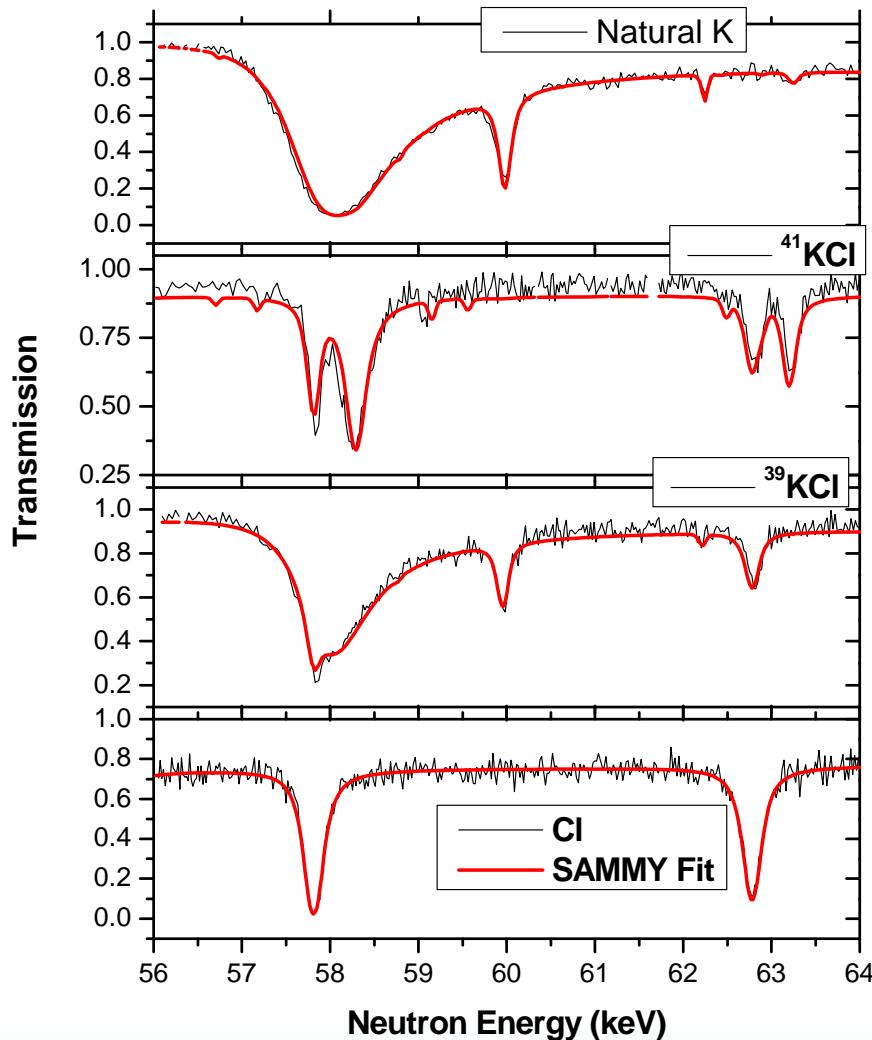
ORELA capture data for ^{41}KCl compared to JENDL3.3 and Preliminary SAMMY fits.

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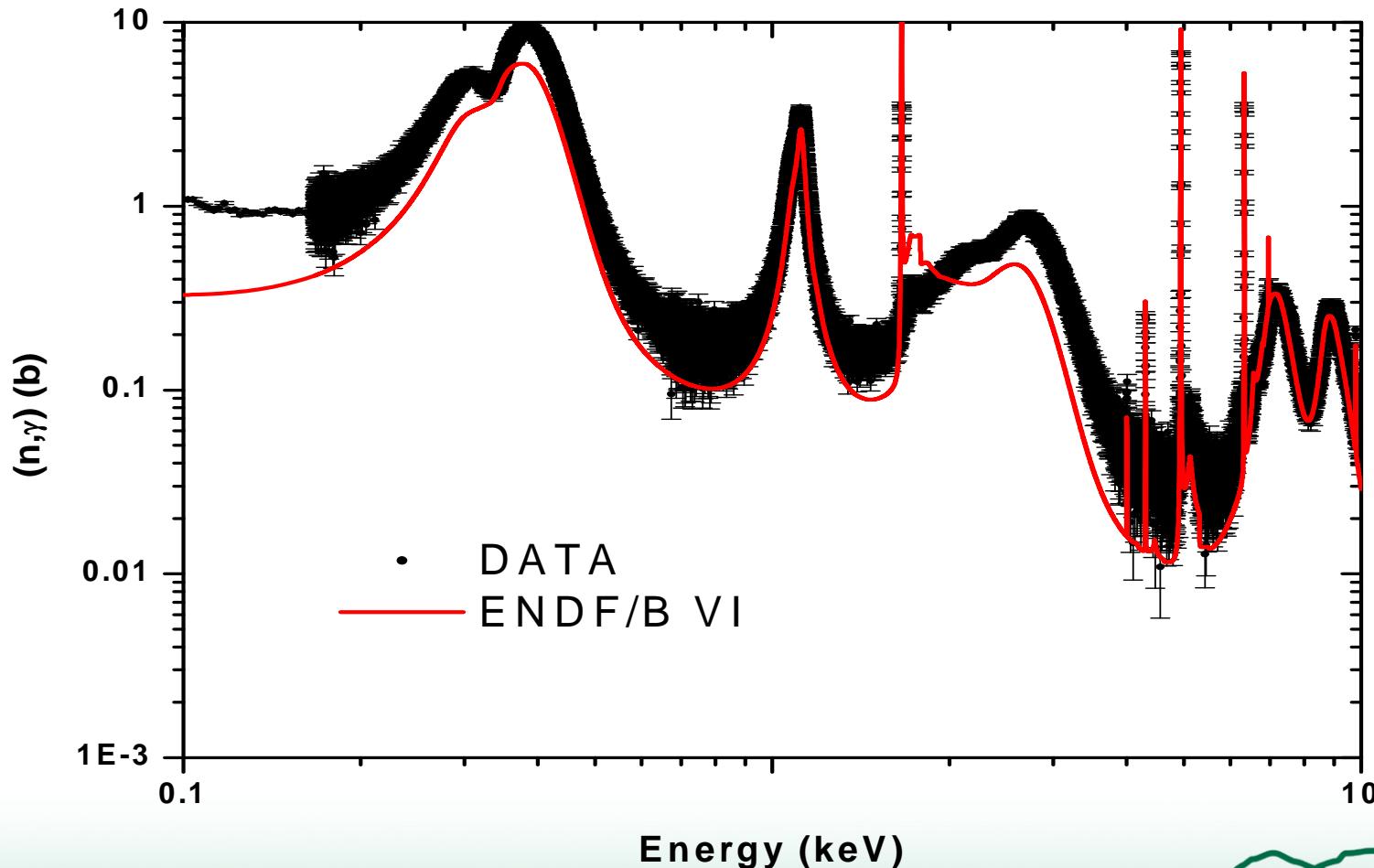
Preliminary fit to the Transmission data of metallic K and $^{41,39}\text{KCl}$

- Preliminary SAMMY fits to the transmission data of the $^{39,41}\text{KCl}$ samples and natural metallic K.
- With the use of the resonance parameter from the most recent Cl evaluation (Sayer et. al. Phys. Rev. C) it is possible to extract reliable K parameters.

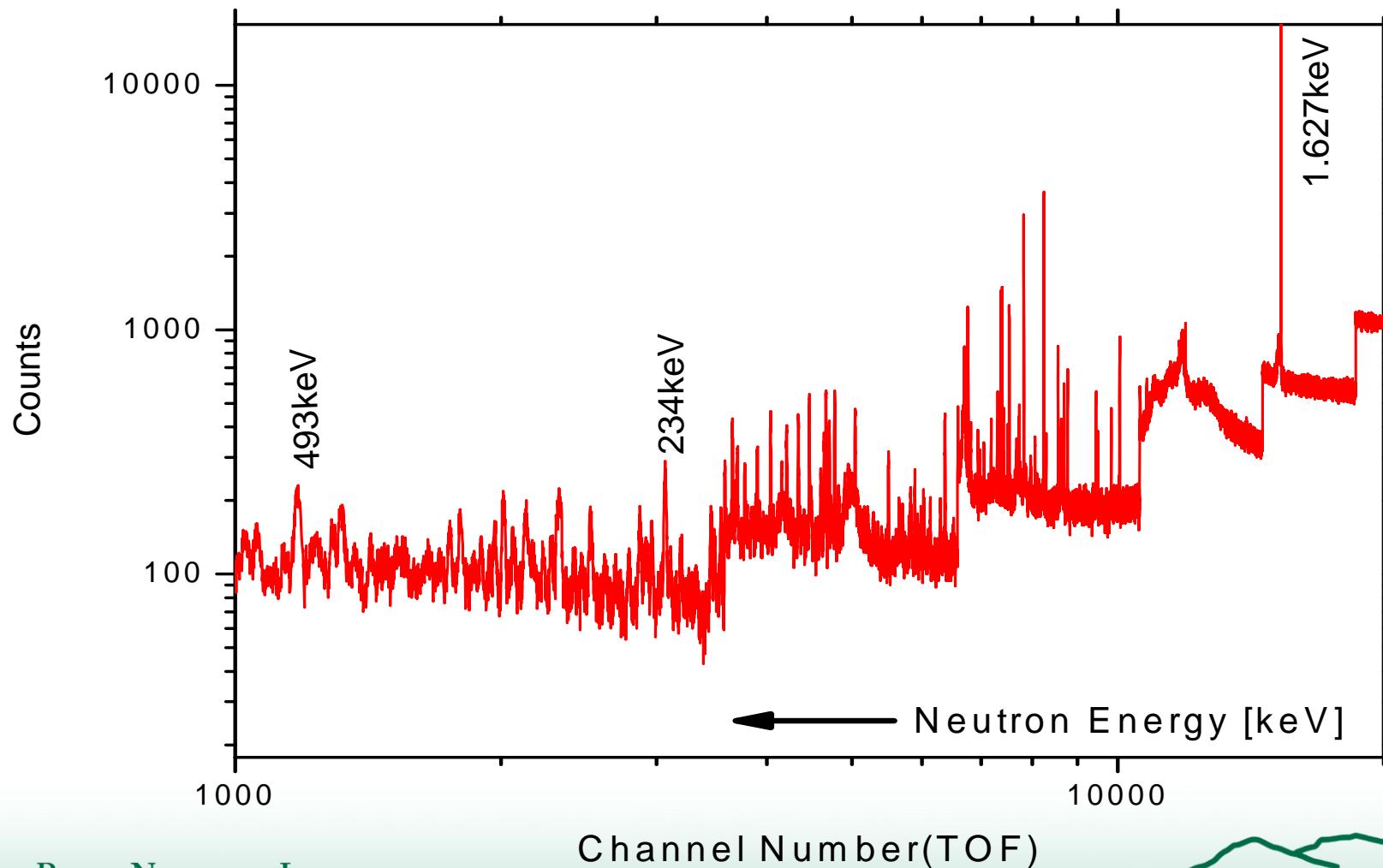


New ORELA Mn Neutron Capture compared to ENDF/B VI Evaluation

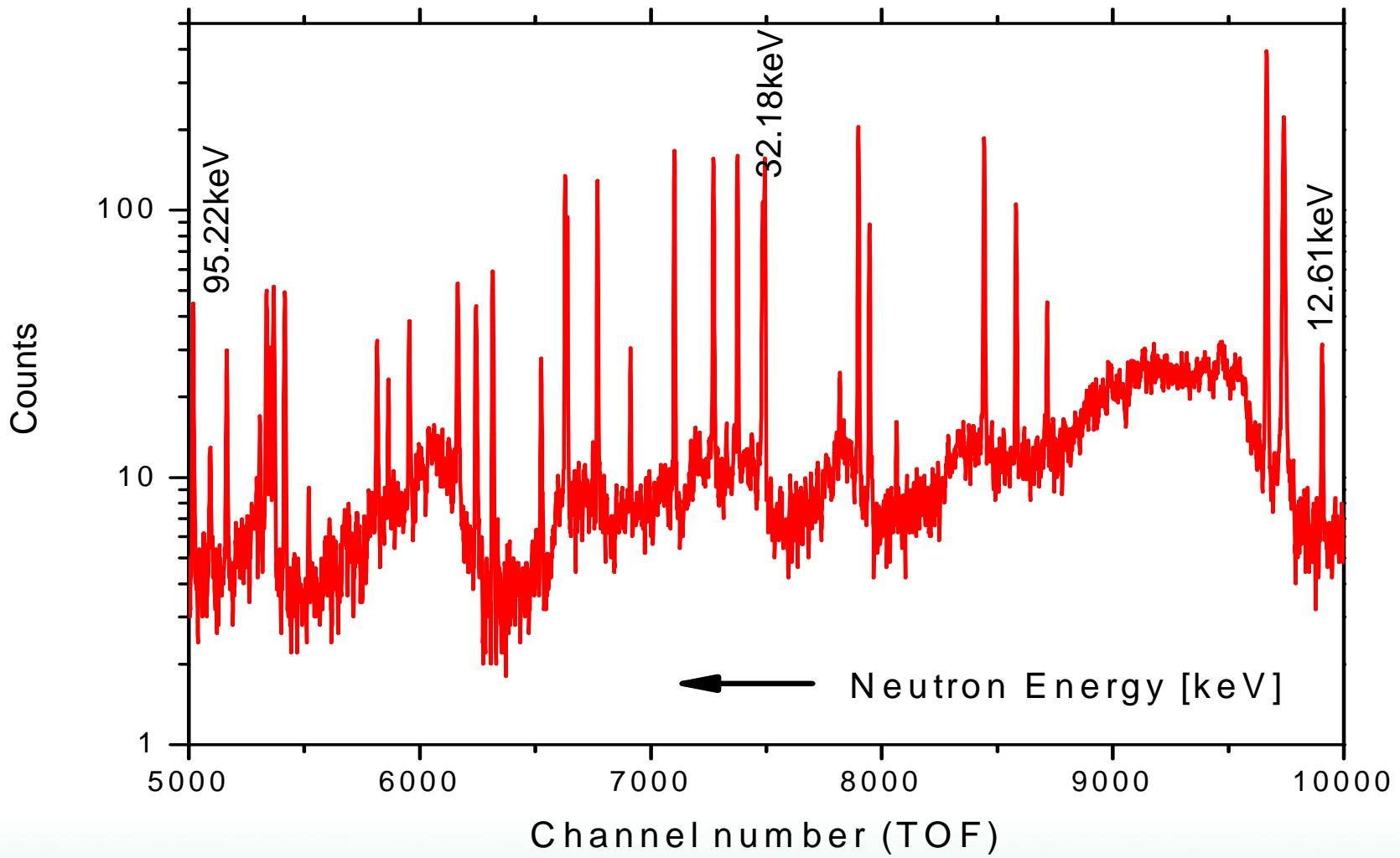
Very thick sample used, multiple scattering effect enormous



New ORELA Nat.Cr Capture TOF data; using a metallic sample

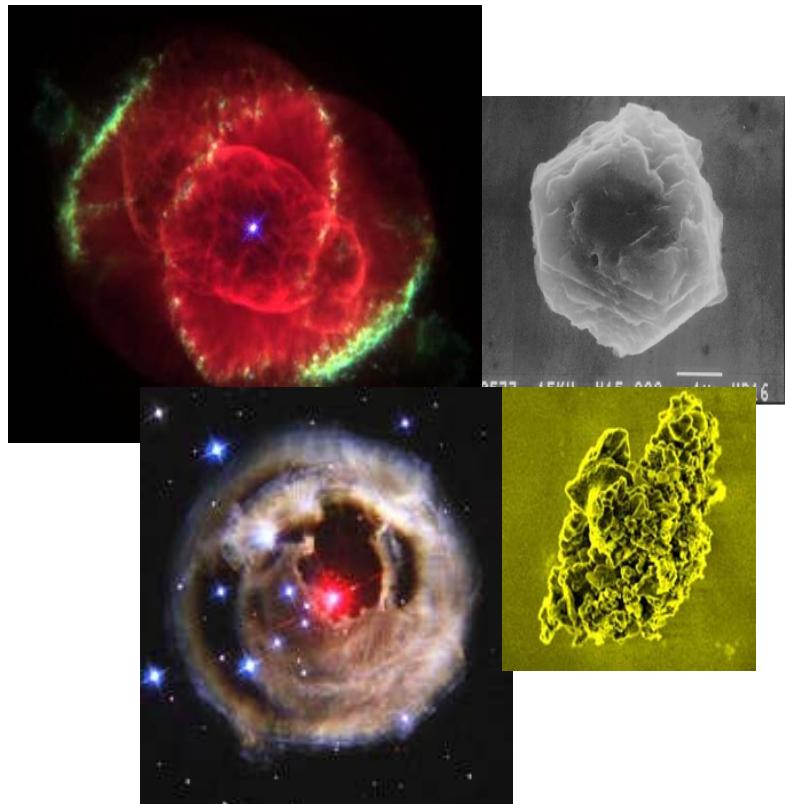


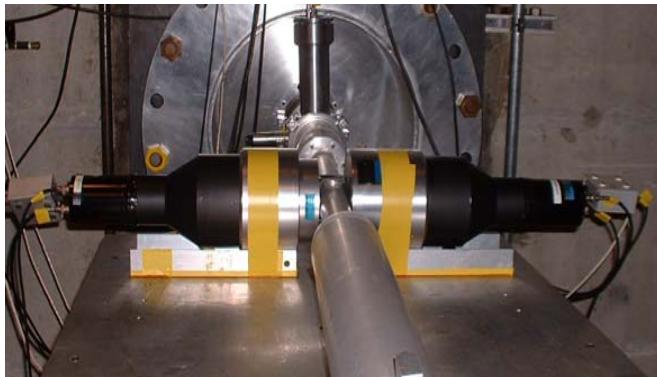
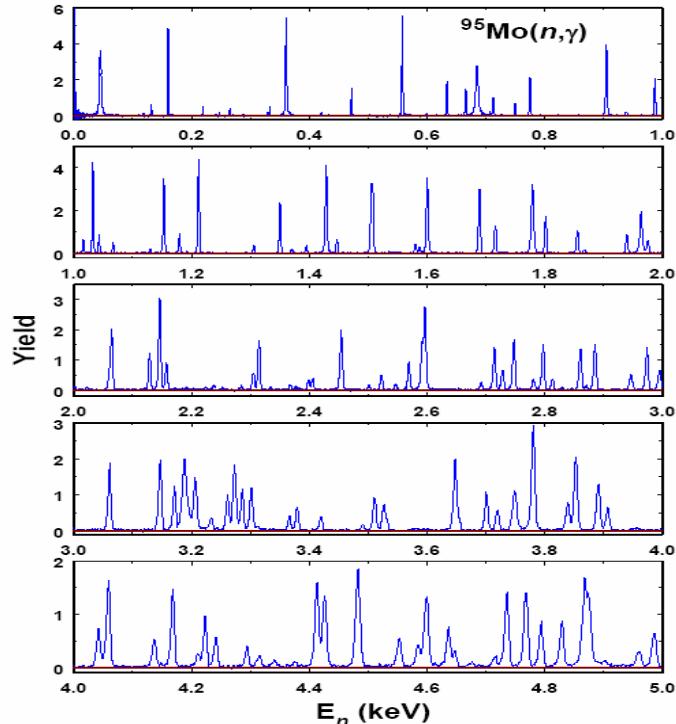
New ORELA ^{58}Ni Capture TOF data; using a metallic sample; 8 hours running time



New $^{95}\text{Mo}(n,\gamma)$ experiments

- AGB stellar models over predict the abundance of ^{95}Mo compared to observation in SiC grain which origin from an AGB star where the s-process takes place.
- M. Lugaro et. al. 2003: calculations show a 30% enhancement in the (n,γ) cross section for ^{95}Mo would solve the problem.





Old set up

New C_6D_6 Apparatus on FP6 in 40-m Station

- New system for (n,γ) experiments built in collaboration with JNC.
- Improved set up compared to last year. Less structural material. Two new C_6D_6 -detectors.
- $L = 38.5$ m, 2.5-cm diameter beam at sample.
- ^6Li -glass flux monitor.
- First test measurement: $^{95}\text{Mo}(n,\gamma)$. Finished!
- Completed transmission measurements.

Outlook

- Perform or finish new total and capture cross section measurements for the NCSP, i.e. nat. Cr, ^{53}Cr , ^{58}Ni , ^{60}Ni , ^{63}Cu , ^{65}Cu
- Continue nuclear astrophysics experiments
Finish $^{64}\text{Zn}(n,\alpha)$
Future experiments include (n,γ) and σ_t for $^{86,87}\text{Sr}$ and $^{149}\text{Sm}(n,\alpha)$

